

τ

$$J = \frac{1}{2}$$

τ discovery paper was PERL 75. $e^+ e^- \rightarrow \tau^+ \tau^-$ cross-section threshold behavior and magnitude are consistent with pointlike spin-1/2 Dirac particle. BRANDELIK 78 ruled out pointlike spin-0 or spin-1 particle. FELDMAN 78 ruled out $J = 3/2$. KIRKBY 79 also ruled out $J=\text{integer}$, $J = 3/2$.

τ MASS

| VALUE (MeV) | EVTS | DOCUMENT ID | TECN | COMMENT |
|---|-------|-------------|-----------|---|
| 1776.86±0.12 OUR AVERAGE | | | | |
| 1776.91±0.12 ^{+0.10} _{-0.13} | 1171 | 1 ABLIKIM | 14D BES3 | $23.3 \text{ pb}^{-1}, E_{\text{cm}}^{\text{ee}} = 3.54\text{--}3.60 \text{ GeV}$ |
| 1776.68±0.12±0.41 | 682k | 2 AUBERT | 09AK BABR | $423 \text{ fb}^{-1}, E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$ |
| 1776.81 ^{+0.25} _{-0.23} ±0.15 | 81 | ANASHIN | 07 KEDR | $6.7 \text{ pb}^{-1}, E_{\text{cm}}^{\text{ee}} = 3.54\text{--}3.78 \text{ GeV}$ |
| 1776.61±0.13±0.35 | | 2 BELOUS | 07 BELL | $414 \text{ fb}^{-1} E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$ |
| 1775.1 ±1.6 ±1.0 | 13.3k | 3 ABBIENDI | 00A OPAL | 1990–1995 LEP runs |
| 1778.2 ±0.8 ±1.2 | | ANASTASSOV | 97 CLEO | $E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$ |
| 1776.96 ^{+0.18} _{-0.21} ^{+0.25} _{-0.17} | 65 | 4 BAI | 96 BES | $E_{\text{cm}}^{\text{ee}} = 3.54\text{--}3.57 \text{ GeV}$ |
| 1776.3 ±2.4 ±1.4 | 11k | 5 ALBRECHT | 92M ARG | $E_{\text{cm}}^{\text{ee}} = 9.4\text{--}10.6 \text{ GeV}$ |
| 1783 ⁺³ ₋₄ | 692 | 6 BACINO | 78B DLCO | $E_{\text{cm}}^{\text{ee}} = 3.1\text{--}7.4 \text{ GeV}$ |
| • • • We do not use the following data for averages, fits, limits, etc. • • • | | | | |
| 1777.8 ±0.7 ±1.7 | 35k | 7 BALEST | 93 CLEO | Repl. by ANASTASSOV 97 |
| 1776.9 ^{+0.4} _{-0.5} ±0.2 | 14 | 8 BAI | 92 BES | Repl. by BAI 96 |

¹ ABLIKIM 14D fit $\sigma(e^+ e^- \rightarrow \tau^+ \tau^-)$ at different energies near threshold.

² AUBERT 09AK and BELOUS 07 fit τ pseudomass spectrum in $\tau \rightarrow \pi\pi^+\pi^-\nu_\tau$ decays.
Result assumes $m_{\nu_\tau} = 0$.

³ ABBIENDI 00A fit τ pseudomass spectrum in $\tau \rightarrow \pi^\pm \leq 2\pi^0\nu_\tau$ and
 $\tau \rightarrow \pi^\pm\pi^+\pi^- \leq 1\pi^0\nu_\tau$ decays. Result assumes $m_{\nu_\tau} = 0$.

⁴ BAI 96 fit $\sigma(e^+ e^- \rightarrow \tau^+ \tau^-)$ at different energies near threshold.

⁵ ALBRECHT 92M fit τ pseudomass spectrum in $\tau^- \rightarrow 2\pi^-\pi^+\nu_\tau$ decays. Result assumes $m_{\nu_\tau} = 0$.

⁶ BACINO 78B value comes from $e^\pm X^\mp$ threshold. Published mass 1782 MeV increased by 1 MeV using the high precision $\psi(2S)$ mass measurement of ZHOLENTZ 80 to eliminate the absolute SPEAR energy calibration uncertainty.

⁷ BALEST 93 fit spectra of minimum kinematically allowed τ mass in events of the type $e^+ e^- \rightarrow \tau^+ \tau^- \rightarrow (\pi^+ n\pi^0\nu_\tau)(\pi^- m\pi^0\nu_\tau)$ $n \leq 2$, $m \leq 2$, $1 \leq n+m \leq 3$. If $m_{\nu_\tau} \neq 0$, result increases by $(m_{\nu_\tau}^2/1100 \text{ MeV})$.

⁸ BAI 92 fit $\sigma(e^+ e^- \rightarrow \tau^+ \tau^-)$ near threshold using $e\mu$ events.

$(m_{\tau^+} - m_{\tau^-})/m_{\text{average}}$
A test of *CPT* invariance.

| VALUE | CL% | DOCUMENT ID | TECN | COMMENT |
|---|-----|---------------------|-----------|--|
| $<2.8 \times 10^{-4}$ | 90 | BELOUS | 07 | BELL $414 \text{ fb}^{-1}, E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$ |
| • • • We do not use the following data for averages, fits, limits, etc. • • • | | | | |
| $<5.5 \times 10^{-4}$ | 90 | ¹ AUBERT | 09AK BABR | $423 \text{ fb}^{-1}, E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$ |
| $<3.0 \times 10^{-3}$ | 90 | ABBIENDI | 00A OPAL | 1990–1995 LEP runs |
| ¹ AUBERT 09AK quote both the listed upper limit and $(m_{\tau^+} - m_{\tau^-})/m_{\text{average}} = (-3.4 \pm 1.3 \pm 0.3) \times 10^{-4}$. | | | | |

 $\tau \text{ MEAN LIFE}$

| VALUE (10^{-15} s) | EVTS | DOCUMENT ID | TECN | COMMENT |
|--|-------|-------------|------|---|
| 290.3 ± 0.5 OUR AVERAGE | | | | |
| 290.17 ± 0.53 ± 0.33 | 1.1M | BELOUS | 14 | BELL $711 \text{ fb}^{-1} E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$ |
| 290.9 ± 1.4 ± 1.0 | | ABDALLAH | 04T | DLPH 1991–1995 LEP runs |
| 293.2 ± 2.0 ± 1.5 | | ACCIARRI | 00B | L3 1991–1995 LEP runs |
| 290.1 ± 1.5 ± 1.1 | | BARATE | 97R | ALEP 1989–1994 LEP runs |
| 289.2 ± 1.7 ± 1.2 | | ALEXANDER | 96E | OPAL 1990–1994 LEP runs |
| 289.0 ± 2.8 ± 4.0 | 57.4k | BALEST | 96 | CLEO $E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$ |
| • • • We do not use the following data for averages, fits, limits, etc. • • • | | | | |
| 291.2 ± 2.0 ± 1.2 | | BARATE | 97I | ALEP Repl. by BARATE 97R |
| 291.4 ± 3.0 | | ABREU | 96B | DLPH Repl. by ABDALLAH 04T |
| 290.1 ± 4.0 | 34k | ACCIARRI | 96K | L3 Repl. by ACCIARRI 00B |
| 297 ± 9 ± 5 | 1671 | ABE | 95Y | SLD 1992–1993 SLC runs |
| 304 ± 14 ± 7 | 4100 | BATTLE | 92 | CLEO $E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$ |
| 301 ± 29 | 3780 | KLEINWORT | 89 | JADE $E_{\text{cm}}^{\text{ee}} = 35\text{--}46 \text{ GeV}$ |
| 288 ± 16 ± 17 | 807 | AMIDEI | 88 | MRK2 $E_{\text{cm}}^{\text{ee}} = 29 \text{ GeV}$ |
| 306 ± 20 ± 14 | 695 | BRAUNSCH... | 88C | TASS $E_{\text{cm}}^{\text{ee}} = 36 \text{ GeV}$ |
| 299 ± 15 ± 10 | 1311 | ABACHI | 87C | HRS $E_{\text{cm}}^{\text{ee}} = 29 \text{ GeV}$ |
| 295 ± 14 ± 11 | 5696 | ALBRECHT | 87P | ARG $E_{\text{cm}}^{\text{ee}} = 9.3\text{--}10.6 \text{ GeV}$ |
| 309 ± 17 ± 7 | 3788 | BAND | 87B | MAC $E_{\text{cm}}^{\text{ee}} = 29 \text{ GeV}$ |
| 325 ± 14 ± 18 | 8470 | BEBEK | 87C | CLEO $E_{\text{cm}}^{\text{ee}} = 10.5 \text{ GeV}$ |
| 460 ± 190 | 102 | FELDMAN | 82 | MRK2 $E_{\text{cm}}^{\text{ee}} = 29 \text{ GeV}$ |

 $(\tau_{\tau^+} - \tau_{\tau^-}) / \tau_{\text{average}}$
Test of *CPT* invariance.

| VALUE | CL% | DOCUMENT ID | TECN | COMMENT |
|-----------------------|-----|---------------------|------|---|
| $<7.0 \times 10^{-3}$ | 90 | ¹ BELOUS | 14 | BELL $711 \text{ fb}^{-1} E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$ |

¹BELOUS 14 quote limit on the absolute value of the relative lifetime difference.

τ MAGNETIC MOMENT ANOMALY

The q^2 dependence is expected to be small providing no thresholds are nearby.

$$\mu_\tau/(e\hbar/2m_\tau) - 1 = (g_\tau - 2)/2$$

For a theoretical calculation [$(g_\tau - 2)/2 = 117\,721(5) \times 10^{-8}$], see EIDELMAN 07.

| VALUE | CL% | DOCUMENT ID | TECN | COMMENT |
|---|-----|-----------------------------|------|--|
| > -0.052 and < 0.013 (CL = 95%) OUR LIMIT | | | | |
| > -0.052 and < 0.013 | 95 | ¹ ABDALLAH 04K | DLPH | $e^+ e^- \rightarrow e^+ e^- \tau^+ \tau^-$ at LEP2 |
| • • • We do not use the following data for averages, fits, limits, etc. • • • | | | | |
| < 0.107 | 95 | ² ACHARD 04G | L3 | $e^+ e^- \rightarrow e^+ e^- \tau^+ \tau^-$ at LEP2 |
| > -0.007 and < 0.005 | 95 | ³ GONZALEZ-S..00 | RVUE | $e^+ e^- \rightarrow \tau^+ \tau^-$ and $W \rightarrow \tau \nu_\tau$ |
| > -0.052 and < 0.058 | 95 | ⁴ ACCIARRI 98E | L3 | 1991–1995 LEP runs |
| > -0.068 and < 0.065 | 95 | ⁵ ACKERSTAFF 98N | OPAL | 1990–1995 LEP runs |
| > -0.004 and < 0.006 | 95 | ⁶ ESCRIBANO 97 | RVUE | $Z \rightarrow \tau^+ \tau^-$ at LEP |
| < 0.01 | 95 | ⁷ ESCRIBANO 93 | RVUE | $Z \rightarrow \tau^+ \tau^-$ at LEP |
| < 0.12 | 90 | GRIFOLS 91 | RVUE | $Z \rightarrow \tau \tau \gamma$ at LEP |
| < 0.023 | 95 | ⁸ SILVERMAN 83 | RVUE | $e^+ e^- \rightarrow \tau^+ \tau^-$ at PETRA |

¹ ABDALLAH 04K limit is derived from $e^+ e^- \rightarrow e^+ e^- \tau^+ \tau^-$ total cross-section measurements at \sqrt{s} between 183 and 208 GeV. In addition to the limits, the authors also quote a value of -0.018 ± 0.017 .

² ACHARD 04G limit is derived from $e^+ e^- \rightarrow e^+ e^- \tau^+ \tau^-$ total cross-section measurements at \sqrt{s} between 189 and 206 GeV, and is on the absolute value of the magnetic moment anomaly.

³ GONZALEZ-SPRINBERG 00 use data on tau lepton production at LEP1, SLC, and LEP2, and data from colliders and LEP2 to determine limits. Assume imaginary component is zero.

⁴ ACCIARRI 98E use $Z \rightarrow \tau^+ \tau^- \gamma$ events. In addition to the limits, the authors also quote a value of $0.004 \pm 0.027 \pm 0.023$.

⁵ ACKERSTAFF 98N use $Z \rightarrow \tau^+ \tau^- \gamma$ events. The limit applies to an average of the form factor for off-shell τ 's having p^2 ranging from m_τ^2 to $(M_Z - m_\tau)^2$.

⁶ ESCRIBANO 97 use preliminary experimental results.

⁷ ESCRIBANO 93 limit derived from $\Gamma(Z \rightarrow \tau^+ \tau^-)$, and is on the absolute value of the magnetic moment anomaly.

⁸ SILVERMAN 83 limit is derived from $e^+ e^- \rightarrow \tau^+ \tau^-$ total cross-section measurements for q^2 up to $(37 \text{ GeV})^2$.

τ ELECTRIC DIPOLE MOMENT (d_τ)

A nonzero value is forbidden by both T invariance and P invariance.

The q^2 dependence is expected to be small providing no thresholds are nearby.

$$\text{Re}(d_\tau)$$

| VALUE (10^{-16} ecm) | CL% | DOCUMENT ID | TECN | COMMENT |
|----------------------------------|-----|-----------------------|------|---|
| - 0.22 to 0.45 | 95 | ¹ INAMI 03 | BELL | $E_{\text{cm}}^{ee} = 10.6 \text{ GeV}$ |

• • • We do not use the following data for averages, fits, limits, etc. • • •

| | | | | | |
|------------------|----|--------------------------|-----|------|---|
| < 2.3 | 90 | ² GROZIN | 09A | RVUE | From e EDM limit |
| < 3.7 | 95 | ³ ABDALLAH | 04K | DLPH | $e^+e^- \rightarrow e^+e^-\tau^+\tau^-$ at LEP2 |
| < 11.4 | 95 | ⁴ ACHARD | 04G | L3 | $e^+e^- \rightarrow e^+e^-\tau^+\tau^-$ at LEP2 |
| < 4.6 | 95 | ⁵ ALBRECHT | 00 | ARG | $E_{\text{cm}}^{\text{ee}} = 10.4 \text{ GeV}$ |
| > -3.1 and < 3.1 | 95 | ACCIARRI | 98E | L3 | 1991–1995 LEP runs |
| > -3.8 and < 3.6 | 95 | ⁶ ACKERSTAFF | 98N | OPAL | 1990–1995 LEP runs |
| < 0.11 | 95 | ^{7,8} ESCRIBANO | 97 | RVUE | $Z \rightarrow \tau^+\tau^-$ at LEP |
| < 0.5 | 95 | ⁹ ESCRIBANO | 93 | RVUE | $Z \rightarrow \tau^+\tau^-$ at LEP |
| < 7 | 90 | GRIFOLS | 91 | RVUE | $Z \rightarrow \tau\tau\gamma$ at LEP |
| < 1.6 | 90 | DELAGUILA | 90 | RVUE | $e^+e^- \rightarrow \tau^+\tau^-$ $E_{\text{cm}}^{\text{ee}} = 35 \text{ GeV}$ |

¹ INAMI 03 use $e^+e^- \rightarrow \tau^+\tau^-$ events.

² GROZIN 09A calculate the contribution to the electron electric dipole moment from the τ electric dipole moment appearing in loops, which is $\Delta d_e = 6.9 \times 10^{-12} d_\tau$. Dividing the REGAN 02 upper limit $|d_e| \leq 1.6 \times 10^{-27} \text{ e cm}$ at CL=90% by 6.9×10^{-12} gives this limit.

³ ABDALLAH 04K limit is derived from $e^+e^- \rightarrow e^+e^-\tau^+\tau^-$ total cross-section measurements at \sqrt{s} between 183 and 208 GeV and is on the absolute value of d_τ .

⁴ ACHARD 04G limit is derived from $e^+e^- \rightarrow e^+e^-\tau^+\tau^-$ total cross-section measurements at \sqrt{s} between 189 and 206 GeV, and is on the absolute value of d_τ .

⁵ ALBRECHT 00 use $e^+e^- \rightarrow \tau^+\tau^-$ events. Limit is on the absolute value of $\text{Re}(d_\tau)$.

⁶ ACKERSTAFF 98N use $Z \rightarrow \tau^+\tau^-\gamma$ events. The limit applies to an average of the form factor for off-shell τ 's having p^2 ranging from m_τ^2 to $(M_Z - m_\tau)^2$.

⁷ ESCRIBANO 97 derive the relationship $|d_\tau| = \cot \theta_W |d_\tau^W|$ using effective Lagrangian methods, and use a conference result $|d_\tau^W| < 5.8 \times 10^{-18} \text{ e cm}$ at 95% CL (L. Silvestris, ICHEP96) to obtain this result.

⁸ ESCRIBANO 97 use preliminary experimental results.

⁹ ESCRIBANO 93 limit derived from $\Gamma(Z \rightarrow \tau^+\tau^-)$, and is on the absolute value of the electric dipole moment.

Im(d_τ)

| VALUE (10^{-16} ecm) | CL% | DOCUMENT ID | TECN | COMMENT |
|----------------------------------|-----|-----------------------|------|--|
| -0.25 to 0.008 | 95 | ¹ INAMI 03 | BELL | $E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$ |

• • • We do not use the following data for averages, fits, limits, etc. • • •

| | | | | |
|-------|----|--------------------------|-----|--|
| < 1.8 | 95 | ² ALBRECHT 00 | ARG | $E_{\text{cm}}^{\text{ee}} = 10.4 \text{ GeV}$ |
|-------|----|--------------------------|-----|--|

¹ INAMI 03 use $e^+e^- \rightarrow \tau^+\tau^-$ events.

² ALBRECHT 00 use $e^+e^- \rightarrow \tau^+\tau^-$ events. Limit is on the absolute value of $\text{Im}(d_\tau)$.

τ WEAK DIPOLE MOMENT (d_τ^W)

A nonzero value is forbidden by CP invariance.

The q^2 dependence is expected to be small providing no thresholds are nearby.

Re(d_τ^W)

| VALUE (10^{-17} ecm) | CL% | DOCUMENT ID | TECN | COMMENT |
|----------------------------------|-----|--------------------------|------|--------------------|
| <0.50 | 95 | ¹ HEISTER 03F | ALEP | 1990–1995 LEP runs |

• • • We do not use the following data for averages, fits, limits, etc. • • •

| | | | | | |
|-------|----|-----------------------|-----|------|--------------------------------------|
| <3.0 | 90 | ¹ ACCIARRI | 98C | L3 | 1991–1995 LEP runs |
| <0.56 | 95 | ACKERSTAFF | 97L | OPAL | 1991–1995 LEP runs |
| <0.78 | 95 | ² AKERS | 95F | OPAL | Repl. by ACKERSTAFF 97L |
| <1.5 | 95 | ² BUSKULIC | 95C | ALEP | Repl. by HEISTER 03F |
| <7.0 | 95 | ² ACTON | 92F | OPAL | $Z \rightarrow \tau^+ \tau^-$ at LEP |
| <3.7 | 95 | ² BUSKULIC | 92J | ALEP | Repl. by BUSKULIC 95C |

¹ Limit is on the absolute value of the real part of the weak dipole moment.

² Limit is on the absolute value of the real part of the weak dipole moment, and applies for $q^2 = m_Z^2$.

$\text{Im}(\alpha_\tau^w)$

| VALUE (10^{-17} ecm) | CL% | DOCUMENT ID | TECN | COMMENT |
|-------------------------|-----|----------------------|------|-------------------------|
| <1.1 | 95 | ¹ HEISTER | 03F | ALEP 1990–1995 LEP runs |

• • • We do not use the following data for averages, fits, limits, etc. • • •

| | | | | | |
|------|----|--------------------|-----|------|-------------------------|
| <1.5 | 95 | ACKERSTAFF | 97L | OPAL | 1991–1995 LEP runs |
| <4.5 | 95 | ² AKERS | 95F | OPAL | Repl. by ACKERSTAFF 97L |

¹ HEISTER 03F limit is on the absolute value of the imaginary part of the weak dipole moment.

² Limit is on the absolute value of the imaginary part of the weak dipole moment, and applies for $q^2 = m_Z^2$.

τ WEAK ANOMALOUS MAGNETIC DIPOLE MOMENT (α_τ^w)

Electroweak radiative corrections are expected to contribute at the 10^{-6} level. See BERNABEU 95.

The q^2 dependence is expected to be small providing no thresholds are nearby.

$\text{Re}(\alpha_\tau^w)$

| VALUE | CL% | DOCUMENT ID | TECN | COMMENT |
|----------------------------------|-----|----------------------|------|-------------------------|
| <1.1 × 10⁻³ | 95 | ¹ HEISTER | 03F | ALEP 1990–1995 LEP runs |

• • • We do not use the following data for averages, fits, limits, etc. • • •

| | | | | |
|-------------------------|----|------------------------------|------|--|
| > -0.0024 and < 0.0025 | 95 | ² GONZALEZ-S...00 | RVUE | $e^+ e^- \rightarrow \tau^+ \tau^-$ and $W \rightarrow \tau \nu_\tau$ |
| <4.5 × 10 ⁻³ | 90 | ¹ ACCIARRI | 98C | L3 1991–1995 LEP runs |

¹ Limit is on the absolute value of the real part of the weak anomalous magnetic dipole moment.

² GONZALEZ-SPRINBERG 00 use data on tau lepton production at LEP1, SLC, and LEP2, and data from colliders and LEP2 to determine limits. Assume imaginary component is zero.

$\text{Im}(\alpha_\tau^w)$

| VALUE | CL% | DOCUMENT ID | TECN | COMMENT |
|----------------------------------|-----|----------------------|------|-------------------------|
| <2.7 × 10⁻³ | 95 | ¹ HEISTER | 03F | ALEP 1990–1995 LEP runs |

• • • We do not use the following data for averages, fits, limits, etc. • • •

| | | | | |
|-------------------------|----|-----------------------|-----|-----------------------|
| <9.9 × 10 ⁻³ | 90 | ¹ ACCIARRI | 98C | L3 1991–1995 LEP runs |
|-------------------------|----|-----------------------|-----|-----------------------|

¹ Limit is on the absolute value of the imaginary part of the weak anomalous magnetic dipole moment.

τ^- DECAY MODES

τ^+ modes are charge conjugates of the modes below. “ h^\pm ” stands for π^\pm or K^\pm . “ ℓ ” stands for e or μ . “Neutrals” stands for γ 's and/or π^0 's.

| Mode | Fraction (Γ_i/Γ) | Scale factor/ Confidence level |
|--|--|-----------------------------------|
| Modes with one charged particle | | |
| Γ_1 particle $^- \geq 0$ neutrals $\geq 0 K^0 \nu_\tau$ ("1-prong") | (85.24 \pm 0.06) % | |
| Γ_2 particle $^- \geq 0$ neutrals $\geq 0 K_L^0 \nu_\tau$ | (84.58 \pm 0.06) % | |
| Γ_3 $\mu^- \bar{\nu}_\mu \nu_\tau$ | [a] (17.39 \pm 0.04) % | |
| Γ_4 $\mu^- \bar{\nu}_\mu \nu_\tau \gamma$ | [b] (3.68 \pm 0.10) $\times 10^{-3}$ | |
| Γ_5 $e^- \bar{\nu}_e \nu_\tau$ | [a] (17.82 \pm 0.04) % | |
| Γ_6 $e^- \bar{\nu}_e \nu_\tau \gamma$ | [b] (1.84 \pm 0.05) % | |
| Γ_7 $h^- \geq 0 K_L^0 \nu_\tau$ | (12.03 \pm 0.05) % | |
| Γ_8 $h^- \nu_\tau$ | (11.51 \pm 0.05) % | |
| Γ_9 $\pi^- \nu_\tau$ | [a] (10.82 \pm 0.05) % | |
| Γ_{10} $K^- \nu_\tau$ | [a] (6.96 \pm 0.10) $\times 10^{-3}$ | |
| Γ_{11} $h^- \geq 1$ neutrals ν_τ | (37.00 \pm 0.09) % | |
| Γ_{12} $h^- \geq 1 \pi^0 \nu_\tau$ (ex. K^0) | (36.51 \pm 0.09) % | |
| Γ_{13} $h^- \pi^0 \nu_\tau$ | (25.93 \pm 0.09) % | |
| Γ_{14} $\pi^- \pi^0 \nu_\tau$ | [a] (25.49 \pm 0.09) % | |
| Γ_{15} $\pi^- \pi^0$ non- $\rho(770) \nu_\tau$ | (3.0 \pm 3.2) $\times 10^{-3}$ | |
| Γ_{16} $K^- \pi^0 \nu_\tau$ | [a] (4.33 \pm 0.15) $\times 10^{-3}$ | |
| Γ_{17} $h^- \geq 2 \pi^0 \nu_\tau$ | (10.81 \pm 0.09) % | |
| Γ_{18} $h^- 2 \pi^0 \nu_\tau$ | (9.48 \pm 0.10) % | |
| Γ_{19} $h^- 2 \pi^0 \nu_\tau$ (ex. K^0) | (9.32 \pm 0.10) % | |
| Γ_{20} $\pi^- 2 \pi^0 \nu_\tau$ (ex. K^0) | [a] (9.26 \pm 0.10) % | |
| Γ_{21} $\pi^- 2 \pi^0 \nu_\tau$ (ex. K^0), scalar | < 9 $\times 10^{-3}$ CL=95% | |
| Γ_{22} $\pi^- 2 \pi^0 \nu_\tau$ (ex. K^0), vector | < 7 $\times 10^{-3}$ CL=95% | |
| Γ_{23} $K^- 2 \pi^0 \nu_\tau$ (ex. K^0) | [a] (6.5 \pm 2.2) $\times 10^{-4}$ | |
| Γ_{24} $h^- \geq 3 \pi^0 \nu_\tau$ | (1.34 \pm 0.07) % | |
| Γ_{25} $h^- \geq 3 \pi^0 \nu_\tau$ (ex. K^0) | (1.25 \pm 0.07) % | |
| Γ_{26} $h^- 3 \pi^0 \nu_\tau$ | (1.18 \pm 0.07) % | |
| Γ_{27} $\pi^- 3 \pi^0 \nu_\tau$ (ex. K^0) | [a] (1.04 \pm 0.07) % | |
| Γ_{28} $K^- 3 \pi^0 \nu_\tau$ (ex. K^0 , η) | [a] (4.8 \pm 2.1) $\times 10^{-4}$ | |
| Γ_{29} $h^- 4 \pi^0 \nu_\tau$ (ex. K^0) | (1.6 \pm 0.4) $\times 10^{-3}$ | |
| Γ_{30} $h^- 4 \pi^0 \nu_\tau$ (ex. K^0, η) | [a] (1.1 \pm 0.4) $\times 10^{-3}$ | |
| Γ_{31} $a_1(1260) \nu_\tau \rightarrow \pi^- \gamma \nu_\tau$ | (3.8 \pm 1.5) $\times 10^{-4}$ | |
| Γ_{32} $K^- \geq 0 \pi^0 \geq 0 K^0 \geq 0 \gamma \nu_\tau$ | (1.552 \pm 0.029) % | |
| Γ_{33} $K^- \geq 1 (\pi^0 \text{ or } K^0 \text{ or } \gamma) \nu_\tau$ | (8.59 \pm 0.28) $\times 10^{-3}$ | |

Modes with K^0 's

| | | |
|---------------|--|--|
| Γ_{34} | K_S^0 (particles) $^- \nu_\tau$ | (9.44 \pm 0.28) $\times 10^{-3}$ |
| Γ_{35} | $h^- \bar{K}^0 \nu_\tau$ | (9.87 \pm 0.14) $\times 10^{-3}$ |
| Γ_{36} | $\pi^- \bar{K}^0 \nu_\tau$ | [a] (8.40 \pm 0.14) $\times 10^{-3}$ |
| Γ_{37} | $\pi^- \bar{K}^0 (\text{non-}K^*(892)^-) \nu_\tau$ | (5.4 \pm 2.1) $\times 10^{-4}$ |
| Γ_{38} | $K^- K^0 \nu_\tau$ | [a] (1.48 \pm 0.05) $\times 10^{-3}$ |
| Γ_{39} | $K^- K^0 \geq 0 \pi^0 \nu_\tau$ | (2.98 \pm 0.08) $\times 10^{-3}$ |
| Γ_{40} | $h^- \bar{K}^0 \pi^0 \nu_\tau$ | (5.32 \pm 0.13) $\times 10^{-3}$ |
| Γ_{41} | $\pi^- \bar{K}^0 \pi^0 \nu_\tau$ | [a] (3.82 \pm 0.13) $\times 10^{-3}$ |
| Γ_{42} | $\bar{K}^0 \rho^- \nu_\tau$ | (2.2 \pm 0.5) $\times 10^{-3}$ |
| Γ_{43} | $K^- K^0 \pi^0 \nu_\tau$ | [a] (1.50 \pm 0.07) $\times 10^{-3}$ |
| Γ_{44} | $\pi^- \bar{K}^0 \geq 1 \pi^0 \nu_\tau$ | (4.08 \pm 0.25) $\times 10^{-3}$ |
| Γ_{45} | $\pi^- \bar{K}^0 \pi^0 \pi^0 \nu_\tau$ (ex. K^0) | [a] (2.6 \pm 2.3) $\times 10^{-4}$ |
| Γ_{46} | $K^- K^0 \pi^0 \pi^0 \nu_\tau$ | < 1.6 $\times 10^{-4}$ CL=95% |
| Γ_{47} | $\pi^- K^0 \bar{K}^0 \nu_\tau$ | (1.55 \pm 0.24) $\times 10^{-3}$ |
| Γ_{48} | $\pi^- K_S^0 K_S^0 \nu_\tau$ | [a] (2.33 \pm 0.07) $\times 10^{-4}$ |
| Γ_{49} | $\pi^- K_S^0 K_L^0 \nu_\tau$ | [a] (1.08 \pm 0.24) $\times 10^{-3}$ |
| Γ_{50} | $\pi^- K_L^0 K_L^0 \nu_\tau$ | (2.33 \pm 0.07) $\times 10^{-4}$ |
| Γ_{51} | $\pi^- K^0 \bar{K}^0 \pi^0 \nu_\tau$ | (3.6 \pm 1.2) $\times 10^{-4}$ |
| Γ_{52} | $\pi^- K_S^0 K_S^0 \pi^0 \nu_\tau$ | [a] (1.82 \pm 0.21) $\times 10^{-5}$ |
| Γ_{53} | $K^{*-} K^0 \pi^0 \nu_\tau \rightarrow \pi^- K_S^0 K_S^0 \pi^0 \nu_\tau$ | (1.08 \pm 0.21) $\times 10^{-5}$ |
| Γ_{54} | $f_1(1285) \pi^- \nu_\tau \rightarrow \pi^- K_S^0 K_S^0 \pi^0 \nu_\tau$ | (6.8 \pm 1.5) $\times 10^{-6}$ |
| Γ_{55} | $f_1(1420) \pi^- \nu_\tau \rightarrow \pi^- K_S^0 K_S^0 \pi^0 \nu_\tau$ | (2.4 \pm 0.8) $\times 10^{-6}$ |
| Γ_{56} | $\pi^- K_S^0 K_S^0 \pi^0 \nu_\tau$ | [a] (3.2 \pm 1.2) $\times 10^{-4}$ |
| Γ_{57} | $\pi^- K_L^0 K_L^0 \pi^0 \nu_\tau$ | (1.82 \pm 0.21) $\times 10^{-5}$ |
| Γ_{58} | $K^- K_S^0 K_S^0 \nu_\tau$ | < 6.3 $\times 10^{-7}$ CL=90% |
| Γ_{59} | $K^- K_S^0 K_S^0 \pi^0 \nu_\tau$ | < 4.0 $\times 10^{-7}$ CL=90% |
| Γ_{60} | $K^0 h^+ h^- h^- \geq 0$ neutrals ν_τ | < 1.7 $\times 10^{-3}$ CL=95% |
| Γ_{61} | $K^0 h^+ h^- h^- \nu_\tau$ | [a] (2.5 \pm 2.0) $\times 10^{-4}$ |

Modes with three charged particles

| | | |
|---------------|---|------------------------|
| Γ_{62} | $h^- h^- h^+ \geq 0$ neutrals $\geq 0 K_L^0 \nu_\tau$ | (15.21 \pm 0.06) % |
| Γ_{63} | $h^- h^- h^+ \geq 0$ neutrals ν_τ (ex. $K_S^0 \rightarrow \pi^+ \pi^-$) ("3-prong") | (14.55 \pm 0.06) % |
| Γ_{64} | $h^- h^- h^+ \nu_\tau$ | (9.80 \pm 0.05) % |
| Γ_{65} | $h^- h^- h^+ \nu_\tau$ (ex. K^0) | (9.46 \pm 0.05) % |
| Γ_{66} | $h^- h^- h^+ \nu_\tau$ (ex. K^0, ω) | (9.43 \pm 0.05) % |
| Γ_{67} | $\pi^- \pi^+ \pi^- \nu_\tau$ | (9.31 \pm 0.05) % |
| Γ_{68} | $\pi^- \pi^+ \pi^- \nu_\tau$ (ex. K^0) | (9.02 \pm 0.05) % |

| | | | | |
|----------------|--|-------|--|--------|
| Γ_{69} | $\pi^- \pi^+ \pi^- \nu_\tau$ (ex. K^0), non-axial vector | < 2.4 | % | CL=95% |
| Γ_{70} | $\pi^- \pi^+ \pi^- \nu_\tau$ (ex. K^0, ω) | [a] | (8.99 \pm 0.05) % | |
| Γ_{71} | $h^- h^- h^+ \geq 1$ neutrals ν_τ | | (5.29 \pm 0.05) % | |
| Γ_{72} | $h^- h^- h^+ \geq 1 \pi^0 \nu_\tau$ (ex. K^0) | | (5.09 \pm 0.05) % | |
| Γ_{73} | $h^- h^- h^+ \pi^0 \nu_\tau$ | | (4.76 \pm 0.05) % | |
| Γ_{74} | $h^- h^- h^+ \pi^0 \nu_\tau$ (ex. K^0) | | (4.57 \pm 0.05) % | |
| Γ_{75} | $h^- h^- h^+ \pi^0 \nu_\tau$ (ex. K^0, ω) | | (2.79 \pm 0.07) % | |
| Γ_{76} | $\pi^- \pi^+ \pi^- \pi^0 \nu_\tau$ | | (4.62 \pm 0.05) % | |
| Γ_{77} | $\pi^- \pi^+ \pi^- \pi^0 \nu_\tau$ (ex. K^0) | | (4.49 \pm 0.05) % | |
| Γ_{78} | $\pi^- \pi^+ \pi^- \pi^0 \nu_\tau$ (ex. K^0, ω) | [a] | (2.74 \pm 0.07) % | |
| Γ_{79} | $h^- \rho \pi^0 \nu_\tau$ | | | |
| Γ_{80} | $h^- \rho^+ h^- \nu_\tau$ | | | |
| Γ_{81} | $h^- \rho^- h^+ \nu_\tau$ | | | |
| Γ_{82} | $h^- h^- h^+ \geq 2 \pi^0 \nu_\tau$ (ex. K^0) | | (5.17 \pm 0.31) $\times 10^{-3}$ | |
| Γ_{83} | $h^- h^- h^+ 2 \pi^0 \nu_\tau$ | | (5.05 \pm 0.31) $\times 10^{-3}$ | |
| Γ_{84} | $h^- h^- h^+ 2 \pi^0 \nu_\tau$ (ex. K^0) | | (4.95 \pm 0.31) $\times 10^{-3}$ | |
| Γ_{85} | $h^- h^- h^+ 2 \pi^0 \nu_\tau$ (ex. K^0, ω, η) | [a] | (10 \pm 4) $\times 10^{-4}$ | |
| Γ_{86} | $h^- h^- h^+ 3 \pi^0 \nu_\tau$ | | (2.12 \pm 0.30) $\times 10^{-4}$ | |
| Γ_{87} | $2 \pi^- \pi^+ 3 \pi^0 \nu_\tau$ (ex. K^0) | | (1.94 \pm 0.30) $\times 10^{-4}$ | |
| Γ_{88} | $2 \pi^- \pi^+ 3 \pi^0 \nu_\tau$ (ex. $K^0, \eta, f_1(1285)$) | | (1.7 \pm 0.4) $\times 10^{-4}$ | |
| Γ_{89} | $2 \pi^- \pi^+ 3 \pi^0 \nu_\tau$ (ex. $K^0, \eta, \omega, f_1(1285)$) | [a] | (1.4 \pm 2.7) $\times 10^{-5}$ | |
| Γ_{90} | $K^- h^+ h^- \geq 0$ neutrals ν_τ | | (6.29 \pm 0.14) $\times 10^{-3}$ | |
| Γ_{91} | $K^- h^+ \pi^- \nu_\tau$ (ex. K^0) | | (4.37 \pm 0.07) $\times 10^{-3}$ | |
| Γ_{92} | $K^- h^+ \pi^- \pi^0 \nu_\tau$ (ex. K^0) | | (8.6 \pm 1.2) $\times 10^{-4}$ | |
| Γ_{93} | $K^- \pi^+ \pi^- \geq 0$ neutrals ν_τ | | (4.77 \pm 0.14) $\times 10^{-3}$ | |
| Γ_{94} | $K^- \pi^+ \pi^- \geq 0 \pi^0 \nu_\tau$ (ex. K^0) | | (3.73 \pm 0.13) $\times 10^{-3}$ | |
| Γ_{95} | $K^- \pi^+ \pi^- \nu_\tau$ | | (3.45 \pm 0.07) $\times 10^{-3}$ | |
| Γ_{96} | $K^- \pi^+ \pi^- \nu_\tau$ (ex. K^0) | | (2.93 \pm 0.07) $\times 10^{-3}$ | |
| Γ_{97} | $K^- \pi^+ \pi^- \nu_\tau$ (ex. K^0, ω) | [a] | (2.93 \pm 0.07) $\times 10^{-3}$ | |
| Γ_{98} | $K^- \rho^0 \nu_\tau \rightarrow K^- \pi^+ \pi^- \nu_\tau$ | | (1.4 \pm 0.5) $\times 10^{-3}$ | |
| Γ_{99} | $K^- \pi^+ \pi^- \pi^0 \nu_\tau$ | | (1.31 \pm 0.12) $\times 10^{-3}$ | |
| Γ_{100} | $K^- \pi^+ \pi^- \pi^0 \nu_\tau$ (ex. K^0) | | (7.9 \pm 1.2) $\times 10^{-4}$ | |
| Γ_{101} | $K^- \pi^+ \pi^- \pi^0 \nu_\tau$ (ex. K^0, η) | | (7.6 \pm 1.2) $\times 10^{-4}$ | |
| Γ_{102} | $K^- \pi^+ \pi^- \pi^0 \nu_\tau$ (ex. K^0, ω) | | (3.7 \pm 0.9) $\times 10^{-4}$ | |
| Γ_{103} | $K^- \pi^+ \pi^- \pi^0 \nu_\tau$ (ex. K^0, ω, η) | [a] | (3.9 \pm 1.4) $\times 10^{-4}$ | |
| Γ_{104} | $K^- \pi^+ K^- \geq 0$ neut. ν_τ | < 9 | $\times 10^{-4}$ | CL=95% |
| Γ_{105} | $K^- K^+ \pi^- \geq 0$ neut. ν_τ | | (1.496 \pm 0.033) $\times 10^{-3}$ | |
| Γ_{106} | $K^- K^+ \pi^- \nu_\tau$ | [a] | (1.435 \pm 0.027) $\times 10^{-3}$ | |
| Γ_{107} | $K^- K^+ \pi^- \pi^0 \nu_\tau$ | [a] | (6.1 \pm 1.8) $\times 10^{-5}$ | |

| | | | |
|----------------|--|--------------------------------|--------|
| Γ_{108} | $K^- K^+ K^- \nu_\tau$ | $(2.2 \pm 0.8) \times 10^{-5}$ | S=5.4 |
| Γ_{109} | $K^- K^+ K^- \nu_\tau$ (ex. ϕ) | $< 2.5 \times 10^{-6}$ | CL=90% |
| Γ_{110} | $K^- K^+ K^- \pi^0 \nu_\tau$ | $< 4.8 \times 10^{-6}$ | CL=90% |
| Γ_{111} | $\pi^- K^+ \pi^- \geq 0$ neutrals ν_τ | $< 2.5 \times 10^{-3}$ | CL=95% |
| Γ_{112} | $e^- e^- e^+ \bar{\nu}_e \nu_\tau$ | $(2.8 \pm 1.5) \times 10^{-5}$ | |
| Γ_{113} | $\mu^- e^- e^+ \bar{\nu}_\mu \nu_\tau$ | $< 3.6 \times 10^{-5}$ | CL=90% |

Modes with five charged particles

| | | | |
|----------------|---|--------------------------------------|--------|
| Γ_{114} | $3h^- 2h^+ \geq 0$ neutrals ν_τ (ex. $K_S^0 \rightarrow \pi^- \pi^+$) ("5-prong") | $(9.9 \pm 0.4) \times 10^{-4}$ | |
| Γ_{115} | $3h^- 2h^+ \nu_\tau$ (ex. K^0) | $(8.22 \pm 0.32) \times 10^{-4}$ | |
| Γ_{116} | $3\pi^- 2\pi^+ \nu_\tau$ (ex. K^0, ω) | $(8.21 \pm 0.31) \times 10^{-4}$ | |
| Γ_{117} | $3\pi^- 2\pi^+ \nu_\tau$ (ex. $K^0, \omega, f_1(1285)$) | [a] $(7.69 \pm 0.30) \times 10^{-4}$ | |
| Γ_{118} | $K^- 2\pi^- 2\pi^+ \nu_\tau$ (ex. K^0) | [a] $(6 \pm 12) \times 10^{-7}$ | |
| Γ_{119} | $K^+ 3\pi^- \pi^+ \nu_\tau$ | $< 5.0 \times 10^{-6}$ | CL=90% |
| Γ_{120} | $K^+ K^- 2\pi^- \pi^+ \nu_\tau$ | $< 4.5 \times 10^{-7}$ | CL=90% |
| Γ_{121} | $3h^- 2h^+ \pi^0 \nu_\tau$ (ex. K^0) | $(1.64 \pm 0.11) \times 10^{-4}$ | |
| Γ_{122} | $3\pi^- 2\pi^+ \pi^0 \nu_\tau$ (ex. K^0) | $(1.62 \pm 0.11) \times 10^{-4}$ | |
| Γ_{123} | $3\pi^- 2\pi^+ \pi^0 \nu_\tau$ (ex. $K^0, \eta, f_1(1285)$) | $(1.11 \pm 0.10) \times 10^{-4}$ | |
| Γ_{124} | $3\pi^- 2\pi^+ \pi^0 \nu_\tau$ (ex. $K^0, \eta, \omega, f_1(1285)$) | [a] $(3.8 \pm 0.9) \times 10^{-5}$ | |
| Γ_{125} | $K^- 2\pi^- 2\pi^+ \pi^0 \nu_\tau$ (ex. K^0) | [a] $(1.1 \pm 0.6) \times 10^{-6}$ | |
| Γ_{126} | $K^+ 3\pi^- \pi^+ \pi^0 \nu_\tau$ | $< 8 \times 10^{-7}$ | CL=90% |
| Γ_{127} | $3h^- 2h^+ 2\pi^0 \nu_\tau$ | $< 3.4 \times 10^{-6}$ | CL=90% |

Miscellaneous other allowed modes

| | | | |
|----------------|---|----------------------------------|--------|
| Γ_{128} | $(5\pi)^- \nu_\tau$ | $(7.8 \pm 0.5) \times 10^{-3}$ | |
| Γ_{129} | $4h^- 3h^+ \geq 0$ neutrals ν_τ ("7-prong") | $< 3.0 \times 10^{-7}$ | CL=90% |
| Γ_{130} | $4h^- 3h^+ \nu_\tau$ | $< 4.3 \times 10^{-7}$ | CL=90% |
| Γ_{131} | $4h^- 3h^+ \pi^0 \nu_\tau$ | $< 2.5 \times 10^{-7}$ | CL=90% |
| Γ_{132} | $X^- (S=-1) \nu_\tau$ | $(2.92 \pm 0.04) \%$ | |
| Γ_{133} | $K^*(892)^- \geq 0$ neutrals $\geq 0K_L^0 \nu_\tau$ | $(1.42 \pm 0.18) \%$ | S=1.4 |
| Γ_{134} | $K^*(892)^- \nu_\tau$ | $(1.20 \pm 0.07) \%$ | S=1.8 |
| Γ_{135} | $K^*(892)^- \nu_\tau \rightarrow \pi^- \bar{K}^0 \nu_\tau$ | $(7.83 \pm 0.26) \times 10^{-3}$ | |
| Γ_{136} | $K^*(892)^0 K^- \geq 0$ neutrals ν_τ | $(3.2 \pm 1.4) \times 10^{-3}$ | |
| Γ_{137} | $K^*(892)^0 K^- \nu_\tau$ | $(2.1 \pm 0.4) \times 10^{-3}$ | |
| Γ_{138} | $\bar{K}^*(892)^0 \pi^- \geq 0$ neutrals ν_τ | $(3.8 \pm 1.7) \times 10^{-3}$ | |
| Γ_{139} | $\bar{K}^*(892)^0 \pi^- \nu_\tau$ | $(2.2 \pm 0.5) \times 10^{-3}$ | |
| Γ_{140} | $(\bar{K}^*(892)\pi)^- \nu_\tau \rightarrow \pi^- \bar{K}^0 \pi^0 \nu_\tau$ | $(1.0 \pm 0.4) \times 10^{-3}$ | |

| | | | |
|----------------|--|--------------------------------------|--------|
| Γ_{141} | $K_1(1270)^-\nu_\tau$ | $(4.7 \pm 1.1) \times 10^{-3}$ | |
| Γ_{142} | $K_1(1400)^-\nu_\tau$ | $(1.7 \pm 2.6) \times 10^{-3}$ | S=1.7 |
| Γ_{143} | $K^*(1410)^-\nu_\tau$ | $(1.5 \pm 1.4) \times 10^{-3}$ | |
| Γ_{144} | $K_0^*(1430)^-\nu_\tau$ | $< 5 \times 10^{-4}$ | CL=95% |
| Γ_{145} | $K_2^*(1430)^-\nu_\tau$ | $< 3 \times 10^{-3}$ | CL=95% |
| Γ_{146} | $a_0(980)^-\geq 0 \text{ neutrals } \nu_\tau$ | | |
| Γ_{147} | $\eta\pi^-\nu_\tau$ | $< 9.9 \times 10^{-5}$ | CL=95% |
| Γ_{148} | $\eta\pi^-\pi^0\nu_\tau$ | [a] $(1.39 \pm 0.07) \times 10^{-3}$ | |
| Γ_{149} | $\eta\pi^-\pi^0\pi^0\nu_\tau$ | [a] $(1.9 \pm 0.4) \times 10^{-4}$ | |
| Γ_{150} | $\eta K^-\nu_\tau$ | [a] $(1.55 \pm 0.08) \times 10^{-4}$ | |
| Γ_{151} | $\eta K^*(892)^-\nu_\tau$ | $(1.38 \pm 0.15) \times 10^{-4}$ | |
| Γ_{152} | $\eta K^-\pi^0\nu_\tau$ | [a] $(4.8 \pm 1.2) \times 10^{-5}$ | |
| Γ_{153} | $\eta K^-\pi^0(\text{non-}K^*(892))\nu_\tau$ | $< 3.5 \times 10^{-5}$ | CL=90% |
| Γ_{154} | $\eta\bar{K}^0\pi^-\nu_\tau$ | [a] $(9.4 \pm 1.5) \times 10^{-5}$ | |
| Γ_{155} | $\eta\bar{K}^0\pi^-\pi^0\nu_\tau$ | $< 5.0 \times 10^{-5}$ | CL=90% |
| Γ_{156} | $\eta K^-K^0\nu_\tau$ | $< 9.0 \times 10^{-6}$ | CL=90% |
| Γ_{157} | $\eta\pi^+\pi^-\pi^-\geq 0 \text{ neutrals } \nu_\tau$ | $< 3 \times 10^{-3}$ | CL=90% |
| Γ_{158} | $\eta\pi^-\pi^+\pi^-\nu_\tau (\text{ex. } K^0)$ | [a] $(2.19 \pm 0.13) \times 10^{-4}$ | |
| Γ_{159} | $\eta\pi^-\pi^+\pi^-\nu_\tau (\text{ex. } K^0, f_1(1285))$ | $(9.9 \pm 1.6) \times 10^{-5}$ | |
| Γ_{160} | $\eta a_1(1260)^-\nu_\tau \rightarrow \eta\pi^-\rho^0\nu_\tau$ | $< 3.9 \times 10^{-4}$ | CL=90% |
| Γ_{161} | $\eta\eta\pi^-\nu_\tau$ | $< 7.4 \times 10^{-6}$ | CL=90% |
| Γ_{162} | $\eta\eta\pi^-\pi^0\nu_\tau$ | $< 2.0 \times 10^{-4}$ | CL=95% |
| Γ_{163} | $\eta\eta K^-\nu_\tau$ | $< 3.0 \times 10^{-6}$ | CL=90% |
| Γ_{164} | $\eta'(958)\pi^-\nu_\tau$ | $< 4.0 \times 10^{-6}$ | CL=90% |
| Γ_{165} | $\eta'(958)\pi^-\pi^0\nu_\tau$ | $< 1.2 \times 10^{-5}$ | CL=90% |
| Γ_{166} | $\eta'(958)K^-\nu_\tau$ | $< 2.4 \times 10^{-6}$ | CL=90% |
| Γ_{167} | $\phi\pi^-\nu_\tau$ | $(3.4 \pm 0.6) \times 10^{-5}$ | |
| Γ_{168} | $\phi K^-\nu_\tau$ | [a] $(4.4 \pm 1.6) \times 10^{-5}$ | |
| Γ_{169} | $f_1(1285)\pi^-\nu_\tau$ | $(3.9 \pm 0.5) \times 10^{-4}$ | S=1.9 |
| Γ_{170} | $f_1(1285)\pi^-\nu_\tau \rightarrow \eta\pi^-\pi^+\pi^-\nu_\tau$ | $(1.18 \pm 0.07) \times 10^{-4}$ | S=1.3 |
| Γ_{171} | $f_1(1285)\pi^-\nu_\tau \rightarrow 3\pi^-2\pi^+\nu_\tau$ | [a] $(5.2 \pm 0.4) \times 10^{-5}$ | |
| Γ_{172} | $\pi(1300)^-\nu_\tau \rightarrow (\rho\pi)^-\nu_\tau \rightarrow (3\pi)^-\nu_\tau$ | $< 1.0 \times 10^{-4}$ | CL=90% |
| Γ_{173} | $\pi(1300)^-\nu_\tau \rightarrow ((\pi\pi)_{S-\text{wave}}\pi)^-\nu_\tau \rightarrow (3\pi)^-\nu_\tau$ | $< 1.9 \times 10^{-4}$ | CL=90% |
| Γ_{174} | $h^-\omega\geq 0 \text{ neutrals } \nu_\tau$ | $(2.40 \pm 0.08) \%$ | |
| Γ_{175} | $h^-\omega\nu_\tau$ | $(1.99 \pm 0.06) \%$ | |
| Γ_{176} | $\pi^-\omega\nu_\tau$ | [a] $(1.95 \pm 0.06) \%$ | |
| Γ_{177} | $K^-\omega\nu_\tau$ | [a] $(4.1 \pm 0.9) \times 10^{-4}$ | |
| Γ_{178} | $h^-\omega\pi^0\nu_\tau$ | [a] $(4.1 \pm 0.4) \times 10^{-3}$ | |
| Γ_{179} | $h^-\omega 2\pi^0\nu_\tau$ | $(1.4 \pm 0.5) \times 10^{-4}$ | |

| | | | | | |
|----------------|---|-----|------------------------------------|------------------|--------|
| Γ_{180} | $\pi^- \omega 2\pi^0 \nu_\tau$ | [a] | $(7.1 \pm 1.6) \times 10^{-5}$ | | |
| Γ_{181} | $h^- 2\omega \nu_\tau$ | | < 5.4 | $\times 10^{-7}$ | CL=90% |
| Γ_{182} | $2h^- h^+ \omega \nu_\tau$ | | $(1.20 \pm 0.22) \times 10^{-4}$ | | |
| Γ_{183} | $2\pi^- \pi^+ \omega \nu_\tau$ (ex. K^0) | [a] | $(8.4 \pm 0.6) \times 10^{-5}$ | | |

**Lepton Family number (*LF*), Lepton number (*L*),
or Baryon number (*B*) violating modes**

L means lepton number violation (e.g. $\tau^- \rightarrow e^+ \pi^- \pi^-$). Following common usage, *LF* means lepton family violation *and not* lepton number violation (e.g. $\tau^- \rightarrow e^- \pi^+ \pi^-$). *B* means baryon number violation.

| | | | | | |
|----------------|--|-----------|---------|------------------|--------|
| Γ_{184} | $e^- \gamma$ | <i>LF</i> | < 3.3 | $\times 10^{-8}$ | CL=90% |
| Γ_{185} | $\mu^- \gamma$ | <i>LF</i> | < 4.4 | $\times 10^{-8}$ | CL=90% |
| Γ_{186} | $e^- \pi^0$ | <i>LF</i> | < 8.0 | $\times 10^{-8}$ | CL=90% |
| Γ_{187} | $\mu^- \pi^0$ | <i>LF</i> | < 1.1 | $\times 10^{-7}$ | CL=90% |
| Γ_{188} | $e^- K_S^0$ | <i>LF</i> | < 2.6 | $\times 10^{-8}$ | CL=90% |
| Γ_{189} | $\mu^- K_S^0$ | <i>LF</i> | < 2.3 | $\times 10^{-8}$ | CL=90% |
| Γ_{190} | $e^- \eta$ | <i>LF</i> | < 9.2 | $\times 10^{-8}$ | CL=90% |
| Γ_{191} | $\mu^- \eta$ | <i>LF</i> | < 6.5 | $\times 10^{-8}$ | CL=90% |
| Γ_{192} | $e^- \rho^0$ | <i>LF</i> | < 1.8 | $\times 10^{-8}$ | CL=90% |
| Γ_{193} | $\mu^- \rho^0$ | <i>LF</i> | < 1.2 | $\times 10^{-8}$ | CL=90% |
| Γ_{194} | $e^- \omega$ | <i>LF</i> | < 4.8 | $\times 10^{-8}$ | CL=90% |
| Γ_{195} | $\mu^- \omega$ | <i>LF</i> | < 4.7 | $\times 10^{-8}$ | CL=90% |
| Γ_{196} | $e^- K^*(892)^0$ | <i>LF</i> | < 3.2 | $\times 10^{-8}$ | CL=90% |
| Γ_{197} | $\mu^- K^*(892)^0$ | <i>LF</i> | < 5.9 | $\times 10^{-8}$ | CL=90% |
| Γ_{198} | $e^- \bar{K}^*(892)^0$ | <i>LF</i> | < 3.4 | $\times 10^{-8}$ | CL=90% |
| Γ_{199} | $\mu^- \bar{K}^*(892)^0$ | <i>LF</i> | < 7.0 | $\times 10^{-8}$ | CL=90% |
| Γ_{200} | $e^- \eta'(958)$ | <i>LF</i> | < 1.6 | $\times 10^{-7}$ | CL=90% |
| Γ_{201} | $\mu^- \eta'(958)$ | <i>LF</i> | < 1.3 | $\times 10^{-7}$ | CL=90% |
| Γ_{202} | $e^- f_0(980) \rightarrow e^- \pi^+ \pi^-$ | <i>LF</i> | < 3.2 | $\times 10^{-8}$ | CL=90% |
| Γ_{203} | $\mu^- f_0(980) \rightarrow \mu^- \pi^+ \pi^-$ | <i>LF</i> | < 3.4 | $\times 10^{-8}$ | CL=90% |
| Γ_{204} | $e^- \phi$ | <i>LF</i> | < 3.1 | $\times 10^{-8}$ | CL=90% |
| Γ_{205} | $\mu^- \phi$ | <i>LF</i> | < 8.4 | $\times 10^{-8}$ | CL=90% |
| Γ_{206} | $e^- e^+ e^-$ | <i>LF</i> | < 2.7 | $\times 10^{-8}$ | CL=90% |
| Γ_{207} | $e^- \mu^+ \mu^-$ | <i>LF</i> | < 2.7 | $\times 10^{-8}$ | CL=90% |
| Γ_{208} | $e^+ \mu^- \mu^-$ | <i>LF</i> | < 1.7 | $\times 10^{-8}$ | CL=90% |
| Γ_{209} | $\mu^- e^+ e^-$ | <i>LF</i> | < 1.8 | $\times 10^{-8}$ | CL=90% |
| Γ_{210} | $\mu^+ e^- e^-$ | <i>LF</i> | < 1.5 | $\times 10^{-8}$ | CL=90% |
| Γ_{211} | $\mu^- \mu^+ \mu^-$ | <i>LF</i> | < 2.1 | $\times 10^{-8}$ | CL=90% |
| Γ_{212} | $e^- \pi^+ \pi^-$ | <i>LF</i> | < 2.3 | $\times 10^{-8}$ | CL=90% |
| Γ_{213} | $e^+ \pi^- \pi^-$ | <i>L</i> | < 2.0 | $\times 10^{-8}$ | CL=90% |
| Γ_{214} | $\mu^- \pi^+ \pi^-$ | <i>LF</i> | < 2.1 | $\times 10^{-8}$ | CL=90% |
| Γ_{215} | $\mu^+ \pi^- \pi^-$ | <i>L</i> | < 3.9 | $\times 10^{-8}$ | CL=90% |
| Γ_{216} | $e^- \pi^+ K^-$ | <i>LF</i> | < 3.7 | $\times 10^{-8}$ | CL=90% |
| Γ_{217} | $e^- \pi^- K^+$ | <i>LF</i> | < 3.1 | $\times 10^{-8}$ | CL=90% |

| | | | | | |
|----------------|----------------------------|--------|---------|------------------|--------|
| Γ_{218} | $e^+ \pi^- K^-$ | L | < 3.2 | $\times 10^{-8}$ | CL=90% |
| Γ_{219} | $e^- K_S^0 K_S^0$ | LF | < 7.1 | $\times 10^{-8}$ | CL=90% |
| Γ_{220} | $e^- K^+ K^-$ | LF | < 3.4 | $\times 10^{-8}$ | CL=90% |
| Γ_{221} | $e^+ K^- K^-$ | L | < 3.3 | $\times 10^{-8}$ | CL=90% |
| Γ_{222} | $\mu^- \pi^+ K^-$ | LF | < 8.6 | $\times 10^{-8}$ | CL=90% |
| Γ_{223} | $\mu^- \pi^- K^+$ | LF | < 4.5 | $\times 10^{-8}$ | CL=90% |
| Γ_{224} | $\mu^+ \pi^- K^-$ | L | < 4.8 | $\times 10^{-8}$ | CL=90% |
| Γ_{225} | $\mu^- K_S^0 K_S^0$ | LF | < 8.0 | $\times 10^{-8}$ | CL=90% |
| Γ_{226} | $\mu^- K^+ K^-$ | LF | < 4.4 | $\times 10^{-8}$ | CL=90% |
| Γ_{227} | $\mu^+ K^- K^-$ | L | < 4.7 | $\times 10^{-8}$ | CL=90% |
| Γ_{228} | $e^- \pi^0 \pi^0$ | LF | < 6.5 | $\times 10^{-6}$ | CL=90% |
| Γ_{229} | $\mu^- \pi^0 \pi^0$ | LF | < 1.4 | $\times 10^{-5}$ | CL=90% |
| Γ_{230} | $e^- \eta \eta$ | LF | < 3.5 | $\times 10^{-5}$ | CL=90% |
| Γ_{231} | $\mu^- \eta \eta$ | LF | < 6.0 | $\times 10^{-5}$ | CL=90% |
| Γ_{232} | $e^- \pi^0 \eta$ | LF | < 2.4 | $\times 10^{-5}$ | CL=90% |
| Γ_{233} | $\mu^- \pi^0 \eta$ | LF | < 2.2 | $\times 10^{-5}$ | CL=90% |
| Γ_{234} | $p \mu^- \mu^-$ | L, B | < 4.4 | $\times 10^{-7}$ | CL=90% |
| Γ_{235} | $\bar{p} \mu^+ \mu^-$ | L, B | < 3.3 | $\times 10^{-7}$ | CL=90% |
| Γ_{236} | $\bar{p} \gamma$ | L, B | < 3.5 | $\times 10^{-6}$ | CL=90% |
| Γ_{237} | $\bar{p} \pi^0$ | L, B | < 1.5 | $\times 10^{-5}$ | CL=90% |
| Γ_{238} | $\bar{p} 2\pi^0$ | L, B | < 3.3 | $\times 10^{-5}$ | CL=90% |
| Γ_{239} | $\bar{p} \eta$ | L, B | < 8.9 | $\times 10^{-6}$ | CL=90% |
| Γ_{240} | $\bar{p} \pi^0 \eta$ | L, B | < 2.7 | $\times 10^{-5}$ | CL=90% |
| Γ_{241} | $\Lambda \pi^-$ | L, B | < 7.2 | $\times 10^{-8}$ | CL=90% |
| Γ_{242} | $\bar{\Lambda} \pi^-$ | L, B | < 1.4 | $\times 10^{-7}$ | CL=90% |
| Γ_{243} | $e^- \text{light boson}$ | LF | < 2.7 | $\times 10^{-3}$ | CL=95% |
| Γ_{244} | $\mu^- \text{light boson}$ | LF | < 5 | $\times 10^{-3}$ | CL=95% |

[a] Basis mode for the τ .

[b] See the Particle Listings below for the energy limits used in this measurement.

CONSTRAINED FIT INFORMATION

An overall fit to 85 branching ratios uses 169 measurements and one constraint to determine 46 parameters. The overall fit has a $\chi^2 = 134.9$ for 124 degrees of freedom.

The following *off-diagonal* array elements are the correlation coefficients $\langle \delta x_i \delta x_j \rangle / (\delta x_i \cdot \delta x_j)$, in percent, from the fit to the branching fractions, $x_i \equiv \Gamma_i / \Gamma_{\text{total}}$.

| x_5 | 18 | | | | | | | | | |
|-----------|-------|-------|-------|----------|----------|----------|----------|----------|----------|----------|
| x_9 | 2 | -1 | | | | | | | | |
| x_{10} | 3 | 4 | 5 | | | | | | | |
| x_{14} | -18 | -19 | -17 | -5 | | | | | | |
| x_{16} | -1 | -1 | 1 | -2 | -9 | | | | | |
| x_{20} | -11 | -11 | -14 | -4 | -46 | -1 | | | | |
| x_{23} | -1 | 0 | -2 | -3 | -1 | -14 | -10 | | | |
| x_{27} | -6 | -5 | -10 | -1 | 0 | 0 | -39 | 1 | | |
| x_{28} | -1 | -1 | -1 | -2 | 0 | -13 | -3 | -23 | -11 | |
| x_{30} | -4 | -4 | -11 | -1 | -9 | 0 | 7 | -2 | -44 | 2 |
| x_{36} | -2 | -2 | -3 | -1 | -1 | 0 | -2 | 0 | -1 | 0 |
| x_{38} | 0 | 0 | 0 | 0 | 0 | -2 | 0 | -3 | 0 | -3 |
| x_{41} | -2 | -2 | -2 | -1 | -1 | 0 | -2 | 0 | -1 | 0 |
| x_{43} | -1 | -1 | -1 | -1 | 0 | -3 | 0 | -5 | 0 | -5 |
| x_{45} | -5 | -5 | -5 | -2 | -3 | -1 | -5 | -2 | -1 | -2 |
| x_{48} | 0 | 0 | 0 | 0 | 0 | 0 | 0 | 0 | 0 | 0 |
| x_{49} | -5 | -5 | -5 | -2 | -3 | -1 | -5 | -2 | -1 | -2 |
| x_{52} | 0 | 0 | 0 | 0 | 0 | 0 | 0 | -1 | 0 | -1 |
| x_{56} | -2 | -2 | -2 | -1 | -1 | -1 | -2 | -1 | -1 | -1 |
| x_{61} | -5 | -5 | -5 | -2 | -3 | -1 | -4 | -2 | -1 | -2 |
| x_{70} | -7 | -9 | 4 | -2 | -6 | 3 | -12 | -2 | -7 | -1 |
| x_{78} | -4 | -4 | -5 | 0 | -9 | 0 | 1 | 1 | -1 | 1 |
| x_{85} | 0 | 0 | -2 | 0 | -2 | 0 | 0 | 0 | 2 | 0 |
| x_{89} | 0 | 0 | 0 | 0 | 0 | 0 | 0 | 0 | 0 | 0 |
| x_{97} | -2 | -2 | -1 | -1 | -1 | -1 | -4 | -1 | -2 | -1 |
| x_{103} | 1 | 1 | 0 | -1 | 1 | -1 | -1 | -1 | 0 | -1 |
| x_{106} | -2 | -2 | 2 | -1 | -1 | 2 | -2 | -1 | -1 | -1 |
| x_{107} | 0 | 0 | 0 | 0 | 0 | 0 | 0 | 0 | 0 | 0 |
| x_{117} | -1 | 0 | 0 | 0 | 0 | 0 | -1 | 0 | -1 | 0 |
| x_{118} | 0 | 0 | 0 | 0 | 0 | 0 | 0 | 0 | 0 | 0 |
| | x_3 | x_5 | x_9 | x_{10} | x_{14} | x_{16} | x_{20} | x_{23} | x_{27} | x_{28} |

| | x_3 | x_5 | x_9 | x_{10} | x_{14} | x_{16} | x_{20} | x_{23} | x_{27} | x_{28} |
|-----------|-------|-------|-------|----------|----------|----------|----------|----------|----------|----------|
| x_{124} | 0 | 0 | 0 | 0 | 0 | 0 | 0 | 0 | 0 | 0 |
| x_{125} | 0 | 0 | 0 | 0 | 0 | 0 | 0 | 0 | 0 | 0 |
| x_{148} | -1 | -1 | -1 | 0 | -1 | 0 | -2 | -1 | 0 | -1 |
| x_{149} | -1 | -1 | -1 | 0 | 0 | 0 | -1 | 0 | 0 | 0 |
| x_{150} | 0 | 0 | 0 | 0 | 0 | 0 | 0 | -1 | 0 | -1 |
| x_{152} | 0 | 0 | 0 | 0 | 0 | 0 | 0 | 0 | 0 | 0 |
| x_{154} | 0 | 0 | 0 | 0 | 0 | 0 | 0 | 0 | 0 | 0 |
| x_{158} | -1 | -1 | -1 | 0 | 0 | 0 | -1 | 0 | 0 | 0 |
| x_{168} | 0 | 0 | 0 | 0 | 0 | 0 | 0 | 0 | 0 | 0 |
| x_{171} | 0 | 0 | 0 | 0 | 0 | 0 | -1 | 0 | 0 | 0 |
| x_{176} | -3 | -3 | -3 | -1 | -4 | -1 | -1 | 0 | -1 | 0 |
| x_{177} | 0 | 0 | 0 | 0 | 0 | 0 | 0 | 0 | 0 | 0 |
| x_{178} | -2 | -2 | -5 | -1 | -3 | 0 | -2 | -1 | 2 | -1 |
| x_{180} | 0 | 0 | 0 | 0 | 0 | 0 | 0 | 0 | 0 | 0 |
| x_{183} | -1 | 0 | 0 | 0 | 0 | 0 | -1 | 0 | 0 | 0 |

| | x_{36} | x_{38} | x_{41} | x_{43} | x_{45} | x_{48} | x_{49} | x_{52} | x_{56} | |
|-----------|----------|----------|----------|----------|----------|----------|----------|----------|----------|----|
| | 0 | 0 | -22 | | | | | | | |
| x_{41} | 0 | -13 | 4 | | | | | | | |
| x_{43} | 0 | 2 | -21 | -20 | | | | | | |
| x_{45} | 0 | -3 | 0 | -6 | 0 | | | | | |
| x_{48} | 0 | -1 | 1 | -4 | 1 | 0 | | | | |
| x_{49} | 0 | -5 | 0 | -4 | -1 | -10 | 0 | | | |
| x_{52} | 0 | 0 | 7 | 0 | 5 | 0 | -7 | 0 | | |
| x_{56} | 0 | -2 | 0 | -2 | -1 | -4 | 0 | -8 | 0 | |
| x_{61} | 0 | -2 | 0 | -2 | 0 | -4 | 0 | -4 | 0 | -2 |
| x_{70} | -5 | -2 | 0 | -1 | 0 | -4 | 1 | -4 | 0 | -2 |
| x_{78} | 3 | 1 | 0 | 1 | 0 | 2 | 0 | 2 | 0 | 1 |
| x_{85} | 2 | 0 | 0 | 0 | 0 | 0 | 0 | 0 | 0 | 0 |
| x_{89} | 0 | 0 | 0 | 0 | 0 | 0 | 0 | 0 | -1 | 0 |
| x_{97} | -1 | -1 | 0 | -1 | 0 | -2 | 0 | -2 | 0 | -1 |
| x_{103} | -1 | -1 | 0 | -1 | 0 | -1 | 0 | -1 | 0 | -1 |
| x_{106} | -1 | -1 | 0 | 0 | 0 | -1 | 0 | -1 | 0 | 0 |
| x_{107} | 0 | 0 | 0 | 0 | 0 | 0 | 0 | 0 | 0 | 0 |
| x_{117} | -1 | 0 | 0 | 0 | 0 | -1 | 0 | -1 | 0 | 0 |
| x_{118} | 0 | 0 | 0 | 0 | 0 | 0 | 0 | 0 | 0 | 0 |
| x_{124} | 0 | 0 | 0 | 0 | 0 | 0 | 0 | 0 | 0 | 0 |
| x_{125} | 0 | 0 | 0 | 0 | 0 | 0 | 0 | 0 | 0 | 0 |
| x_{148} | -2 | 0 | 0 | 0 | 0 | -1 | 1 | -1 | 0 | 0 |
| x_{149} | 0 | 0 | 0 | 0 | 0 | -1 | 0 | -1 | 0 | 0 |
| x_{150} | 0 | 0 | 0 | 0 | 0 | 0 | 1 | 0 | 0 | 0 |
| x_{152} | 0 | 0 | 0 | 0 | 0 | 0 | 0 | 0 | 0 | 0 |
| x_{154} | 0 | 0 | 0 | 0 | 0 | 0 | 0 | -1 | 0 | 0 |
| x_{158} | -1 | 0 | 0 | 0 | 0 | -1 | 0 | -1 | 0 | 0 |
| x_{168} | 0 | 0 | 0 | 0 | 0 | 0 | 0 | 0 | 0 | 0 |
| x_{171} | 0 | 0 | 0 | 0 | 0 | 0 | 0 | 0 | 0 | 0 |
| x_{176} | 1 | -1 | 0 | 0 | 0 | -1 | 0 | -1 | 0 | 0 |
| x_{177} | 0 | 0 | 0 | 0 | 0 | 0 | 0 | 0 | 0 | 0 |
| x_{178} | 2 | -1 | 0 | 0 | 0 | -1 | 0 | -1 | 0 | 0 |
| x_{180} | 0 | 0 | 0 | 0 | 0 | 0 | 0 | 0 | 0 | 0 |
| x_{183} | -1 | 0 | 0 | 0 | 0 | 0 | 0 | 0 | 0 | 0 |

$x_{30} \quad x_{36} \quad x_{38} \quad x_{41} \quad x_{43} \quad x_{45} \quad x_{48} \quad x_{49} \quad x_{52} \quad x_{56}$

| x_{70} | -4 | | | | | | | | | |
|-----------|----|-----|-----|-----|-----|----|-----|----|---|----|
| x_{78} | 2 | -19 | | | | | | | | |
| x_{85} | 0 | -1 | -8 | | | | | | | |
| x_{89} | 0 | -1 | -1 | 0 | | | | | | |
| x_{97} | -2 | 19 | -6 | 0 | 0 | | | | | |
| x_{103} | -1 | -4 | -14 | -1 | 0 | -1 | | | | |
| x_{106} | -1 | 15 | -4 | 0 | 0 | 0 | -1 | | | |
| x_{107} | 0 | -1 | -1 | 0 | 0 | 0 | -3 | 0 | | |
| x_{117} | -1 | 0 | 0 | 0 | -4 | 0 | 0 | 0 | 0 | |
| x_{118} | 0 | 0 | 0 | 0 | 0 | 0 | 0 | 0 | 0 | -1 |
| x_{124} | 0 | 0 | 0 | 0 | 0 | 0 | 0 | 0 | 0 | 3 |
| x_{125} | 0 | 0 | 0 | 0 | 0 | 0 | 0 | 0 | 0 | -1 |
| x_{148} | -1 | 0 | 0 | -5 | 0 | 0 | 0 | 0 | 0 | 0 |
| x_{149} | -1 | -1 | 0 | 0 | -11 | 0 | 0 | 0 | 0 | 10 |
| x_{150} | 0 | 2 | 0 | 0 | 0 | 0 | -1 | 1 | 0 | 0 |
| x_{152} | 0 | 0 | 0 | -1 | 0 | 0 | 0 | 0 | 0 | 0 |
| x_{154} | 0 | 0 | 0 | 0 | -2 | 0 | 0 | 0 | 0 | 0 |
| x_{158} | -1 | -1 | 0 | 0 | -8 | 0 | 0 | 0 | 0 | 47 |
| x_{168} | 0 | -1 | 0 | 0 | 0 | 1 | 0 | 1 | 0 | 0 |
| x_{171} | 0 | 0 | 0 | 0 | -2 | 0 | 0 | 0 | 0 | 35 |
| x_{176} | -1 | -9 | -67 | -3 | 0 | -2 | 10 | -2 | 0 | 0 |
| x_{177} | 0 | 0 | 12 | 0 | 0 | -2 | -58 | 0 | 0 | 0 |
| x_{178} | -1 | -2 | -11 | -64 | -1 | -1 | -1 | -1 | 0 | 0 |
| x_{180} | 0 | 0 | 0 | 0 | -16 | 0 | 0 | 0 | 0 | 8 |
| x_{183} | 0 | 0 | 0 | 0 | -4 | 0 | 0 | 0 | 0 | 39 |
| x_{61} | | | | | | | | | | |
| x_{70} | | | | | | | | | | |
| x_{78} | | | | | | | | | | |
| x_{85} | | | | | | | | | | |
| x_{89} | | | | | | | | | | |
| x_{97} | | | | | | | | | | |
| x_{103} | | | | | | | | | | |
| x_{106} | | | | | | | | | | |
| x_{107} | | | | | | | | | | |
| x_{117} | | | | | | | | | | |

| | | | | | | | | | |
|-----------|-----------|-----------|-----------|-----------|-----------|-----------|-----------|-----------|-----------|
| x_{124} | 0 | | | | | | | | |
| x_{125} | 0 | -1 | | | | | | | |
| x_{148} | 0 | 0 | 0 | | | | | | |
| x_{149} | 0 | 2 | 0 | 0 | | | | | |
| x_{150} | 0 | 0 | 0 | 4 | 0 | | | | |
| x_{152} | 0 | 0 | 0 | 1 | 0 | 1 | | | |
| x_{154} | 0 | 0 | 0 | 2 | -1 | 1 | 0 | | |
| x_{158} | -1 | 3 | -1 | 0 | 25 | 0 | 0 | 0 | |
| x_{168} | 0 | 0 | 0 | 0 | 0 | 0 | 0 | 0 | |
| x_{171} | -1 | 1 | 0 | 0 | 4 | 0 | 0 | 0 | 20 |
| x_{176} | 0 | 0 | 0 | 0 | 0 | 0 | 0 | 0 | 0 |
| x_{177} | 0 | 0 | 0 | 0 | 0 | 0 | 0 | 0 | 0 |
| x_{178} | 0 | 0 | 0 | 0 | 0 | 0 | 0 | 0 | 0 |
| x_{180} | 0 | 2 | 0 | 0 | 10 | 0 | 0 | -1 | 20 |
| x_{183} | -1 | -2 | -1 | 0 | 17 | 0 | 0 | 39 | 0 |
| | x_{118} | x_{124} | x_{125} | x_{148} | x_{149} | x_{150} | x_{152} | x_{154} | x_{158} |
| | x_{176} | 0 | | | | | | | |
| | x_{177} | 0 | -14 | | | | | | |
| | x_{178} | 0 | -4 | 0 | | | | | |
| | x_{180} | 3 | 0 | 0 | 0 | | | | |
| | x_{183} | 17 | 0 | 0 | 0 | 14 | | | |
| | x_{171} | x_{176} | x_{177} | x_{178} | x_{180} | | | | |

A REVIEW GOES HERE – Check our WWW List of Reviews

$$(\Gamma(\tau^+) - \Gamma(\tau^-)) / (\Gamma(\tau^+) + \Gamma(\tau^-))$$

$$\tau^\pm \rightarrow \pi^\pm K_S^0 \nu_\tau \text{ (RATE DIFFERENCE) / (RATE SUM)}$$

| VALUE (%) | DOCUMENT ID | TECN | COMMENT |
|---|-------------|----------|---|
| $-0.36 \pm 0.23 \pm 0.11$ | LEES | 12M BABR | $476 \text{ fb}^{-1} E_{\text{cm}}^{ee} = 10.6 \text{ GeV}$ |

τ^- BRANCHING RATIOS

$$\Gamma(\text{particle}^- \geq 0 \text{ neutrals} \geq 0 K^0 \nu_\tau \text{ ("1-prong")}) / \Gamma_{\text{total}} \quad \Gamma_1 / \Gamma$$

$$\begin{aligned} \Gamma_1 / \Gamma = & (\Gamma_3 + \Gamma_5 + \Gamma_9 + \Gamma_{10} + \Gamma_{14} + \Gamma_{16} + \Gamma_{20} + \Gamma_{23} + \Gamma_{27} + \Gamma_{28} + \Gamma_{30} + \Gamma_{36} + \Gamma_{38} + \Gamma_{41} + \\ & \Gamma_{43} + \Gamma_{45} + \Gamma_{48} + \Gamma_{49} + \Gamma_{50} + \Gamma_{52} + \Gamma_{56} + \Gamma_{57} + 0.7212 \Gamma_{148} + 0.7212 \Gamma_{150} + \\ & 0.7212 \Gamma_{152} + 0.7212 \Gamma_{154} + 0.342 \Gamma_{168} + 0.0828 \Gamma_{176} + 0.0828 \Gamma_{177} + 0.0828 \Gamma_{178}) / \Gamma \end{aligned}$$

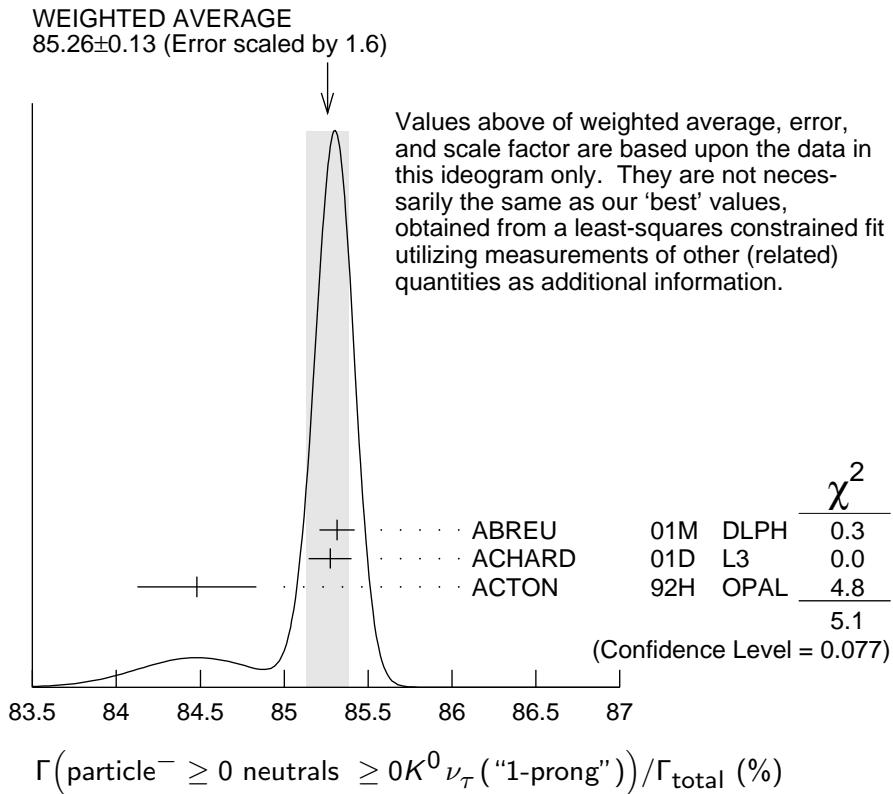
The charged particle here can be e , μ , or hadron. In many analyses, the sum of the topological branching fractions (1, 3, and 5 prongs) is constrained to be unity. Since the 5-prong fraction is very small, the measured 1-prong and 3-prong fractions are highly correlated and cannot be treated as independent quantities in our overall fit.

We arbitrarily choose to use the 3-prong fraction in our fit, and leave the 1-prong fraction out. We do, however, use these 1-prong measurements in our average below. The measurements used only for the average are marked “avg,” whereas “f&a” marks a result used for the fit and the average.

| VALUE (%) | EVTS | DOCUMENT ID | TECN | COMMENT |
|---|---|---------------------|----------|-----------------------|
| 85.24 ±0.06 OUR FIT | | | | |
| 85.26 ±0.13 OUR AVERAGE | Error includes scale factor of 1.6. See the ideogram below. | | | |
| • • • We use the following data for averages but not for fits. • • • | | | | |
| 85.316 ±0.093 ±0.049 | 78k | ¹ ABREU | 01M DLPH | 1992–1995 LEP runs |
| 85.274 ±0.105 ±0.073 | | ² ACHARD | 01D L3 | 1992–1995 LEP runs |
| 84.48 ±0.27 ±0.23 | | ACTON | 92H OPAL | 1990–1991 LEP runs |
| • • • We do not use the following data for averages, fits, limits, etc. • • • | | | | |
| 85.45 $\begin{array}{l} +0.69 \\ -0.73 \end{array}$ ±0.65 | | DECAMP | 92C ALEP | Repl. by SCHAEFEL 05C |

¹ The correlation coefficients between this measurement and the ABREU 01M measurements of $B(\tau \rightarrow \text{3-prong})$ and $B(\tau \rightarrow \text{5-prong})$ are -0.98 and -0.08 respectively.

² The correlation coefficients between this measurement and the ACHARD 01D measurements of $B(\tau \rightarrow \text{"3-prong"})$ and $B(\tau \rightarrow \text{"5-prong"})$ are -0.978 and -0.082 respectively.



$\Gamma(\text{particle}^- \geq 0 K_L^0 \nu_\tau) / \Gamma_{\text{total}}$ Γ_2 / Γ

$$\begin{aligned}\Gamma_2 / \Gamma = & (\Gamma_3 + \Gamma_5 + \Gamma_9 + \Gamma_{10} + \Gamma_{14} + \Gamma_{16} + \Gamma_{20} + \Gamma_{23} + \Gamma_{27} + \Gamma_{28} + \Gamma_{30} + 0.6534 \Gamma_{36} + \\ & 0.6534 \Gamma_{38} + 0.6534 \Gamma_{41} + 0.6534 \Gamma_{43} + 0.6534 \Gamma_{45} + 0.0942 \Gamma_{48} + 0.3069 \Gamma_{49} + \Gamma_{50} + \\ & 0.0942 \Gamma_{52} + 0.3069 \Gamma_{56} + \Gamma_{57} + 0.7212 \Gamma_{148} + 0.7212 \Gamma_{150} + 0.7212 \Gamma_{152} + 0.4712 \Gamma_{154} + \\ & 0.1049 \Gamma_{168} + 0.0828 \Gamma_{176} + 0.0828 \Gamma_{177} + 0.0828 \Gamma_{178}) / \Gamma\end{aligned}$$

| VALUE (%) | EVTS | DOCUMENT ID | TECN | COMMENT |
|-----------|------|-------------|------|---------|
|-----------|------|-------------|------|---------|

84.58 ± 0.06 OUR FIT**85.1 ± 0.4 OUR AVERAGE**

• • • We use the following data for averages but not for fits. • • •

| | | | | |
|------------------|------|---------------------|----------|---|
| 85.6 ± 0.6 ± 0.3 | 3300 | ¹ ADEVA | 91F L3 | $E_{\text{cm}}^{\text{ee}} = 88.3\text{--}94.3 \text{ GeV}$ |
| 84.9 ± 0.4 ± 0.3 | | BEHREND | 89B CELL | $E_{\text{cm}}^{\text{ee}} = 14\text{--}47 \text{ GeV}$ |
| 84.7 ± 0.8 ± 0.6 | | ² AIHARA | 87B TPC | $E_{\text{cm}}^{\text{ee}} = 29 \text{ GeV}$ |

• • • We do not use the following data for averages, fits, limits, etc. • • •

| | | | | |
|--|-----|----------------------|----------|--|
| 86.4 ± 0.3 ± 0.3 | | ABACHI | 89B HRS | $E_{\text{cm}}^{\text{ee}} = 29 \text{ GeV}$ |
| 87.1 ± 1.0 ± 0.7 | | ³ BURCHAT | 87 MRK2 | $E_{\text{cm}}^{\text{ee}} = 29 \text{ GeV}$ |
| 87.2 ± 0.5 ± 0.8 | | SCHMIDKE | 86 MRK2 | $E_{\text{cm}}^{\text{ee}} = 29 \text{ GeV}$ |
| 84.7 ± 1.1 ^{+1.6} _{-1.3} | 169 | ⁴ ALTHOFF | 85 TASS | $E_{\text{cm}}^{\text{ee}} = 34.5 \text{ GeV}$ |
| 86.1 ± 0.5 ± 0.9 | | BARTEL | 85F JADE | $E_{\text{cm}}^{\text{ee}} = 34.6 \text{ GeV}$ |
| 87.8 ± 1.3 ± 3.9 | | ⁵ BERGER | 85 PLUT | $E_{\text{cm}}^{\text{ee}} = 34.6 \text{ GeV}$ |
| 86.7 ± 0.3 ± 0.6 | | FERNANDEZ | 85 MAC | $E_{\text{cm}}^{\text{ee}} = 29 \text{ GeV}$ |

¹ Not independent of ADEVA 91F $\Gamma(h^- h^- h^+ \geq 0 K_L^0 \nu_\tau) / \Gamma_{\text{total}}$ value.

² Not independent of AIHARA 87B $\Gamma(\mu^- \bar{\nu}_\mu \nu_\tau) / \Gamma_{\text{total}}$, $\Gamma(e^- \bar{\nu}_e \nu_\tau) / \Gamma_{\text{total}}$, and $\Gamma(h^- \geq 0 \text{ neutrals} \geq 0 K_L^0 \nu_\tau) / \Gamma_{\text{total}}$ values.

³ Not independent of SCHMIDKE 86 value (also not independent of BURCHAT 87 value for $\Gamma(h^- h^- h^+ \geq 0 \text{ neutrals} \geq 0 K_L^0 \nu_\tau) / \Gamma_{\text{total}}$).

⁴ Not independent of ALTHOFF 85 $\Gamma(\mu^- \bar{\nu}_\mu \nu_\tau) / \Gamma_{\text{total}}$, $\Gamma(e^- \bar{\nu}_e \nu_\tau) / \Gamma_{\text{total}}$, $\Gamma(h^- \geq 0 \text{ neutrals} \geq 0 K_L^0 \nu_\tau) / \Gamma_{\text{total}}$, and $\Gamma(h^- h^- h^+ \geq 0 \text{ neutrals} \geq 0 K_L^0 \nu_\tau) / \Gamma_{\text{total}}$ values.

⁵ Not independent of (1-prong + π^0) and (1-prong + $\geq 1\pi^0$) values.

 $\Gamma(\mu^- \bar{\nu}_\mu \nu_\tau) / \Gamma_{\text{total}}$ Γ_3 / Γ

To minimize the effect of experiments with large systematic errors, we exclude experiments which together would contribute 5% of the weight in the average.

| VALUE (%) | EVTS | DOCUMENT ID | TECN | COMMENT |
|-----------|------|-------------|------|---------|
|-----------|------|-------------|------|---------|

17.39 ± 0.04 OUR FIT**17.33 ± 0.05 OUR AVERAGE**

| | | | | |
|------------------------|-------|-----------------------|----------|--------------------|
| 17.319 ± 0.070 ± 0.032 | 54k | ¹ SCHAEL | 05C ALEP | 1991-1995 LEP runs |
| 17.34 ± 0.09 ± 0.06 | 31.4k | ABBIENDI | 03 OPAL | 1990-1995 LEP runs |
| 17.342 ± 0.110 ± 0.067 | 21.5k | ² ACCIARRI | 01F L3 | 1991-1995 LEP runs |
| 17.325 ± 0.095 ± 0.077 | 27.7k | ABREU | 99X DLPH | 1991-1995 LEP runs |

• • • We use the following data for averages but not for fits. • • •

| | | | | |
|---------------------|--|----------------------------|------|--|
| 17.37 ± 0.08 ± 0.18 | | ³ ANASTASSOV 97 | CLEO | $E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$ |
|---------------------|--|----------------------------|------|--|

• • • We do not use the following data for averages, fits, limits, etc. • • •

| | | | | |
|---------------------|-------|----------|----------|----------------------|
| 17.31 ± 0.11 ± 0.05 | 20.7k | BUSKULIC | 96C ALEP | Repl. by SCHAEL 05C |
| 17.02 ± 0.19 ± 0.24 | 6586 | ABREU | 95T DLPH | Repl. by ABREU 99X |
| 17.36 ± 0.27 | 7941 | AKERS | 95I OPAL | Repl. by ABBIENDI 03 |

| | | | | | | | |
|-------|------------|------------------|------|-----------------------------|-----|------|--|
| 17.6 | ± 0.4 | ± 0.4 | 2148 | ADRIANI | 93M | L3 | Repl. by ACCIARRI 01F |
| 17.4 | ± 0.3 | ± 0.5 | | ⁴ ALBRECHT | 93G | ARG | $E_{\text{cm}}^{\text{ee}} = 9.4\text{--}10.6 \text{ GeV}$ |
| 17.35 | ± 0.41 | ± 0.37 | | DECAMP | 92C | ALEP | 1989-1990 LEP runs |
| 17.7 | ± 0.8 | ± 0.4 | 568 | BEHREND | 90 | CELL | $E_{\text{cm}}^{\text{ee}} = 35 \text{ GeV}$ |
| 17.4 | ± 1.0 | | 2197 | ADEVA | 88 | MRKJ | $E_{\text{cm}}^{\text{ee}} = 14\text{--}16 \text{ GeV}$ |
| 17.7 | ± 1.2 | ± 0.7 | | AIHARA | 87B | TPC | $E_{\text{cm}}^{\text{ee}} = 29 \text{ GeV}$ |
| 18.3 | ± 0.9 | ± 0.8 | | BURCHAT | 87 | MRK2 | $E_{\text{cm}}^{\text{ee}} = 29 \text{ GeV}$ |
| 18.6 | ± 0.8 | ± 0.7 | 558 | ⁵ BARTEL | 86D | JADE | $E_{\text{cm}}^{\text{ee}} = 34.6 \text{ GeV}$ |
| 12.9 | ± 1.7 | $+0.7$ -0.5 | | ALTHOFF | 85 | TASS | $E_{\text{cm}}^{\text{ee}} = 34.5 \text{ GeV}$ |
| 18.0 | ± 0.9 | ± 0.5 | 473 | ⁵ ASH | 85B | MAC | $E_{\text{cm}}^{\text{ee}} = 29 \text{ GeV}$ |
| 18.0 | ± 1.0 | ± 0.6 | | ⁶ BALTRUSAIT..85 | | MRK3 | $E_{\text{cm}}^{\text{ee}} = 3.77 \text{ GeV}$ |
| 19.4 | ± 1.6 | ± 1.7 | 153 | BERGER | 85 | PLUT | $E_{\text{cm}}^{\text{ee}} = 34.6 \text{ GeV}$ |
| 17.6 | ± 2.6 | ± 2.1 | 47 | BEHREND | 83C | CELL | $E_{\text{cm}}^{\text{ee}} = 34 \text{ GeV}$ |
| 17.8 | ± 2.0 | ± 1.8 | | BERGER | 81B | PLUT | $E_{\text{cm}}^{\text{ee}} = 9\text{--}32 \text{ GeV}$ |

¹ See footnote to SCHael 05C $\Gamma(\tau^- \rightarrow e^- \bar{\nu}_e \nu_\tau)/\Gamma_{\text{total}}$ measurement for correlations with other measurements.

² The correlation coefficient between this measurement and the ACCIARRI 01F measurement of $B(\tau^- \rightarrow e^- \bar{\nu}_e \nu_\tau)$ is 0.08.

³ The correlation coefficients between this measurement and the ANASTASSOV 97 measurements of $B(e\bar{\nu}_e \nu_\tau)$, $B(\mu\bar{\nu}_\mu \nu_\tau)/B(e\bar{\nu}_e \nu_\tau)$, $B(h^- \nu_\tau)$, and $B(h^- \nu_\tau)/B(e\bar{\nu}_e \nu_\tau)$ are 0.50, 0.58, 0.50, and 0.08 respectively.

⁴ Not independent of ALBRECHT 92D $\Gamma(\mu^- \bar{\nu}_\mu \nu_\tau)/\Gamma(e^- \bar{\nu}_e \nu_\tau)$ and ALBRECHT 93G $\Gamma(\mu^- \bar{\nu}_\mu \nu_\tau) \times \Gamma(e^- \bar{\nu}_e \nu_\tau)/\Gamma_{\text{total}}^2$ values.

⁵ Modified using $B(e^- \bar{\nu}_e \nu_\tau)/B(\text{"1 prong"})$ and $B(\text{"1 prong"}) = 0.855$.

⁶ Error correlated with BALTRUSAITIS 85 $e\nu\bar{\nu}$ value.

| $\Gamma(\mu^- \bar{\nu}_\mu \nu_\tau \gamma)/\Gamma_{\text{total}}$ | | | | Γ_4/Γ |
|--|------|------------------------|------|---|
| VALUE (%) | EVTS | DOCUMENT ID | TECN | COMMENT |
| 0.368± 0.010 OUR AVERAGE | | | | |
| 0.369 $\pm 0.003 \pm 0.010$ | 16k | ¹ LEES | 15G | BABR $431 \text{ fb}^{-1} E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$ |
| 0.361 $\pm 0.016 \pm 0.035$ | | ² BERGFELD | 00 | CLEO $E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$ |
| • • • We do not use the following data for averages, fits, limits, etc. • • • | | | | |
| 0.30 $\pm 0.04 \pm 0.05$ | 116 | ³ ALEXANDER | 96S | OPAL 1991-1994 LEP runs |
| 0.23 ± 0.10 | 10 | ⁴ WU | 90 | MRK2 $E_{\text{cm}}^{\text{ee}} = 29 \text{ GeV}$ |

¹ LEES 15G impose requirements on detected γ 's corresponding to a τ -rest-frame energy cutoff $E_\gamma^* > 10 \text{ MeV}$.

² BERGFELD 00 impose requirements on detected γ 's corresponding to a τ -rest-frame energy cutoff $E_\gamma^* > 10 \text{ MeV}$. For $E_\gamma^* > 20 \text{ MeV}$, they quote $(3.04 \pm 0.14 \pm 0.30) \times 10^{-3}$.

³ ALEXANDER 96S impose requirements on detected γ 's corresponding to a τ -rest-frame energy cutoff $E_\gamma > 20 \text{ MeV}$.

⁴ WU 90 reports $\Gamma(\mu^- \bar{\nu}_\mu \nu_\tau \gamma)/\Gamma(\mu^- \bar{\nu}_\mu \nu_\tau) = 0.013 \pm 0.006$, which is converted to $\Gamma(\mu^- \bar{\nu}_\mu \nu_\tau \gamma)/\Gamma_{\text{total}}$ using $\Gamma(\mu^- \bar{\nu}_\mu \nu_\tau \gamma)/\Gamma_{\text{total}} = 17.35\%$. Requirements on detected γ 's correspond to a τ rest frame energy cutoff $E_\gamma > 37 \text{ MeV}$.

$\Gamma(e^-\bar{\nu}_e\nu_\tau)/\Gamma_{\text{total}}$ Γ_5/Γ

To minimize the effect of experiments with large systematic errors, we exclude experiments which together would contribute 5% of the weight in the average.

| VALUE (%) | EVTS | DOCUMENT ID | TECN | COMMENT |
|---|-------|-----------------------------|----------|--|
| 17.82 ±0.04 OUR FIT | | | | |
| 17.82 ±0.05 OUR AVERAGE | | | | |
| 17.837±0.072±0.036 | 56k | ¹ SCHABEL | 05C ALEP | 1991–1995 LEP runs |
| 17.806±0.104±0.076 | 24.7k | ² ACCIARRI | 01F L3 | 1991–1995 LEP runs |
| 17.81 ± 0.09 ± 0.06 | 33.1k | ABBIENDI | 99H OPAL | 1991–1995 LEP runs |
| 17.877±0.109±0.110 | 23.3k | ABREU | 99X DLPH | 1991–1995 LEP runs |
| 17.76 ± 0.06 ± 0.17 | | ³ ANASTASSOV | 97 CLEO | $E_{\text{cm}}^{ee} = 10.6 \text{ GeV}$ |
| • • • We do not use the following data for averages, fits, limits, etc. • • • | | | | |
| 17.78 ± 0.10 ± 0.09 | 25.3k | ALEXANDER | 96D OPAL | Repl. by ABBIENDI 99H |
| 17.79 ± 0.12 ± 0.06 | 20.6k | BUSKULIC | 96C ALEP | Repl. by SCHABEL 05C |
| 17.51 ± 0.23 ± 0.31 | 5059 | ABREU | 95T DLPH | Repl.. by ABREU 99X |
| 17.9 ± 0.4 ± 0.4 | 2892 | ADRIANI | 93M L3 | Repl. by ACCIARRI 01F |
| 17.5 ± 0.3 ± 0.5 | | ⁴ ALBRECHT | 93G ARG | $E_{\text{cm}}^{ee} = 9.4\text{--}10.6 \text{ GeV}$ |
| 17.97 ± 0.14 ± 0.23 | 3970 | AKERIB | 92 CLEO | Repl. by ANASTASSOV 97 |
| 19.1 ± 0.4 ± 0.6 | 2960 | ⁵ AMMAR | 92 CLEO | $E_{\text{cm}}^{ee} = 10.5\text{--}10.9 \text{ GeV}$ |
| 18.09 ± 0.45 ± 0.45 | | DECAMP | 92C ALEP | Repl. by SCHABEL 05C |
| 17.0 ± 0.5 ± 0.6 | 1.7k | ABACHI | 90 HRS | $E_{\text{cm}}^{ee} = 29 \text{ GeV}$ |
| 18.4 ± 0.8 ± 0.4 | 644 | BEHREND | 90 CELL | $E_{\text{cm}}^{ee} = 35 \text{ GeV}$ |
| 16.3 ± 0.3 ± 3.2 | | JANSSEN | 89 CBAL | $E_{\text{cm}}^{ee} = 9.4\text{--}10.6 \text{ GeV}$ |
| 18.4 ± 1.2 ± 1.0 | | AIHARA | 87B TPC | $E_{\text{cm}}^{ee} = 29 \text{ GeV}$ |
| 19.1 ± 0.8 ± 1.1 | | BURCHAT | 87 MRK2 | $E_{\text{cm}}^{ee} = 29 \text{ GeV}$ |
| 16.8 ± 0.7 ± 0.9 | 515 | ⁵ BARTEL | 86D JADE | $E_{\text{cm}}^{ee} = 34.6 \text{ GeV}$ |
| 20.4 ± 3.0 ± 1.4 | | ALTHOFF | 85 TASS | $E_{\text{cm}}^{ee} = 34.5 \text{ GeV}$ |
| 17.8 ± 0.9 ± 0.6 | 390 | ⁵ ASH | 85B MAC | $E_{\text{cm}}^{ee} = 29 \text{ GeV}$ |
| 18.2 ± 0.7 ± 0.5 | | ⁶ BALTRUSAIT..85 | 85 MRK3 | $E_{\text{cm}}^{ee} = 3.77 \text{ GeV}$ |
| 13.0 ± 1.9 ± 2.9 | | BERGER | 85 PLUT | $E_{\text{cm}}^{ee} = 34.6 \text{ GeV}$ |
| 18.3 ± 2.4 ± 1.9 | 60 | BEHREND | 83C CELL | $E_{\text{cm}}^{ee} = 34 \text{ GeV}$ |
| 16.0 ± 1.3 | 459 | ⁷ BACINO | 78B DLCO | $E_{\text{cm}}^{ee} = 3.1\text{--}7.4 \text{ GeV}$ |

¹ Correlation matrix for SCHABEL 05C branching fractions, in percent:

- (1) $\Gamma(\tau^- \rightarrow e^-\bar{\nu}_e\nu_\tau)/\Gamma_{\text{total}}$
- (2) $\Gamma(\tau^- \rightarrow \mu^-\bar{\nu}_\mu\nu_\tau)/\Gamma_{\text{total}}$
- (3) $\Gamma(\tau^- \rightarrow \pi^-\nu_\tau)/\Gamma_{\text{total}}$
- (4) $\Gamma(\tau^- \rightarrow \pi^-\pi^0\nu_\tau)/\Gamma_{\text{total}}$
- (5) $\Gamma(\tau^- \rightarrow \pi^-2\pi^0\nu_\tau(\text{ex. } K^0))/\Gamma_{\text{total}}$
- (6) $\Gamma(\tau^- \rightarrow \pi^-3\pi^0\nu_\tau(\text{ex. } K^0))/\Gamma_{\text{total}}$
- (7) $\Gamma(\tau^- \rightarrow h^-4\pi^0\nu_\tau(\text{ex. } K^0, \eta))/\Gamma_{\text{total}}$
- (8) $\Gamma(\tau^- \rightarrow \pi^-\pi^+\pi^-\nu_\tau(\text{ex. } K^0, \omega))/\Gamma_{\text{total}}$
- (9) $\Gamma(\tau^- \rightarrow \pi^-\pi^+\pi^-\pi^0\nu_\tau(\text{ex. } K^0))/\Gamma_{\text{total}}$
- (10) $\Gamma(\tau^- \rightarrow h^-h^-h^+2\pi^0\nu_\tau(\text{ex. } K^0))/\Gamma_{\text{total}}$
- (11) $\Gamma(\tau^- \rightarrow h^-h^-h^+3\pi^0\nu_\tau)/\Gamma_{\text{total}}$
- (12) $\Gamma(\tau^- \rightarrow 3h^-2h^+\nu_\tau(\text{ex. } K^0))/\Gamma_{\text{total}}$

(13) $\Gamma(\tau^- \rightarrow 3h^- 2h^+ \pi^0 \nu_\tau (\text{ex. } K^0)) / \Gamma_{\text{total}}$

| | | | | | | | | | | | |
|------|-----|-----|-----|-----|-----|-----|-----|-----|------|------|-------|
| (1) | (2) | (3) | (4) | (5) | (6) | (7) | (8) | (9) | (10) | (11) | (12) |
| (2) | -20 | | | | | | | | | | |
| (3) | -9 | -6 | | | | | | | | | |
| (4) | -16 | -12 | 2 | | | | | | | | |
| (5) | -5 | -5 | -17 | -37 | | | | | | | |
| (6) | 0 | -4 | -15 | 2 | -27 | | | | | | |
| (7) | -2 | -4 | -24 | -15 | 20 | -47 | | | | | |
| (8) | -14 | -9 | 15 | -5 | -17 | -14 | -8 | | | | |
| (9) | -13 | -12 | -25 | -30 | 4 | -2 | 16 | -15 | | | |
| (10) | 0 | -2 | -23 | -14 | 4 | 10 | 13 | -6 | -17 | | |
| (11) | 1 | 0 | -5 | 1 | 4 | 6 | 0 | -9 | -2 | -11 | |
| (12) | 0 | 1 | 9 | 4 | -8 | -4 | -6 | 9 | -5 | -4 | -2 |
| (13) | 1 | -4 | -3 | -5 | 3 | 2 | -4 | -3 | -1 | 4 | 1 -24 |

² The correlation coefficient between this measurement and the ACCIARRI 01F measurement of $B(\tau^- \rightarrow \mu^- \bar{\nu}_\mu \nu_\tau)$ is 0.08.³ The correlation coefficients between this measurement and the ANASTASSOV 97 measurements of $B(\mu \bar{\nu}_\mu \nu_\tau)$, $B(\mu \bar{\nu}_\mu \nu_\tau) / B(e \bar{\nu}_e \nu_\tau)$, $B(h^- \nu_\tau)$, and $B(h^- \nu_\tau) / B(e \bar{\nu}_e \nu_\tau)$ are 0.50, -0.42, 0.48, and -0.39 respectively.⁴ Not independent of ALBRECHT 92D $\Gamma(\mu^- \bar{\nu}_\mu \nu_\tau) / \Gamma(e^- \bar{\nu}_e \nu_\tau)$ and ALBRECHT 93G $\Gamma(\mu^- \bar{\nu}_\mu \nu_\tau) \times \Gamma(e^- \bar{\nu}_e \nu_\tau) / \Gamma_{\text{total}}^2$ values.⁵ Modified using $B(e^- \bar{\nu}_e \nu_\tau) / B(\text{"1 prong"})$ and $B(\text{"1 prong"})$, = 0.855.⁶ Error correlated with BALTRUSAITIS 85 $\Gamma(\mu^- \bar{\nu}_\mu \nu_\tau) / \Gamma_{\text{total}}$.⁷ BACINO 78B value comes from fit to events with e^\pm and one other nonelectron charged prong. **$\Gamma(\mu^- \bar{\nu}_\mu \nu_\tau) / \Gamma(e^- \bar{\nu}_e \nu_\tau)$** **$\Gamma_3 / \Gamma_5$**

Standard Model prediction including mass effects is 0.9726.

| VALUE (units 10^{-2}) | EVTS | DOCUMENT ID | TECN | COMMENT |
|--------------------------|------|-------------|------|---------|
|--------------------------|------|-------------|------|---------|

97.62 ± 0.28 OUR FIT**97.9 ± 0.4 OUR AVERAGE**97.96 ± 0.16 ± 0.36 731k ¹AUBERT 10F BABR 467 fb^{-1} $E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$ 97.77 ± 0.63 ± 0.87 ²ANASTASSOV 97 CLEO $E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$ 99.7 ± 3.5 ± 4.0 ALBRECHT 92D ARG $E_{\text{cm}}^{\text{ee}} = 9.4\text{--}10.6 \text{ GeV}$ ¹ Correlation matrix for AUBERT 10F branching fractions:(1) $\Gamma(\tau^- \rightarrow \mu^- \bar{\nu}_\mu \nu_\tau) / \Gamma(\tau^- \rightarrow e^- \bar{\nu}_e \nu_\tau)$ (2) $\Gamma(\tau^- \rightarrow \pi^- \nu_\tau) / \Gamma(\tau^- \rightarrow e^- \bar{\nu}_e \nu_\tau)$ (3) $\Gamma(\tau^- \rightarrow K^- \nu_\tau) / \Gamma(\tau^- \rightarrow e^- \bar{\nu}_e \nu_\tau)$

(1) (2)

(2) 0.25

(3) 0.12 0.33

² The correlation coefficients between this measurement and the ANASTASSOV 97 measurements of $B(\mu \bar{\nu}_\mu \nu_\tau)$, $B(e \bar{\nu}_e \nu_\tau)$, $B(h^- \nu_\tau)$, and $B(h^- \nu_\tau) / B(e \bar{\nu}_e \nu_\tau)$ are 0.58, -0.42, 0.07, and 0.45 respectively.

| $\Gamma(e^-\bar{\nu}_e\nu_\tau\gamma)/\Gamma_{\text{total}}$ | | | | | Γ_6/Γ |
|--|------|-------------|----------|--|-------------------|
| VALUE (%) | EVTS | DOCUMENT ID | TECN | COMMENT | |
| 1.84 ± 0.05 OUR AVERAGE | | | | | |
| 1.847 ± 0.015 ± 0.052 | 18k | 1 LEES | 15G BABR | 431 fb ⁻¹ $E_{\text{cm}}^{ee} = 10.6$ GeV | |
| 1.75 ± 0.06 ± 0.17 | | 2 BERGFELD | 00 CLEO | $E_{\text{cm}}^{ee} = 10.6$ GeV | |
| 1 LEES 15G impose requirements on detected γ 's corresponding to a τ -rest-frame energy cutoff $E_\gamma^* > 10$ MeV. | | | | | |
| 2 BERGFELD 00 impose requirements on detected γ 's corresponding to a τ -rest-frame energy cutoff $E_\gamma^* > 10$ MeV. | | | | | |

| $\Gamma(h^- \geq 0 K_L^0 \nu_\tau)/\Gamma_{\text{total}}$ | | | | | Γ_7/Γ |
|---|------|-------------|----------|---|-------------------|
| VALUE (%) | EVTS | DOCUMENT ID | TECN | COMMENT | |
| 12.03 ± 0.05 OUR FIT | | | | | |
| 12.2 ± 0.4 OUR AVERAGE | | | | | |
| 12.47 ± 0.26 ± 0.43 | 2967 | 1 ACCIARRI | 95 L3 | 1992 LEP run | |
| 12.4 ± 0.7 ± 0.7 | 283 | 2 ABREU | 92N DLPH | 1990 LEP run | |
| 12.1 ± 0.7 ± 0.5 | 309 | ALEXANDER | 91D OPAL | 1990 LEP run | |
| • • • We use the following data for averages but not for fits. • • • | | | | | |
| 11.3 ± 0.5 ± 0.8 | 798 | 3 FORD | 87 MAC | $E_{\text{cm}}^{ee} = 29$ GeV | |
| • • • We do not use the following data for averages, fits, limits, etc. • • • | | | | | |
| 12.44 ± 0.11 ± 0.11 | 15k | 4 BUSKULIC | 96 ALEP | Repl. by SCHael 05C | |
| 11.7 ± 0.6 ± 0.8 | | 5 ALBRECHT | 92D ARG | $E_{\text{cm}}^{ee} = 9.4\text{--}10.6$ GeV | |
| 12.98 ± 0.44 ± 0.33 | | 6 DECOMP | 92C ALEP | Repl. by SCHael 05C | |
| 12.3 ± 0.9 ± 0.5 | 1338 | BEHREND | 90 CELL | $E_{\text{cm}}^{ee} = 35$ GeV | |
| 11.1 ± 1.1 ± 1.4 | | 7 BURCHAT | 87 MRK2 | $E_{\text{cm}}^{ee} = 29$ GeV | |
| 12.3 ± 0.6 ± 1.1 | 328 | 8 BARTEL | 86D JADE | $E_{\text{cm}}^{ee} = 34.6$ GeV | |
| 13.0 ± 2.0 ± 4.0 | | BERGER | 85 PLUT | $E_{\text{cm}}^{ee} = 34.6$ GeV | |
| 11.2 ± 1.7 ± 1.2 | 34 | 9 BEHREND | 83C CELL | $E_{\text{cm}}^{ee} = 34$ GeV | |

¹ ACCIARRI 95 with 0.65% added to remove their correction for $\pi^- K_L^0$ backgrounds.

² ABREU 92N with 0.5% added to remove their correction for $K^*(892)^-$ backgrounds.

³ FORD 87 result for $B(\pi^- \nu_\tau)$ with 0.67% added to remove their K^- correction and adjusted for 1992 B ("1 prong").

⁴ BUSKULIC 96 quote $11.78 \pm 0.11 \pm 0.13$ We add 0.66 to undo their correction for unseen K_L^0 and modify the systematic error accordingly.

⁵ Not independent of ALBRECHT 92D $\Gamma(\mu^- \bar{\nu}_\mu \nu_\tau)/\Gamma(e^- \bar{\nu}_e \nu_\tau)$, $\Gamma(\mu^- \bar{\nu}_\mu \nu_\tau) \times \Gamma(e^- \bar{\nu}_e \nu_\tau)$, and $\Gamma(h^- \geq 0 K_L^0 \nu_\tau)/\Gamma(e^- \bar{\nu}_e \nu_\tau)$ values.

⁶ DECOMP 92C quote $B(h^- \geq 0 K_L^0 \geq 0 (K_S^0 \rightarrow \pi^+ \pi^-) \nu_\tau) = 13.32 \pm 0.44 \pm 0.33$. We subtract 0.35 to correct for their inclusion of the K_S^0 decays.

⁷ BURCHAT 87 with 1.1% added to remove their correction for K^- and $K^*(892)^-$ backgrounds.

⁸ BARTEL 86D result for $B(\pi^- \nu_\tau)$ with 0.59% added to remove their K^- correction and adjusted for 1992 B ("1 prong").

⁹ BEHREND 83C quote $B(\pi^- \nu_\tau) = 9.9 \pm 1.7 \pm 1.3$ after subtracting 1.3 ± 0.5 to correct for $B(K^- \nu_\tau)$.

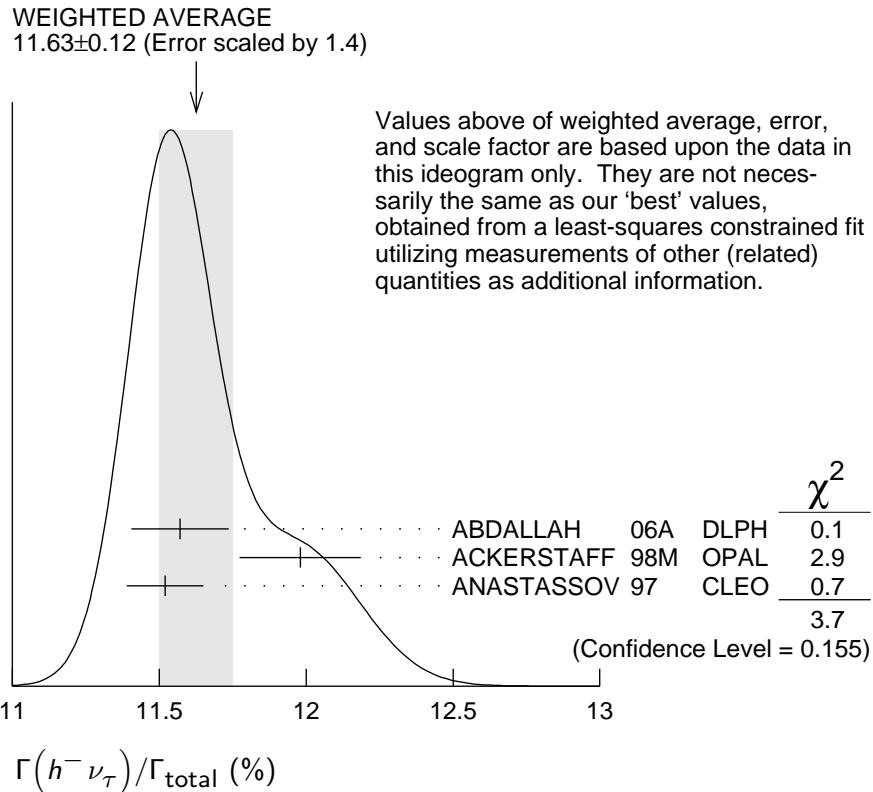
| $\Gamma(h^- \nu_\tau)/\Gamma_{\text{total}}$ | $\Gamma_8/\Gamma = (\Gamma_9 + \Gamma_{10})/\Gamma$ | | | |
|--|---|---|------|---------|
| VALUE (%) | EVTS | DOCUMENT ID | TECN | COMMENT |
| 11.51 ± 0.05 OUR FIT | | | | |
| 11.63 ± 0.12 OUR AVERAGE | | Error includes scale factor of 1.4. See the ideogram below. | | |
| 11.571 ± 0.120 ± 0.114 | 19k | ¹ ABDALLAH 06A DLPH 1992–1995 LEP runs | | |
| 11.98 ± 0.13 ± 0.16 | | ACKERSTAFF 98M OPAL 1991–1995 LEP runs | | |
| 11.52 ± 0.05 ± 0.12 | | ² ANASTASSOV 97 CLEO $E_{\text{cm}}^{ee} = 10.6$ GeV | | |

¹ Correlation matrix for ABDALLAH 06A branching fractions, in percent:

- (1) $\Gamma(\tau^- \rightarrow h^- \nu_\tau)/\Gamma_{\text{total}}$
- (2) $\Gamma(\tau^- \rightarrow h^- \pi^0 \nu_\tau)/\Gamma_{\text{total}}$
- (3) $\Gamma(\tau^- \rightarrow h^- \geq 1\pi^0 \nu_\tau (\text{ex. } K^0))/\Gamma_{\text{total}}$
- (4) $\Gamma(\tau^- \rightarrow h^- 2\pi^0 \nu_\tau (\text{ex. } K^0))/\Gamma_{\text{total}}$
- (5) $\Gamma(\tau^- \rightarrow h^- \geq 3\pi^0 \nu_\tau (\text{ex. } K^0))/\Gamma_{\text{total}}$
- (6) $\Gamma(\tau^- \rightarrow h^- h^- h^+ \nu_\tau (\text{ex. } K^0))/\Gamma_{\text{total}}$
- (7) $\Gamma(\tau^- \rightarrow h^- h^- h^+ \pi^0 \nu_\tau (\text{ex. } K^0))/\Gamma_{\text{total}}$
- (8) $\Gamma(\tau^- \rightarrow h^- h^- h^+ \geq 1\pi^0 \nu_\tau (\text{ex. } K^0))/\Gamma_{\text{total}}$
- (9) $\Gamma(\tau^- \rightarrow h^- h^- h^+ \geq 2\pi^0 \nu_\tau (\text{ex. } K^0))/\Gamma_{\text{total}}$
- (10) $\Gamma(\tau^- \rightarrow 3h^- 2h^+ \nu_\tau (\text{ex. } K^0))/\Gamma_{\text{total}}$
- (11) $\Gamma(\tau^- \rightarrow 3h^- 2h^+ \pi^0 \nu_\tau (\text{ex. } K^0))/\Gamma_{\text{total}}$

| | (1) | (2) | (3) | (4) | (5) | (6) | (7) | (8) | (9) | (10) |
|------|-----|-----|-----|-----|-----|-----|-----|-----|-----|------|
| (2) | -34 | | | | | | | | | |
| (3) | -47 | 56 | | | | | | | | |
| (4) | 6 | -66 | 15 | | | | | | | |
| (5) | -6 | 38 | 11 | -86 | | | | | | |
| (6) | -7 | -8 | 15 | 0 | -2 | | | | | |
| (7) | -2 | -1 | -5 | -3 | 3 | -53 | | | | |
| (8) | -4 | -4 | -13 | -4 | -2 | -56 | 75 | | | |
| (9) | -1 | -1 | -4 | 3 | -6 | 26 | -78 | -16 | | |
| (10) | -1 | -1 | 1 | 0 | 0 | -2 | -3 | -1 | 3 | |
| (11) | 0 | 0 | 0 | 0 | 0 | 1 | 0 | -5 | 5 | -57 |

² The correlation coefficients between this measurement and the ANASTASSOV 97 measurements of $B(\mu \bar{\nu}_\mu \nu_\tau)$, $B(e \bar{\nu}_e \nu_\tau)$, $B(\mu \bar{\nu}_\mu \nu_\tau)/B(e \bar{\nu}_e \nu_\tau)$, and $B(h^- \nu_\tau)/B(e \bar{\nu}_e \nu_\tau)$ are 0.50, 0.48, 0.07, and 0.63 respectively.

 $\Gamma(h^- \nu_\tau)/\Gamma(e^- \bar{\nu}_e \nu_\tau)$

| VALUE (units 10^{-2}) | EVTS | DOCUMENT ID | TECN | COMMENT |
|--------------------------|------|-------------|------|---------|
|--------------------------|------|-------------|------|---------|

 64.62 ± 0.33 OUR FIT **64.0 ± 0.7 OUR AVERAGE** Error includes scale factor of 1.6.

• • • We use the following data for averages but not for fits. • • •

| | | | | | |
|---------------------------|------|----------------------------|----------|--|--|
| $63.33 \pm 0.14 \pm 0.61$ | 394k | ¹ AUBERT | 10F BABR | 467 fb^{-1} | $E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$ |
| $64.84 \pm 0.41 \pm 0.60$ | | ² ANASTASSOV 97 | CLEO | $E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$ | |

¹ Not independent of AUBERT 10F $\Gamma(\tau^- \rightarrow \pi^- \nu_\tau)/\Gamma(\tau^- \rightarrow e^- \bar{\nu}_e \nu_\tau)$ and $\Gamma(\tau^- \rightarrow K^- \nu_\tau)/\Gamma(\tau^- \rightarrow e^- \bar{\nu}_e \nu_\tau)$.

² The correlation coefficients between this measurement and the ANASTASSOV 97 measurements of $B(\mu \bar{\nu}_\mu \nu_\tau)$, $B(e \bar{\nu}_e \nu_\tau)$, $B(\mu \bar{\nu}_\mu \nu_\tau)/B(e \bar{\nu}_e \nu_\tau)$, and $B(h^- \nu_\tau)$ are 0.08, -0.39, 0.45, and 0.63 respectively.

 $\Gamma_8/\Gamma_5 = (\Gamma_9 + \Gamma_{10})/\Gamma_5$ $\Gamma(\pi^- \nu_\tau)/\Gamma_{\text{total}}$

| VALUE (%) | EVTS | DOCUMENT ID | TECN | COMMENT |
|-----------|------|-------------|------|---------|
|-----------|------|-------------|------|---------|

 10.82 ± 0.05 OUR FIT **$10.828 \pm 0.070 \pm 0.078$** 38k ¹ SCHael 05C ALEP 1991-1995 LEP runs

• • • We do not use the following data for averages, fits, limits, etc. • • •

| | | | | |
|---------------------------|------|--------------------------|----------|---|
| $11.06 \pm 0.11 \pm 0.14$ | | ² BUSKULIC 96 | ALEP | Repl. by SCHael 05C |
| $11.7 \pm 0.4 \pm 1.8$ | 1138 | BLOCKER | 82D MRK2 | $E_{\text{cm}}^{\text{ee}} = 3.5\text{--}6.7 \text{ GeV}$ |

¹ See footnote to SCHael 05C $\Gamma(\tau^- \rightarrow e^- \bar{\nu}_e \nu_\tau)/\Gamma_{\text{total}}$ measurement for correlations with other measurements.

² Not independent of BUSKULIC 96 $B(h^- \nu_\tau)$ and $B(K^- \nu_\tau)$ values.

| $\Gamma(\pi^- \nu_\tau)/\Gamma(e^- \bar{\nu}_e \nu_\tau)$ | | Γ_9/Γ_5 | | |
|---|-------------|---------------------|-------------|---|
| <u>VALUE (units 10^{-2})</u> | <u>EVTS</u> | <u>DOCUMENT ID</u> | <u>TECN</u> | <u>COMMENT</u> |
| 60.71 ± 0.32 OUR FIT | | | | |
| $59.45 \pm 0.14 \pm 0.61$ | 369k | ¹ AUBERT | 10F BABR | $467 \text{ fb}^{-1} E_{\text{cm}}^{ee} = 10.6 \text{ GeV}$ |

¹ See footnote to AUBERT 10F $\Gamma(\tau^- \rightarrow \mu^- \bar{\nu}_\mu \nu_\tau)/\Gamma(\tau^- \rightarrow e^- \bar{\nu}_e \nu_\tau)$ for correlations with other measurements.

| $\Gamma(K^- \nu_\tau)/\Gamma_{\text{total}}$ | | Γ_{10}/Γ | | |
|---|-------------|-----------------------|-------------|--|
| <u>VALUE (%)</u> | <u>EVTS</u> | <u>DOCUMENT ID</u> | <u>TECN</u> | <u>COMMENT</u> |
| 0.696 ± 0.010 OUR FIT | | | | |
| 0.685 ± 0.023 OUR AVERAGE | | | | |
| $0.658 \pm 0.027 \pm 0.029$ | | ¹ ABBIENDI | 01J OPAL | 1990–1995 LEP runs |
| $0.696 \pm 0.025 \pm 0.014$ | 2032 | BARATE | 99K ALEP | 1991–1995 LEP runs |
| 0.85 ± 0.18 | 27 | ABREU | 94K DLPH | LEP 1992 Z data |
| $0.66 \pm 0.07 \pm 0.09$ | 99 | BATTLE | 94 CLEO | $E_{\text{cm}}^{ee} \approx 10.6 \text{ GeV}$ |
| $\bullet \bullet \bullet$ We do not use the following data for averages, fits, limits, etc. $\bullet \bullet \bullet$ | | | | |
| $0.72 \pm 0.04 \pm 0.04$ | 728 | BUSKULIC | 96 ALEP | Repl. by BARATE 99K |
| 0.59 ± 0.18 | 16 | MILLS | 84 DLCO | $E_{\text{cm}}^{ee} = 29 \text{ GeV}$ |
| 1.3 ± 0.5 | 15 | BLOCKER | 82B MRK2 | $E_{\text{cm}}^{ee} = 3.9\text{--}6.7 \text{ GeV}$ |

¹ The correlation coefficient between this measurement and the ABBIENDI 01J $B(\tau^- \rightarrow K^- \geq 0 \pi^0 \geq 0 K^0 \geq 0 \gamma \nu_\tau)$ is 0.60.

| $\Gamma(K^- \nu_\tau)/\Gamma(e^- \bar{\nu}_e \nu_\tau)$ | | Γ_{10}/Γ_5 | | |
|---|-------------|------------------------|-------------|---|
| <u>VALUE (units 10^{-2})</u> | <u>EVTS</u> | <u>DOCUMENT ID</u> | <u>TECN</u> | <u>COMMENT</u> |
| 3.91 ± 0.05 OUR FIT | | | | |
| $3.882 \pm 0.032 \pm 0.057$ | 25k | ¹ AUBERT | 10F BABR | $467 \text{ fb}^{-1} E_{\text{cm}}^{ee} = 10.6 \text{ GeV}$ |

¹ See footnote to AUBERT 10F $\Gamma(\tau^- \rightarrow \mu^- \bar{\nu}_\mu \nu_\tau)/\Gamma(\tau^- \rightarrow e^- \bar{\nu}_e \nu_\tau)$ for correlations with other measurements.

| $\Gamma(K^- \nu_\tau)/\Gamma(\pi^- \nu_\tau)$ | | Γ_{10}/Γ_9 | | |
|---|-------------|------------------------|-------------|---|
| <u>VALUE (units 10^{-2})</u> | <u>EVTS</u> | <u>DOCUMENT ID</u> | <u>TECN</u> | <u>COMMENT</u> |
| 6.44 ± 0.09 OUR FIT | | | | |
| $6.531 \pm 0.056 \pm 0.093$ | | ¹ AUBERT | 10F BABR | $467 \text{ fb}^{-1} E_{\text{cm}}^{ee} = 10.6 \text{ GeV}$ |

¹ Not independent of AUBERT 10F $\Gamma(\tau^- \rightarrow \pi^- \nu_\tau)/\Gamma(\tau^- \rightarrow e^- \bar{\nu}_e \nu_\tau)$ and $\Gamma(\tau^- \rightarrow K^- \nu_\tau)/\Gamma(\tau^- \rightarrow e^- \bar{\nu}_e \nu_\tau)$.

| $\Gamma(h^- \geq 1 \text{ neutrals} \nu_\tau)/\Gamma_{\text{total}}$ | | Γ_{11}/Γ | | |
|---|-------------|----------------------|-------------|---|
| <u>VALUE (%)</u> | <u>EVTS</u> | <u>DOCUMENT ID</u> | <u>TECN</u> | <u>COMMENT</u> |
| 37.00 ± 0.09 OUR FIT | | | | |
| $\bullet \bullet \bullet$ We do not use the following data for averages, fits, limits, etc. $\bullet \bullet \bullet$ | | | | |
| $36.14 \pm 0.33 \pm 0.58$ | | ¹ AKERS | 94E OPAL | 1991–1992 LEP runs |
| $38.4 \pm 1.2 \pm 1.0$ | | ² BURCHAT | 87 MRK2 | $E_{\text{cm}}^{ee} = 29 \text{ GeV}$ |
| $42.7 \pm 2.0 \pm 2.9$ | | BERGER | 85 PLUT | $E_{\text{cm}}^{ee} = 34.6 \text{ GeV}$ |

¹ Not independent of ACKERSTAFF 98M $B(h^-\pi^0\nu_\tau)$ and $B(h^-\geq 2\pi^0\nu_\tau)$ values.

² BURCHAT 87 quote for $B(\pi^\pm\geq 1 \text{ neutral}\nu_\tau) = 0.378 \pm 0.012 \pm 0.010$. We add 0.006 to account for contribution from $(K^{*-}\nu_\tau)$ which they fixed at BR = 0.013.

| $\Gamma(h^-\geq 1\pi^0\nu_\tau(\text{ex. } K^0))/\Gamma_{\text{total}}$ | Γ_{12}/Γ |
|--|----------------------|
| $\Gamma_{12}/\Gamma = (\Gamma_{14} + \Gamma_{16} + \Gamma_{20} + \Gamma_{23} + \Gamma_{27} + \Gamma_{28} + \Gamma_{30} + 0.3268\Gamma_{148} + 0.3268\Gamma_{150} + 0.3268\Gamma_{152})/\Gamma$ | |

| VALUE (%) | EVTS | DOCUMENT ID | TECN | COMMENT |
|-----------------------------|------|-------------|------|---------|
| 36.51 ± 0.09 OUR FIT | | | | |

• • • We use the following data for averages but not for fits. • • •

36.641±0.155±0.127 45k ¹ ABDALLAH 06A DLPH 1992–1995 LEP runs

¹ See footnote to ABDALLAH 06A $\Gamma(\tau^-\rightarrow h^-\nu_\tau)/\Gamma_{\text{total}}$ measurement for correlations with other measurements.

| $\Gamma(h^-\pi^0\nu_\tau)/\Gamma_{\text{total}}$ | $\Gamma_{13}/\Gamma = (\Gamma_{14} + \Gamma_{16})/\Gamma$ |
|--|---|
| | |

| VALUE (%) | EVTS | DOCUMENT ID | TECN | COMMENT |
|-----------------------------|------|-------------|------|---------|
| 25.93 ± 0.09 OUR FIT | | | | |

25.73 ± 0.16 OUR AVERAGE

| | | | | | |
|---------------------|------|-----------------------|-----|------|---|
| 25.67 ± 0.01 ± 0.39 | 5.4M | FUJIKAWA | 08 | BELL | $72 \text{ fb}^{-1} E_{\text{cm}}^{\text{ee}}=10.6 \text{ GeV}$ |
| 25.740±0.201±0.138 | 35k | ¹ ABDALLAH | 06A | DLPH | 1992–1995 LEP runs |
| 25.89 ± 0.17 ± 0.29 | | ACKERSTAFF | 98M | OPAL | 1991–1995 LEP runs |
| 25.05 ± 0.35 ± 0.50 | 6613 | ACCIARRI | 95 | L3 | 1992 LEP run |
| 25.87 ± 0.12 ± 0.42 | 51k | ² ARTUSO | 94 | CLEO | $E_{\text{cm}}^{\text{ee}}=10.6 \text{ GeV}$ |

• • • We do not use the following data for averages, fits, limits, etc. • • •

| | | | | | |
|---------------------|------|-----------------------|-----|------|---|
| 25.76 ± 0.15 ± 0.13 | 31k | BUSKULIC | 96 | ALEP | Repl. by SCHAEFEL 05C |
| 25.98 ± 0.36 ± 0.52 | | ³ AKERS | 94E | OPAL | Repl. by ACKER-STAFF 98M |
| 22.9 ± 0.8 ± 1.3 | 283 | ⁴ ABREU | 92N | DLPH | $E_{\text{cm}}^{\text{ee}}=88.2\text{--}94.2 \text{ GeV}$ |
| 23.1 ± 0.4 ± 0.9 | 1249 | ⁵ ALBRECHT | 92Q | ARG | $E_{\text{cm}}^{\text{ee}}=10 \text{ GeV}$ |
| 25.02 ± 0.64 ± 0.88 | 1849 | DECAMP | 92C | ALEP | 1989–1990 LEP runs |
| 22.0 ± 0.8 ± 1.9 | 779 | ANTREASYAN | 91 | CBAL | $E_{\text{cm}}^{\text{ee}}=9.4\text{--}10.6 \text{ GeV}$ |
| 22.6 ± 1.5 ± 0.7 | 1101 | BEHREND | 90 | CELL | $E_{\text{cm}}^{\text{ee}}=35 \text{ GeV}$ |
| 23.1 ± 1.9 ± 1.6 | | BEHREND | 84 | CELL | $E_{\text{cm}}^{\text{ee}}=14,22 \text{ GeV}$ |

¹ See footnote to ABDALLAH 06A $\Gamma(\tau^-\rightarrow h^-\nu_\tau)/\Gamma_{\text{total}}$ measurement for correlations with other measurements.

² ARTUSO 94 reports the combined result from three independent methods, one of which (23% of the $\tau^-\rightarrow h^-\pi^0\nu_\tau$) is normalized to the inclusive one-prong branching fraction, taken as 0.854 ± 0.004 . Renormalization to the present value causes negligible change.

³ AKERS 94E quote $(26.25 \pm 0.36 \pm 0.52) \times 10^{-2}$; we subtract 0.27% from their number to correct for $\tau^-\rightarrow h^-K_L^0\nu_\tau$.

⁴ ABREU 92N with 0.5% added to remove their correction for $K^*(892)^-$ backgrounds.

⁵ ALBRECHT 92Q with 0.5% added to remove their correction for $\tau^-\rightarrow K^*(892)^-\nu_\tau$ background.

| $\Gamma(\pi^- \pi^0 \nu_\tau)/\Gamma_{\text{total}}$ | | | | | Γ_{14}/Γ |
|---|------|-------------------------|----------|--|----------------------|
| VALUE (%) | EVTS | DOCUMENT ID | TECN | COMMENT | |
| 25.49 ± 0.09 OUR FIT | | | | | |
| 25.46 ± 0.12 OUR AVERAGE | | | | | |
| 25.471 ± 0.097 ± 0.085 | 81k | ¹ SCHAEL | 05C ALEP | 1991-1995 LEP runs | |
| • • • We use the following data for averages but not for fits. • • • | | | | | |
| 25.36 ± 0.44 | | ² ARTUSO | 94 CLEO | $E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$ | |
| • • • We do not use the following data for averages, fits, limits, etc. • • • | | | | | |
| 25.30 ± 0.15 ± 0.13 | | ³ BUSKULIC | 96 ALEP | Repl. by SCHAEL 05C | |
| 21.5 ± 0.4 ± 1.9 | 4400 | ^{4,5} ALBRECHT | 88L ARG | $E_{\text{cm}}^{\text{ee}} = 10 \text{ GeV}$ | |
| 23.0 ± 1.3 ± 1.7 | 582 | ADLER | 87B MRK3 | $E_{\text{cm}}^{\text{ee}} = 3.77 \text{ GeV}$ | |
| 25.8 ± 1.7 ± 2.5 | | ⁶ BURCHAT | 87 MRK2 | $E_{\text{cm}}^{\text{ee}} = 29 \text{ GeV}$ | |
| 22.3 ± 0.6 ± 1.4 | 629 | ⁵ YELTON | 86 MRK2 | $E_{\text{cm}}^{\text{ee}} = 29 \text{ GeV}$ | |

¹ See footnote to SCHAEL 05C $\Gamma(\tau^- \rightarrow e^- \bar{\nu}_e \nu_\tau)/\Gamma_{\text{total}}$ measurement for correlations with other measurements.

² Not independent of ARTUSO 94 $B(h^- \pi^0 \nu_\tau)$ and BATTLE 94 $B(K^- \pi^0 \nu_\tau)$ values.

³ Not independent of BUSKULIC 96 $B(h^- \pi^0 \nu_\tau)$ and $B(K^- \pi^0 \nu_\tau)$ values.

⁴ The authors divide by $(\Gamma_3 + \Gamma_5 + \Gamma_9 + \Gamma_{10})/\Gamma = 0.467$ to obtain this result.

⁵ Experiment had no hadron identification. Kaon corrections were made, but insufficient information is given to permit their removal.

⁶ BURCHAT 87 value is not independent of YELTON 86 value. Nonresonant decays included.

| $\Gamma(\pi^- \pi^0 \text{non-}\rho(770) \nu_\tau)/\Gamma_{\text{total}}$ | | | | | Γ_{15}/Γ |
|---|------|----------------------|---------|---|----------------------|
| VALUE (%) | EVTS | DOCUMENT ID | TECN | COMMENT | |
| 0.3±0.1±0.3 | | ¹ BEHREND | 84 CELL | $E_{\text{cm}}^{\text{ee}} = 14,22 \text{ GeV}$ | |

¹ BEHREND 84 assume a flat nonresonant mass distribution down to the $\rho(770)$ mass, using events with mass above 1300 to set the level.

| $\Gamma(K^- \pi^0 \nu_\tau)/\Gamma_{\text{total}}$ | | | | | Γ_{16}/Γ |
|---|------|-------------|-----------|--|----------------------|
| VALUE (%) | EVTS | DOCUMENT ID | TECN | COMMENT | |
| 0.433±0.015 OUR FIT | | | | | |
| 0.426±0.016 OUR AVERAGE | | | | | |
| 0.416 ± 0.003 ± 0.018 | 78k | AUBERT | 07AP BABR | $230 \text{ fb}^{-1} E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$ | |
| 0.471 ± 0.059 ± 0.023 | 360 | ABBIENDI | 04J OPAL | 1991-1995 LEP runs | |
| 0.444 ± 0.026 ± 0.024 | 923 | BARATE | 99K ALEP | 1991-1995 LEP runs | |
| 0.51 ± 0.10 ± 0.07 | 37 | BATTLE | 94 CLEO | $E_{\text{cm}}^{\text{ee}} \approx 10.6 \text{ GeV}$ | |
| • • • We do not use the following data for averages, fits, limits, etc. • • • | | | | | |
| 0.52 ± 0.04 ± 0.05 | 395 | BUSKULIC | 96 ALEP | Repl. by BARATE 99K | |

| $\Gamma(h^- \geq 2\pi^0 \nu_\tau)/\Gamma_{\text{total}}$ | | | | | Γ_{17}/Γ |
|--|------|-------------|----------|--------------------|----------------------|
| VALUE (%) | EVTS | DOCUMENT ID | TECN | COMMENT | |
| 10.81±0.09 OUR FIT | | | | | |
| 9.91±0.31±0.27 | | ACKERSTAFF | 98M OPAL | 1991–1995 LEP runs | |

• • • We do not use the following data for averages, fits, limits, etc. • • •

| | | | | | |
|--------------------------|-----|----------------------|-----|------|--|
| $9.89 \pm 0.34 \pm 0.55$ | | ¹ AKERS | 94E | OPAL | Repl. by ACKER-STAFF 98M |
| $14.0 \pm 1.2 \pm 0.6$ | 938 | ² BEHREND | 90 | CELL | $E_{\text{cm}}^{\text{ee}} = 35 \text{ GeV}$ |
| $12.0 \pm 1.4 \pm 2.5$ | | ³ BURCHAT | 87 | MRK2 | $E_{\text{cm}}^{\text{ee}} = 29 \text{ GeV}$ |
| $13.9 \pm 2.0 \pm 1.9$ | | ⁴ AIHARA | 86E | TPC | $E_{\text{cm}}^{\text{ee}} = 29 \text{ GeV}$ |

¹ AKERS 94E not independent of AKERS 94E $B(h^- \geq 1\pi^0\nu_\tau)$ and $B(h^- \pi^0\nu_\tau)$ measurements.

² No independent of BEHREND 90 $\Gamma(h^- 2\pi^0\nu_\tau)$ (exp. K^0) and $\Gamma(h^- \geq 3\pi^0\nu_\tau)$.

³ Error correlated with BURCHAT 87 $\Gamma(\rho^-\nu_e)/\Gamma(\text{total})$ value.

⁴ AIHARA 86E (TPC) quote $B(2\pi^0\pi^-\nu_\tau) + 1.6B(3\pi^0\pi^-\nu_\tau) + 1.1B(\pi^0\eta\pi^-\nu_\tau)$.

$\Gamma(h^- 2\pi^0\nu_\tau)/\Gamma_{\text{total}}$

$$\Gamma_{18}/\Gamma = (\Gamma_{20} + \Gamma_{23} + 0.15344\Gamma_{36} + 0.15344\Gamma_{38})/\Gamma$$

Γ_{18}/Γ

| VALUE (%) | EVTS | DOCUMENT ID | TECN | COMMENT |
|----------------------------|------|-------------|------|---------|
| 9.48 ± 0.10 OUR FIT | | | | |

• • • We do not use the following data for averages, fits, limits, etc. • • •

| | | | | | |
|--------------------------|-----|-----------------------|----|------|----------------------|
| $9.48 \pm 0.13 \pm 0.10$ | 12k | ¹ BUSKULIC | 96 | ALEP | Repl. by SCHAEEL 05C |
|--------------------------|-----|-----------------------|----|------|----------------------|

¹ BUSKULIC 96 quote $9.29 \pm 0.13 \pm 0.10$. We add 0.19 to undo their correction for $\tau^- \rightarrow h^- K^0\nu_\tau$.

$\Gamma(h^- 2\pi^0\nu_\tau(\text{ex. } K^0))/\Gamma_{\text{total}}$

$$\Gamma_{19}/\Gamma = (\Gamma_{20} + \Gamma_{23})/\Gamma$$

Γ_{19}/Γ

| VALUE (%) | EVTS | DOCUMENT ID | TECN | COMMENT |
|--------------------------------|------|-------------|------|---------|
| 9.32 ± 0.10 OUR FIT | | | | |
| 9.17 ± 0.27 OUR AVERAGE | | | | |

| | | | | | |
|-----------------------------|------|-----------------------|-----|------|--------------------|
| $9.498 \pm 0.320 \pm 0.275$ | 9.5k | ¹ ABDALLAH | 06A | DLPH | 1992–1995 LEP runs |
| $8.88 \pm 0.37 \pm 0.42$ | 1060 | ACCIARRI | 95 | L3 | 1992 LEP run |

• • • We use the following data for averages but not for fits. • • •

| | | | | | |
|--------------------------|--|-----------------------|----|------|--|
| $8.96 \pm 0.16 \pm 0.44$ | | ² PROCARIO | 93 | CLEO | $E_{\text{cm}}^{\text{ee}} \approx 10.6 \text{ GeV}$ |
|--------------------------|--|-----------------------|----|------|--|

• • • We do not use the following data for averages, fits, limits, etc. • • •

| | | | | | |
|---------------------------|-----|-------------------------|-----|------|--|
| $10.38 \pm 0.66 \pm 0.82$ | 809 | ³ DECAMP | 92C | ALEP | Repl. by SCHAEEL 05C |
| $5.7 \pm 0.5 \pm 1.7$ | 133 | ⁴ ANTREASYAN | 91 | CBAL | $E_{\text{cm}}^{\text{ee}} = 9.4\text{--}10.6 \text{ GeV}$ |
| $10.0 \pm 1.5 \pm 1.1$ | 333 | ⁵ BEHREND | 90 | CELL | $E_{\text{cm}}^{\text{ee}} = 35 \text{ GeV}$ |
| $8.7 \pm 0.4 \pm 1.1$ | 815 | ⁶ BAND | 87 | MAC | $E_{\text{cm}}^{\text{ee}} = 29 \text{ GeV}$ |
| $6.2 \pm 0.6 \pm 1.2$ | | ⁷ GAN | 87 | MRK2 | $E_{\text{cm}}^{\text{ee}} = 29 \text{ GeV}$ |
| $6.0 \pm 3.0 \pm 1.8$ | | BEHREND | 84 | CELL | $E_{\text{cm}}^{\text{ee}} = 14,22 \text{ GeV}$ |

¹ See footnote to ABDALLAH 06A $\Gamma(\tau^- \rightarrow h^- \nu_\tau)/\Gamma_{\text{total}}$ measurement for correlations with other measurements.

² PROCARIO 93 entry is obtained from $B(h^- 2\pi^0\nu_\tau)/B(h^- \pi^0\nu_\tau)$ using ARTUSO 94 result for $B(h^- \pi^0\nu_\tau)$.

³ We subtract 0.0015 to account for $\tau^- \rightarrow K^*(892)^-\nu_\tau$ contribution.

⁴ ANTREASYAN 91 subtract 0.001 to account for the $\tau^- \rightarrow K^*(892)^-\nu_\tau$ contribution.

⁵ BEHREND 90 subtract 0.002 to account for the $\tau^- \rightarrow K^*(892)^-\nu_\tau$ contribution.

⁶ BAND 87 assume $B(\pi^- 3\pi^0\nu_\tau) = 0.01$ and $B(\pi^- \pi^0\eta\nu_\tau) = 0.005$.

⁷ GAN 87 analysis use photon multiplicity distribution.

$$\frac{\Gamma(h^- 2\pi^0 \nu_\tau (\text{ex. } K^0))}{\Gamma(h^- \pi^0 \nu_\tau)} / \frac{\Gamma(h^- \pi^0 \nu_\tau)}{\Gamma_{19}/\Gamma_{13}} = \frac{(\Gamma_{20} + \Gamma_{23})}{(\Gamma_{14} + \Gamma_{16})}$$

 Γ_{19}/Γ_{13}

| VALUE (units 10^{-2}) | DOCUMENT ID | TECN | COMMENT |
|--|-------------|------|---------|
| 36.0 ± 0.4 OUR FIT | | | |

 $34.2 \pm 0.6 \pm 1.6$ ¹ PROCARIO 93 CLEO $E_{\text{cm}}^{ee} \approx 10.6$ GeV

¹ PROCARIO 93 quote $0.345 \pm 0.006 \pm 0.016$ after correction for 2 kaon backgrounds assuming $B(K^{*-} \nu_\tau) = 1.42 \pm 0.18\%$ and $B(h^- K^0 \pi^0 \nu_\tau) = 0.48 \pm 0.48\%$. We multiply by 0.990 ± 0.010 to remove these corrections to $B(h^- \pi^0 \nu_\tau)$.

$$\frac{\Gamma(\pi^- 2\pi^0 \nu_\tau (\text{ex. } K^0))}{\Gamma_{\text{total}}} / \frac{\Gamma_{20}}{\Gamma}$$

| VALUE (%) | EVTS | DOCUMENT ID | TECN | COMMENT |
|---|------|-------------|------|---------|
| 9.26 ± 0.10 OUR FIT | | | | |

 $9.239 \pm 0.086 \pm 0.090$ 31k ¹ SCHAEL 05C ALEP 1991-1995 LEP runs

• • • We do not use the following data for averages, fits, limits, etc. • • •

 $9.21 \pm 0.13 \pm 0.11$ ² BUSKULIC 96 ALEP Repl. by SCHAEL 05C

¹ See footnote to SCHAEL 05C $\Gamma(\tau^- \rightarrow e^- \bar{\nu}_e \nu_\tau)/\Gamma_{\text{total}}$ measurement for correlations with other measurements.

² Not independent of BUSKULIC 96 $B(h^- 2\pi^0 \nu_\tau (\text{ex. } K^0))$ and $B(K^- 2\pi^0 \nu_\tau (\text{ex. } K^0))$ values.

$$\frac{\Gamma(\pi^- 2\pi^0 \nu_\tau (\text{ex. } K^0), \text{scalar})}{\Gamma(\pi^- 2\pi^0 \nu_\tau (\text{ex. } K^0))} / \frac{\Gamma_{21}}{\Gamma_{20}}$$

| VALUE | CL% | DOCUMENT ID | TECN | COMMENT |
|------------------|-----|-------------------------|------|---|
| <0.094 | 95 | ¹ BROWDER 00 | CLEO | $4.7 \text{ fb}^{-1} E_{\text{cm}}^{ee} = 10.6 \text{ GeV}$ |

¹ Model-independent limit from structure function analysis on contribution to $B(\tau^- \rightarrow \pi^- 2\pi^0 \nu_\tau (\text{ex. } K^0))$ from scalars.

$$\frac{\Gamma(\pi^- 2\pi^0 \nu_\tau (\text{ex. } K^0), \text{vector})}{\Gamma(\pi^- 2\pi^0 \nu_\tau (\text{ex. } K^0))} / \frac{\Gamma_{22}}{\Gamma_{20}}$$

| VALUE | CL% | DOCUMENT ID | TECN | COMMENT |
|------------------|-----|-------------------------|------|---|
| <0.073 | 95 | ¹ BROWDER 00 | CLEO | $4.7 \text{ fb}^{-1} E_{\text{cm}}^{ee} = 10.6 \text{ GeV}$ |

¹ Model-independent limit from structure function analysis on contribution to $B(\tau^- \rightarrow \pi^- 2\pi^0 \nu_\tau (\text{ex. } K^0))$ from vectors.

$$\frac{\Gamma(K^- 2\pi^0 \nu_\tau (\text{ex. } K^0))}{\Gamma_{\text{total}}} / \frac{\Gamma_{23}}{\Gamma}$$

| VALUE (units 10^{-4}) | EVTS | DOCUMENT ID | TECN | COMMENT |
|---|------|-------------|------|---------|
| 6.5 ± 2.2 OUR FIT | | | | |

 5.8 ± 2.4 OUR AVERAGE $5.6 \pm 2.0 \pm 1.5$ 131 BARATE 99K ALEP 1991–1995 LEP runs $9 \pm 10 \pm 3$ 3 ¹ BATTLE 94 CLEO $E_{\text{cm}}^{ee} \approx 10.6$ GeV

• • • We do not use the following data for averages, fits, limits, etc. • • •

 $8 \pm 2 \pm 2$ 59 BUSKULIC 96 ALEP Repl. by BARATE 99K

¹ BATTLE 94 quote $(14 \pm 10 \pm 3) \times 10^{-4}$ or $< 30 \times 10^{-4}$ at 90% CL. We subtract $(5 \pm 2) \times 10^{-4}$ to account for $\tau^- \rightarrow K^- (K^0 \rightarrow \pi^0 \pi^0) \nu_\tau$ background.

$$\frac{\Gamma(h^- \geq 3\pi^0 \nu_\tau)}{\Gamma_{\text{total}}} / \frac{\Gamma_{24}}{\Gamma}$$

$$\Gamma_{24}/\Gamma = (\Gamma_{27} + \Gamma_{28} + \Gamma_{30} + 0.15344\Gamma_{41} + 0.15344\Gamma_{43} + 0.0942\Gamma_{48} + 0.0942\Gamma_{52} + 0.3268\Gamma_{148} + 0.3268\Gamma_{150} + 0.3268\Gamma_{152} + 0.0501\Gamma_{154})/\Gamma$$

| VALUE (%) | EVTS | DOCUMENT ID | TECN | COMMENT |
|---|------|-------------|------|---------|
| 1.34 ± 0.07 OUR FIT | | | | |

• • • We do not use the following data for averages, fits, limits, etc. • • •

 $1.53 \pm 0.40 \pm 0.46$ 186 DECAMP 92C ALEP Repl. by SCHAEL 05C $3.2 \pm 1.0 \pm 1.0$ BEHREND 90 CELL $E_{\text{cm}}^{ee} = 35$ GeV

$$\Gamma(h^- \geq 3\pi^0 \nu_\tau (\text{ex. } K^0)) / \Gamma_{\text{total}} \quad \Gamma_{25}/\Gamma$$

$$\Gamma_{25}/\Gamma = (\Gamma_{27} + \Gamma_{28} + \Gamma_{30} + 0.3268\Gamma_{148} + 0.3268\Gamma_{150} + 0.3268\Gamma_{152}) / \Gamma$$

| <u>VALUE (%)</u> | <u>EVTS</u> | <u>DOCUMENT ID</u> | <u>TECN</u> | <u>COMMENT</u> |
|------------------------------|-------------|-----------------------|-------------|--------------------|
| 1.25 ± 0.07 OUR FIT | | | | |
| 1.403 ± 0.214 ± 0.224 | 1.1k | ¹ ABDALLAH | 06A DLPH | 1992–1995 LEP runs |

¹ See footnote to ABDALLAH 06A $\Gamma(\tau^- \rightarrow h^- \nu_\tau) / \Gamma_{\text{total}}$ measurement for correlations with other measurements.

$$\Gamma(h^- 3\pi^0 \nu_\tau) / \Gamma_{\text{total}} \quad \Gamma_{26}/\Gamma$$

$$\Gamma_{26}/\Gamma = (\Gamma_{27} + \Gamma_{28} + 0.15344\Gamma_{41} + 0.15344\Gamma_{43} + 0.3268\Gamma_{150}) / \Gamma$$

| <u>VALUE (%)</u> | <u>EVTS</u> | <u>DOCUMENT ID</u> | <u>TECN</u> | <u>COMMENT</u> |
|----------------------------|-------------|--------------------|-------------|----------------|
| 1.18 ± 0.07 OUR FIT | | | | |

1.21 ± 0.17 OUR AVERAGE Error includes scale factor of 1.2.

1.70 ± 0.24 ± 0.38 293 ACCIARRI 95 L3 1992 LEP run

• • • We use the following data for averages but not for fits. • • •

1.15 ± 0.08 ± 0.13 ¹ PROCARIO 93 CLEO $E_{\text{cm}}^{\text{ee}} \approx 10.6 \text{ GeV}$

• • • We do not use the following data for averages, fits, limits, etc. • • •

1.24 ± 0.09 ± 0.11 2.3k ² BUSKULIC 96 ALEP Repl. by SCHAEEL 05C

0.0 $+1.4$ $+1.1$ ³ GAN 87 MRK2 $E_{\text{cm}}^{\text{ee}} = 29 \text{ GeV}$

¹ PROCARIO 93 entry is obtained from $B(h^- 3\pi^0 \nu_\tau) / B(h^- \pi^0 \nu_\tau)$ using ARTUSO 94 result for $B(h^- \pi^0 \nu_\tau)$.

² BUSKULIC 96 quote $B(h^- 3\pi^0 \nu_\tau (\text{ex. } K^0)) = 1.17 \pm 0.09 \pm 0.11$. We add 0.07 to remove their correction for K^0 backgrounds.

³ Highly correlated with GAN 87 $\Gamma(\eta \pi^- \pi^0 \nu_\tau) / \Gamma_{\text{total}}$ value. Authors quote $B(\pi^\pm 3\pi^0 \nu_\tau) + 0.67B(\pi^\pm \eta \pi^0 \nu_\tau) = 0.047 \pm 0.010 \pm 0.011$.

$$\Gamma(h^- 3\pi^0 \nu_\tau) / \Gamma(h^- \pi^0 \nu_\tau) \quad \Gamma_{26}/\Gamma_{13}$$

$$\Gamma_{26}/\Gamma_{13} = (\Gamma_{27} + \Gamma_{28} + 0.15344\Gamma_{41} + 0.15344\Gamma_{43} + 0.3268\Gamma_{150}) / (\Gamma_{14} + \Gamma_{16})$$

| <u>VALUE (units 10^{-2})</u> | <u>DOCUMENT ID</u> | <u>TECN</u> | <u>COMMENT</u> |
|---|--------------------|-------------|----------------|
| 4.54 ± 0.28 OUR FIT | | | |

4.4 ± 0.3 ± 0.5 ¹ PROCARIO 93 CLEO $E_{\text{cm}}^{\text{ee}} \approx 10.6 \text{ GeV}$

¹ PROCARIO 93 quote $0.041 \pm 0.003 \pm 0.005$ after correction for 2 kaon backgrounds assuming $B(K^{*-} \nu_\tau) = 1.42 \pm 0.18\%$ and $B(h^- K^0 \pi^0 \nu_\tau) = 0.48 \pm 0.48\%$. We add 0.003 ± 0.003 and multiply the sum by 0.990 ± 0.010 to remove these corrections.

$$\Gamma(\pi^- 3\pi^0 \nu_\tau (\text{ex. } K^0)) / \Gamma_{\text{total}} \quad \Gamma_{27}/\Gamma$$

| <u>VALUE (%)</u> | <u>EVTS</u> | <u>DOCUMENT ID</u> | <u>TECN</u> | <u>COMMENT</u> |
|------------------------------|-------------|----------------------|-------------|--------------------|
| 1.04 ± 0.07 OUR FIT | | | | |
| 0.977 ± 0.069 ± 0.058 | 6.1k | ¹ SCHAEEL | 05C ALEP | 1991–1995 LEP runs |

¹ See footnote to SCHAEEL 05C $\Gamma(\tau^- \rightarrow e^- \bar{\nu}_e \nu_\tau) / \Gamma_{\text{total}}$ measurement for correlations with other measurements.

$$\Gamma(K^- 3\pi^0 \nu_\tau (\text{ex. } K^0, \eta)) / \Gamma_{\text{total}} \quad \Gamma_{28}/\Gamma$$

| <u>VALUE (units 10^{-4})</u> | <u>EVTS</u> | <u>DOCUMENT ID</u> | <u>TECN</u> | <u>COMMENT</u> |
|---|-------------|--------------------|-------------|----------------|
| 4.8 ± 2.1 OUR FIT | | | | |

3.7 ± 2.1 ± 1.1 22 BARATE 99K ALEP 1991–1995 LEP runs

• • • We do not use the following data for averages, fits, limits, etc. • • •

5 ± 13 ¹ BUSKULIC 94E ALEP Repl. by BARATE 99K

¹ BUSKULIC 94E quote $B(K^- \geq 0\pi^0 \geq 0K^0 \nu_\tau) = [B(K^- \nu_\tau) + B(K^- \pi^0 \nu_\tau) + B(K^- K^0 \nu_\tau) + B(K^- \pi^0 \pi^0 \nu_\tau) + B(K^- \pi^0 K^0 \nu_\tau)] = (5 \pm 13) \times 10^{-4}$ accounting for common systematic errors in BUSKULIC 94E and BUSKULIC 94F measurements of these modes. We assume $B(K^- \geq 2K^0 \nu_\tau)$ and $B(K^- \geq 4\pi^0 \nu_\tau)$ are negligible.

$$\Gamma(h^- 4\pi^0 \nu_\tau (\text{ex.} K^0)) / \Gamma_{\text{total}} \quad \Gamma_{29}/\Gamma$$

$$\Gamma_{29}/\Gamma = (\Gamma_{30} + 0.3268\Gamma_{148} + 0.3268\Gamma_{152}) / \Gamma$$

| VALUE (%) | EVTS | DOCUMENT ID | TECN | COMMENT |
|---|------|--|------|---------|
| 0.16 ± 0.04 OUR FIT | | | | |
| 0.16 ± 0.05 ± 0.05 | | ¹ PROCARIO 93 CLEO $E_{\text{cm}}^{\text{ee}} \approx 10.6 \text{ GeV}$ | | |
| • • • We do not use the following data for averages, fits, limits, etc. • • • | | | | |
| 0.16 ± 0.04 ± 0.09 | 232 | ² BUSKULIC 96 ALEP Repl. by SCHAEEL 05C | | |
| 1 PROCARIO 93 quotes $B(h^- 4\pi^0 \nu_\tau) / B(h^- \pi^0 \nu_\tau) = 0.006 \pm 0.002 \pm 0.002$. We multiply by the ARTUSO 94 result for $B(h^- \pi^0 \nu_\tau)$ to obtain $B(h^- 4\pi^0 \nu_\tau)$. PROCARIO 93 assume $B(h^- \geq 5\pi^0 \nu_\tau)$ is small and do not correct for it. | | | | |
| 2 BUSKULIC 96 quote result for $\tau^- \rightarrow h^- \geq 4\pi^0 \nu_\tau$. We assume $B(h^- \geq 5\pi^0 \nu_\tau)$ is negligible. | | | | |

$$\Gamma(h^- 4\pi^0 \nu_\tau (\text{ex.} K^0, \eta)) / \Gamma_{\text{total}} \quad \Gamma_{30}/\Gamma$$

| VALUE (%) | EVTS | DOCUMENT ID | TECN | COMMENT |
|---|------|--|------|---------|
| 0.11 ± 0.04 OUR FIT | | | | |
| 0.112 ± 0.037 ± 0.035 | 957 | ¹ SCHAEEL 05C ALEP 1991-1995 LEP runs | | |
| 1 See footnote to SCHAEEL 05C $\Gamma(\tau^- \rightarrow e^- \bar{\nu}_e \nu_\tau) / \Gamma_{\text{total}}$ measurement for correlations with other measurements. | | | | |

$$\Gamma(a_1(1260)\nu_\tau \rightarrow \pi^- \gamma \nu_\tau) / \Gamma_{\text{total}} \quad \Gamma_{31}/\Gamma = (0.0021\Gamma_{20} + 0.0021\Gamma_{70}) / \Gamma$$

The uncertainty on $\Gamma(\tau^- \rightarrow a_1(1260)\nu_\tau \rightarrow \pi^- \gamma \nu_\tau) / \Gamma_{\text{total}}$ takes into account the non-negligible contribution from the uncertainty of the coefficient of the relationship that defines $\Gamma(\tau^- \rightarrow a_1(1260)\nu_\tau \rightarrow \pi^- \gamma \nu_\tau)$ in terms of $\Gamma(\tau^- \rightarrow \pi^- 2\pi^0 \nu_\tau (\text{ex.} K^0))$ and $\Gamma(\tau^- \rightarrow \pi^- \pi^+ \pi^- \nu_\tau (\text{ex.} K^0, \omega))$.

| VALUE (units 10^{-4}) | DOCUMENT ID |
|--------------------------|-------------|
| 3.8 ± 1.5 OUR FIT | |

$$\Gamma(K^- \geq 0\pi^0 \geq 0K^0 \geq 0\gamma \nu_\tau) / \Gamma_{\text{total}} \quad \Gamma_{32}/\Gamma$$

$$\Gamma_{32}/\Gamma = (\Gamma_{10} + \Gamma_{16} + \Gamma_{23} + \Gamma_{28} + \Gamma_{38} + \Gamma_{43} + 0.7212\Gamma_{150} + 0.1049\Gamma_{168}) / \Gamma$$

| VALUE (%) | EVTS | DOCUMENT ID | TECN | COMMENT |
|---|------|--|------|---------|
| 1.552 ± 0.029 OUR FIT | | | | |
| 1.53 ± 0.04 OUR AVERAGE | | | | |
| 1.528 ± 0.039 ± 0.040 | | ¹ ABBIENDI 01J OPAL 1990–1995 LEP runs | | |
| 1.54 ± 0.24 | | ABREU 94K DLPH LEP 1992 Z data | | |
| 1.70 ± 0.12 ± 0.19 | 202 | ² BATTLE 94 CLEO $E_{\text{cm}}^{\text{ee}} \approx 10.6 \text{ GeV}$ | | |
| • • • We use the following data for averages but not for fits. • • • | | | | |
| 1.520 ± 0.040 ± 0.041 | 4006 | ³ BARATE 99K ALEP 1991–1995 LEP runs | | |
| • • • We do not use the following data for averages, fits, limits, etc. • • • | | | | |
| 1.70 ± 0.05 ± 0.06 | 1610 | ⁴ BUSKULIC 96 ALEP Repl. by BARATE 99K | | |
| 1.6 ± 0.4 ± 0.2 | 35 | AIHARA 87B TPC $E_{\text{cm}}^{\text{ee}} = 29 \text{ GeV}$ | | |
| 1.71 ± 0.29 | 53 | MILLS 84 DLCO $E_{\text{cm}}^{\text{ee}} = 29 \text{ GeV}$ | | |

¹ The correlation coefficient between this measurement and the ABBIENDI 01J $B(\tau^- \rightarrow K^- \nu_\tau)$ is 0.60.

² BATTLE 94 quote $1.60 \pm 0.12 \pm 0.19$. We add 0.10 ± 0.02 to correct for their rejection of $K_S^0 \rightarrow \pi^+ \pi^-$ decays.

³ Not independent of BARATE 99K $B(K^- \nu_\tau)$, $B(K^- \pi^0 \nu_\tau)$, $B(K^- 2\pi^0 \nu_\tau$ (ex. K^0)), $B(K^- 3\pi^0 \nu_\tau$ (ex. K^0)), $B(K^- K^0 \nu_\tau)$, and $B(K^- K^0 \pi^0 \nu_\tau)$ values.

⁴ Not independent of BUSKULIC 96 $B(K^- \nu_\tau)$, $B(K^- \pi^0 \nu_\tau)$, $B(K^- 2\pi^0 \nu_\tau)$, $B(K^- K^0 \nu_\tau)$, and $B(K^- K^0 \pi^0 \nu_\tau)$ values.

$\Gamma(K^- \geq 1(\pi^0 \text{ or } K^0 \text{ or } \gamma) \nu_\tau)/\Gamma_{\text{total}}$ Γ_{33}/Γ

| $\Gamma_{33}/\Gamma = (\Gamma_{16} + \Gamma_{23} + \Gamma_{28} + \Gamma_{38} + \Gamma_{43} + 0.7212\Gamma_{150} + 0.7212\Gamma_{152} + 0.1049\Gamma_{168})/\Gamma$ | <u>EVTS</u> | <u>DOCUMENT ID</u> | <u>TECN</u> | <u>COMMENT</u> |
|--|-------------|--------------------|-------------|----------------|
|--|-------------|--------------------|-------------|----------------|

0.859 ± 0.028 OUR FIT

0.86 ± 0.05 OUR AVERAGE

• • • We use the following data for averages but not for fits. • • •

| | | | | |
|-----------------------------|-----------------------|-----|------|--------------------|
| $0.869 \pm 0.031 \pm 0.034$ | ¹ ABBIENDI | 01J | OPAL | 1990–1995 LEP runs |
| 0.69 ± 0.25 | ² ABREU | 94K | DLPH | LEP 1992 Z data |

• • • We do not use the following data for averages, fits, limits, etc. • • •

| | | | | | |
|-----------------------|---|--------|-----|-----|--|
| $1.2 \pm 0.5 \pm 0.2$ | 9 | AIHARA | 87B | TPC | $E_{\text{cm}}^{\text{ee}} = 29 \text{ GeV}$ |
|-----------------------|---|--------|-----|-----|--|

¹ Not independent of ABBIENDI 01J $B(\tau^- \rightarrow K^- \nu_\tau)$ and $B(\tau^- \rightarrow K^- \geq 0 \pi^0 \geq 0 K^0 \geq 0 \gamma \nu_\tau)$ values.

² Not independent of ABREU 94K $B(K^- \nu_\tau)$ and $B(K^- \geq 0 \text{ neutrals} \nu_\tau)$ measurements.

$\Gamma(K_S^0(\text{particles})^- \nu_\tau)/\Gamma_{\text{total}}$ Γ_{34}/Γ

| $\Gamma_{34}/\Gamma = (\frac{1}{2}\Gamma_{36} + \frac{1}{2}\Gamma_{38} + \frac{1}{2}\Gamma_{41} + \frac{1}{2}\Gamma_{43} + \frac{1}{2}\Gamma_{45} + \Gamma_{48} + \Gamma_{49} + \Gamma_{52} + \Gamma_{56} + 0.3606\Gamma_{154} + 0.342\Gamma_{168})/\Gamma$ |
|---|
|---|

| <u>VALUE (%)</u> | <u>EVTS</u> | <u>DOCUMENT ID</u> | <u>TECN</u> | <u>COMMENT</u> |
|------------------|-------------|--------------------|-------------|----------------|
|------------------|-------------|--------------------|-------------|----------------|

0.944 ± 0.028 OUR FIT

0.918 ± 0.015 OUR AVERAGE

| | | | | | |
|-----------------------------|-----|--------|-----|------|---|
| $0.970 \pm 0.058 \pm 0.062$ | 929 | BARATE | 98E | ALEP | 1991–1995 LEP runs |
| $0.97 \pm 0.09 \pm 0.06$ | 141 | AKERS | 94G | OPAL | $E_{\text{cm}}^{\text{ee}} = 88–94 \text{ GeV}$ |

• • • We use the following data for averages but not for fits. • • •

| | | | | | |
|-----------------------------|------|------------------|----|------|--|
| $0.915 \pm 0.001 \pm 0.015$ | 398k | ¹ RYU | 14 | BELL | $669 \text{ fb}^{-1} E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$ |
|-----------------------------|------|------------------|----|------|--|

¹ Not independent of RYU 14 measurements of $B(\tau^- \rightarrow \pi^- \bar{K}^0 \nu_\tau)$, $B(\tau^- \rightarrow K^- K^0 \nu_\tau)$, $B(\tau^- \rightarrow \pi^- \bar{K}^0 \pi^0 \nu_\tau)$, $B(\tau^- \rightarrow K^- K^0 \pi^0 \nu_\tau)$, $B(\tau^- \rightarrow \pi^- K_S^0 K_S^0 \nu_\tau)$, and $B(\tau^- \rightarrow \pi^- K_S^0 K_S^0 \pi^0 \nu_\tau)$.

$\Gamma(h^- \bar{K}^0 \nu_\tau)/\Gamma_{\text{total}}$ $\Gamma_{35}/\Gamma = (\Gamma_{36} + \Gamma_{38})/\Gamma$

| <u>VALUE (%)</u> | <u>EVTS</u> | <u>DOCUMENT ID</u> | <u>TECN</u> | <u>COMMENT</u> |
|------------------|-------------|--------------------|-------------|----------------|
|------------------|-------------|--------------------|-------------|----------------|

0.987 ± 0.014 OUR FIT

0.90 ± 0.07 OUR AVERAGE

| | | | | | |
|-----------------------------|------|------|----|------|--|
| $0.855 \pm 0.036 \pm 0.073$ | 1242 | COAN | 96 | CLEO | $E_{\text{cm}}^{\text{ee}} \approx 10.6 \text{ GeV}$ |
|-----------------------------|------|------|----|------|--|

• • • We use the following data for averages but not for fits. • • •

| | | | | | |
|--------------------------|-----|---------------------|-----|------|--------------------|
| $1.01 \pm 0.11 \pm 0.07$ | 555 | ¹ BARATE | 98E | ALEP | 1991–1995 LEP runs |
|--------------------------|-----|---------------------|-----|------|--------------------|

¹ Not independent of BARATE 98E $B(\tau^- \rightarrow \pi^- \bar{K}^0 \nu_\tau)$ and $B(\tau^- \rightarrow K^- K^0 \nu_\tau)$ values.

$\Gamma(\pi^-\bar{K}^0\nu_\tau)/\Gamma_{\text{total}}$ Γ_{36}/Γ VALUE (units 10^{-3})

DOCUMENT ID

TECN

COMMENT

8.40±0.14 OUR FIT**8.39±0.22 OUR AVERAGE**

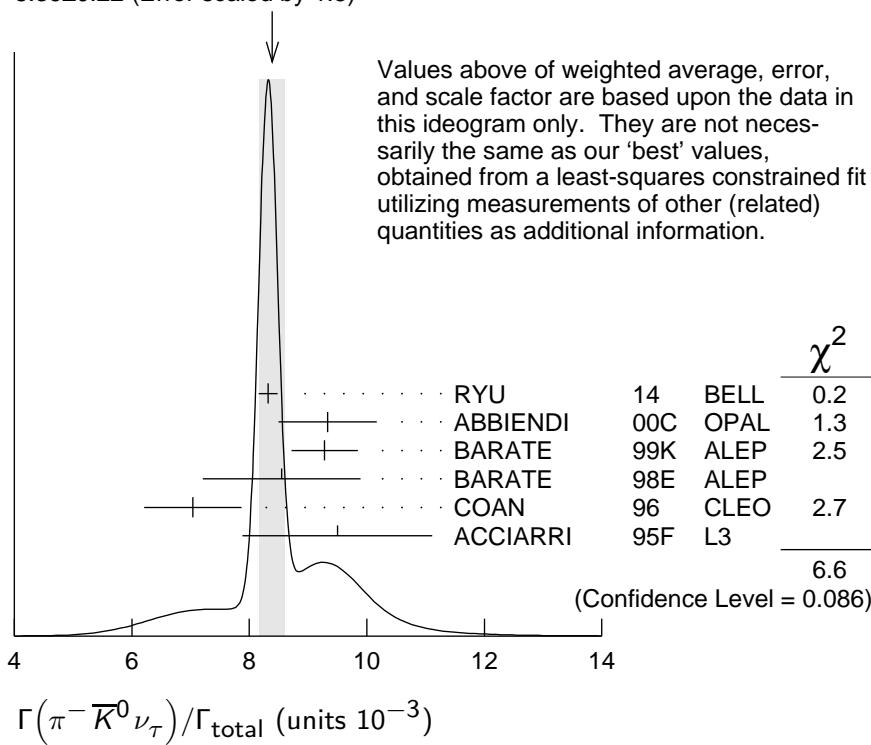
Error includes scale factor of 1.5. See the ideogram below.

| | | | | | | |
|--|------|-----------------------|-----|------|---|---|
| $8.32 \pm 0.02 \pm 0.16$ | 158k | ¹ RYU | 14 | BELL | 669 fb^{-1} | $E_{\text{cm}}^{ee} = 10.6 \text{ GeV}$ |
| $9.33 \pm 0.68 \pm 0.49$ | 377 | ABBIENDI | 00C | OPAL | 1991–1995 LEP runs | |
| $9.28 \pm 0.45 \pm 0.34$ | 937 | ² BARATE | 99K | ALEP | 1991–1995 LEP runs | |
| $9.5 \pm 1.5 \pm 0.6$ | | ³ ACCIARRI | 95F | L3 | 1991–1993 LEP runs | |
| • • • We use the following data for averages but not for fits. • • • | | | | | | |
| $8.55 \pm 1.17 \pm 0.66$ | 509 | ⁴ BARATE | 98E | ALEP | 1991–1995 LEP runs | |
| $7.04 \pm 0.41 \pm 0.72$ | | ⁵ COAN | 96 | CLEO | $E_{\text{cm}}^{ee} \approx 10.6 \text{ GeV}$ | |
| • • • We do not use the following data for averages, fits, limits, etc. • • • | | | | | | |
| $8.08 \pm 0.04 \pm 0.26$ | 53k | EPIFANOV | 07 | BELL | Repl. by RYU 14 | |
| $7.9 \pm 1.0 \pm 0.9$ | 98 | ⁶ BUSKULIC | 96 | ALEP | Repl. by BARATE 99K | |

¹ RYU 14 reconstruct K^0 's using $K_S^0 \rightarrow \pi^+ \pi^-$ decays.² BARATE 99K measure K^0 's by detecting K_L^0 's in their hadron calorimeter.³ ACCIARRI 95F do not identify π^-/K^- and assume $B(K^- K^0 \nu_\tau) = (0.29 \pm 0.12)\%$.⁴ BARATE 98E reconstruct K^0 's using $K_S^0 \rightarrow \pi^+ \pi^-$ decays. Not independent of BARATE 98E $B(K^0 \nu_\tau)$ value.⁵ Not independent of COAN 96 $B(h^- K^0 \nu_\tau)$ and $B(K^- K^0 \nu_\tau)$ measurements.⁶ BUSKULIC 96 measure K^0 's by detecting K_L^0 's in their hadron calorimeter.

WEIGHTED AVERAGE

8.39±0.22 (Error scaled by 1.5)



$\Gamma(\pi^-\bar{K}^0(\text{non-}K^*(892)^-)\nu_\tau)/\Gamma_{\text{total}}$ Γ_{37}/Γ

| <u>VALUE (units 10^{-4})</u> | <u>CL%</u> | <u>DOCUMENT ID</u> | <u>TECN</u> | <u>COMMENT</u> |
|---|------------|--------------------|-------------|--|
| 5.4±2.1 | | 1 EPIFANOV | 07 BELL | $351 \text{ fb}^{-1} E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$ |

• • • We do not use the following data for averages, fits, limits, etc. • • •

<17 95 ACCIARRI 95F L3 1991–1993 LEP runs

¹ EPIFANOV 07 quote $B(\tau^- \rightarrow K^*(892)^- \nu_\tau) B(K^*(892)^- \rightarrow K_S^0 \pi^-) / B(\tau^- \rightarrow K_S^0 \pi^- \nu_\tau) = 0.933 \pm 0.027$. We multiply their $B(\tau^- \rightarrow \bar{K}^0 \pi^- \nu_\tau)$ by $[1 - (0.933 \pm 0.027)]$ to obtain this result.

 $\Gamma(K^-K^0\nu_\tau)/\Gamma_{\text{total}}$ Γ_{38}/Γ

| <u>VALUE (units 10^{-4})</u> | <u>EVTS</u> | <u>DOCUMENT ID</u> | <u>TECN</u> | <u>COMMENT</u> |
|---|-------------|--------------------|-------------|----------------|
| 14.8 ± 0.5 OUR FIT | | | | |
| 14.9 ± 0.5 OUR AVERAGE | | | | |

| | | | | |
|------------------|-----|----------|----------|--|
| 14.80±0.14±0.54 | 33k | 1 RYU | 14 BELL | $669 \text{ fb}^{-1} E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$ |
| 16.2 ± 2.1 ± 1.1 | 150 | 2 BARATE | 99K ALEP | 1991–1995 LEP runs |
| 15.8 ± 4.2 ± 1.7 | 46 | 3 BARATE | 98E ALEP | 1991–1995 LEP runs |
| 15.1 ± 2.1 ± 2.2 | 111 | COAN | 96 CLEO | $E_{\text{cm}}^{\text{ee}} \approx 10.6 \text{ GeV}$ |

• • • We do not use the following data for averages, fits, limits, etc. • • •

26 ± 9 ± 2 13 ⁴ BUSKULIC 96 ALEP Repl. by BARATE 99K

¹ RYU 14 reconstruct K^0 's using $K_S^0 \rightarrow \pi^+ \pi^-$ decays.

² BARATE 99K measure K^0 's by detecting K_L^0 's in their hadron calorimeter.

³ BARATE 98E reconstruct K^0 's using $K_S^0 \rightarrow \pi^+ \pi^-$ decays.

⁴ BUSKULIC 96 measure K^0 's by detecting K_L^0 's in their hadron calorimeter.

 $\Gamma(K^-K^0 \geq 0 \pi^0 \nu_\tau)/\Gamma_{\text{total}}$ $\Gamma_{39}/\Gamma = (\Gamma_{38} + \Gamma_{43})/\Gamma$

| <u>VALUE (%)</u> | <u>EVTS</u> | <u>DOCUMENT ID</u> | <u>TECN</u> | <u>COMMENT</u> |
|----------------------------|-------------|--------------------|-------------|--------------------|
| 0.298±0.008 OUR FIT | | | | |
| 0.330±0.055±0.039 | 124 | ABBIENDI | 00C OPAL | 1991–1995 LEP runs |

 $\Gamma(h^-\bar{K}^0\pi^0\nu_\tau)/\Gamma_{\text{total}}$ $\Gamma_{40}/\Gamma = (\Gamma_{41} + \Gamma_{43})/\Gamma$

| <u>VALUE (%)</u> | <u>EVTS</u> | <u>DOCUMENT ID</u> | <u>TECN</u> | <u>COMMENT</u> |
|--------------------------------|-------------|--------------------|-------------|-------------------------------------|
| 0.532±0.013 OUR FIT | | | | |
| 0.50 ± 0.06 OUR AVERAGE | | | | Error includes scale factor of 1.2. |

0.562±0.050±0.048 264 COAN 96 CLEO $E_{\text{cm}}^{\text{ee}} \approx 10.6 \text{ GeV}$

• • • We use the following data for averages but not for fits. • • •

0.446±0.052±0.046 157 ¹ BARATE 98E ALEP 1991–1995 LEP runs

¹ Not independent of BARATE 98E $B(\tau^- \rightarrow \pi^- \bar{K}^0 \pi^0 \nu_\tau)$ and $B(\tau^- \rightarrow K^- K^0 \pi^0 \nu_\tau)$ values.

 $\Gamma(\pi^-\bar{K}^0\pi^0\nu_\tau)/\Gamma_{\text{total}}$ Γ_{41}/Γ

| <u>VALUE (%)</u> | <u>EVTS</u> | <u>DOCUMENT ID</u> | <u>TECN</u> | <u>COMMENT</u> |
|--------------------------------|-------------|--------------------|-------------|----------------|
| 0.382±0.013 OUR FIT | | | | |
| 0.383±0.014 OUR AVERAGE | | | | |

| | | | | |
|--------------------|-----|------------|----------|--|
| 0.386±0.004±0.014 | 27k | 1 RYU | 14 BELL | $669 \text{ fb}^{-1} E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$ |
| 0.347±0.053±0.037 | 299 | 2 BARATE | 99K ALEP | 1991–1995 LEP runs |
| 0.294±0.073±0.037 | 142 | 3 BARATE | 98E ALEP | 1991–1995 LEP runs |
| 0.41 ± 0.12 ± 0.03 | | 4 ACCIARRI | 95F L3 | 1991–1993 LEP runs |

• • • We use the following data for averages but not for fits. • • •

$0.417 \pm 0.058 \pm 0.044$ ⁵ COAN 96 CLEO $E_{\text{cm}}^{\text{ee}} \approx 10.6 \text{ GeV}$

• • • We do not use the following data for averages, fits, limits, etc. • • •

$0.32 \pm 0.11 \pm 0.05$ 23 ⁶ BUSKULIC 96 ALEP Repl. by BARATE 99K

¹ RYU 14 reconstruct K^0 's using $K_S^0 \rightarrow \pi^+ \pi^-$ decays.

² BARATE 99K measure K^0 's by detecting K_L^0 's in their hadron calorimeter.

³ BARATE 98E reconstruct K^0 's using $K_S^0 \rightarrow \pi^+ \pi^-$ decays.

⁴ ACCIARRI 95F do not identify π^- / K^- and assume $B(K^- K^0 \pi^0 \nu_\tau) = (0.05 \pm 0.05)\%$.

⁵ Not independent of COAN 96 $B(h^- K^0 \pi^0 \nu_\tau)$ and $B(K^- K^0 \pi^0 \nu_\tau)$ measurements.

⁶ BUSKULIC 96 measure K^0 's by detecting K_L^0 's in their hadron calorimeter.

$\Gamma(K^0 \rho^- \nu_\tau) / \Gamma_{\text{total}}$

Γ_{42} / Γ

| VALUE (%) | DOCUMENT ID | TECN | COMMENT |
|-----------|-------------|------|---------|
|-----------|-------------|------|---------|

0.22 ±0.05 OUR AVERAGE

$0.250 \pm 0.057 \pm 0.044$ ¹ BARATE 99K ALEP 1991–1995 LEP runs

$0.188 \pm 0.054 \pm 0.038$ ² BARATE 98E ALEP 1991–1995 LEP runs

¹ BARATE 99K measure K^0 's by detecting K_L^0 's in hadron calorimeter. They determine the $\bar{K}^0 \rho^-$ fraction in $\tau^- \rightarrow \pi^- \bar{K}^0 \pi^0 \nu_\tau$ decays to be $(0.72 \pm 0.12 \pm 0.10)$ and multiply their $B(\pi^- \bar{K}^0 \pi^0 \nu_\tau)$ measurement by this fraction to obtain the quoted result.

² BARATE 98E reconstruct K^0 's using $K_S^0 \rightarrow \pi^+ \pi^-$ decays. They determine the $\bar{K}^0 \rho^-$ fraction in $\tau^- \rightarrow \pi^- \bar{K}^0 \pi^0 \nu_\tau$ decays to be $(0.64 \pm 0.09 \pm 0.10)$ and multiply their $B(\pi^- \bar{K}^0 \pi^0 \nu_\tau)$ measurement by this fraction to obtain the quoted result.

$\Gamma(K^- K^0 \pi^0 \nu_\tau) / \Gamma_{\text{total}}$

Γ_{43} / Γ

| VALUE (units 10^{-4}) | EVTS | DOCUMENT ID | TECN | COMMENT |
|--------------------------|------|-------------|------|---------|
|--------------------------|------|-------------|------|---------|

15.0 ±0.7 OUR FIT

14.9 ±0.7 OUR AVERAGE

$14.96 \pm 0.20 \pm 0.74$ 8.3k ¹ RYU 14 BELL $669 \text{ fb}^{-1} E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$

$14.3 \pm 2.5 \pm 1.5$ 78 ² BARATE 99K ALEP 1991–1995 LEP runs

$15.2 \pm 7.6 \pm 2.1$ 15 ³ BARATE 98E ALEP 1991–1995 LEP runs

$14.5 \pm 3.6 \pm 2.0$ 32 COAN 96 CLEO $E_{\text{cm}}^{\text{ee}} \approx 10.6 \text{ GeV}$

• • • We do not use the following data for averages, fits, limits, etc. • • •

$10 \pm 5 \pm 3$ 5 ⁴ BUSKULIC 96 ALEP Repl. by BARATE 99K

¹ RYU 14 reconstruct K^0 's using $K_S^0 \rightarrow \pi^+ \pi^-$ decays.

² BARATE 99K measure K^0 's by detecting K_L^0 's in their hadron calorimeter.

³ BARATE 98E reconstruct K^0 's using $K_S^0 \rightarrow \pi^+ \pi^-$ decays.

⁴ BUSKULIC 96 measure K^0 's by detecting K_L^0 's in their hadron calorimeter.

$\Gamma(\pi^- \bar{K}^0 \geq 1 \pi^0 \nu_\tau) / \Gamma_{\text{total}}$

$\Gamma_{44} / \Gamma = (\Gamma_{41} + \Gamma_{45}) / \Gamma$

| VALUE (%) | EVTS | DOCUMENT ID | TECN | COMMENT |
|-----------|------|-------------|------|---------|
|-----------|------|-------------|------|---------|

0.408±0.025 OUR FIT

0.324±0.074±0.066 148 ABBIENDI 00C OPAL 1991–1995 LEP runs

$\Gamma(\pi^- \bar{K}^0 \pi^0 \nu_\tau (\text{ex. } K^0)) / \Gamma_{\text{total}}$

Γ_{45} / Γ

| VALUE (units 10^{-3}) | CL% | EVTS | DOCUMENT ID | TECN | COMMENT |
|--------------------------|-----|------|-------------|------|---------|
|--------------------------|-----|------|-------------|------|---------|

0.26±0.23 OUR FIT

0.26±0.24

¹ BARATE 99R ALEP 1991–1995 LEP runs

• • • We do not use the following data for averages, fits, limits, etc. • • •

| | | | | | | |
|----------------|----|----|---------------------|-----|------|--------------------|
| <0.66 | 95 | 17 | ² BARATE | 99K | ALEP | 1991–1995 LEP runs |
| 0.58±0.33±0.14 | | 5 | ³ BARATE | 98E | ALEP | 1991–1995 LEP runs |

¹ BARATE 99R combine the BARATE 98E and BARATE 99K measurements to obtain this value.

² BARATE 99K measure K^0 's by detecting K_L^0 's in their hadron calorimeter.

³ BARATE 98E reconstruct K^0 's using $K_S^0 \rightarrow \pi^+ \pi^-$ decays.

$\Gamma(K^- K^0 \pi^0 \pi^0 \nu_\tau)/\Gamma_{\text{total}}$

Γ_{46}/Γ

| VALUE | CL% | DOCUMENT ID | TECN | COMMENT |
|--|-----|---------------------|------|-------------------------|
| $<0.16 \times 10^{-3}$ | 95 | ¹ BARATE | 99R | ALEP 1991–1995 LEP runs |
| • • • We do not use the following data for averages, fits, limits, etc. • • • | | | | |
| $<0.18 \times 10^{-3}$ | 95 | ² BARATE | 99K | ALEP 1991–1995 LEP runs |
| $<0.39 \times 10^{-3}$ | 95 | ³ BARATE | 98E | ALEP 1991–1995 LEP runs |

¹ BARATE 99R combine the BARATE 98E and BARATE 99K bounds to obtain this value.

² BARATE 99K measure K^0 's by detecting K_L^0 's in hadron calorimeter.

³ BARATE 98E reconstruct K^0 's by using $K_S^0 \rightarrow \pi^+ \pi^-$ decays.

$\Gamma(\pi^- K^0 \bar{K}^0 \nu_\tau)/\Gamma_{\text{total}}$

$\Gamma_{47}/\Gamma = (\Gamma_{48} + \Gamma_{49} + \Gamma_{50})/\Gamma$

| VALUE (%) | EVTS | DOCUMENT ID | TECN | COMMENT |
|-----------|------|-------------|------|---------|
|-----------|------|-------------|------|---------|

0.155±0.024 OUR FIT

• • • We use the following data for averages but not for fits. • • •

| | | | | |
|--------------------------|----|---------------------|-----|-------------------------|
| 0.153±0.030±0.016 | 74 | ¹ BARATE | 98E | ALEP 1991–1995 LEP runs |
|--------------------------|----|---------------------|-----|-------------------------|

• • • We do not use the following data for averages, fits, limits, etc. • • •

| | | | | |
|--------------------|--|-----------------------|-----|-----------------------|
| 0.31 ± 0.12 ± 0.04 | | ² ACCIARRI | 95F | L3 1991–1993 LEP runs |
|--------------------|--|-----------------------|-----|-----------------------|

¹ BARATE 98E obtain this value by adding twice their $B(\pi^- K_S^0 K_S^0 \nu_\tau)$ value to their $B(\pi^- K_S^0 K_L^0 \nu_\tau)$ value.

² ACCIARRI 95F assume $B(\pi^- K_S^0 K_S^0 \nu) = B(\pi^- K_S^0 K_L^0 \nu) = 1/2B(\pi^- K_S^0 K_L^0 \nu)$.

$\Gamma(\pi^- K_S^0 K_S^0 \nu_\tau)/\Gamma_{\text{total}}$

Γ_{48}/Γ

Bose-Einstein correlations might make the mixing fraction different than 1/4.

| VALUE (units 10^{-4}) | EVTS | DOCUMENT ID | TECN | COMMENT |
|--------------------------|------|-------------|------|---------|
|--------------------------|------|-------------|------|---------|

2.33±0.07 OUR FIT

2.32±0.06 OUR AVERAGE

| | | | | |
|-----------------|------|--------|-----|---|
| 2.33±0.03±0.09 | 6.7k | RYU | 14 | BELL $669 \text{ fb}^{-1} E_{\text{cm}}^{\text{ee}}=10.6 \text{ GeV}$ |
| 2.31±0.04±0.08 | 5.0k | LEES | 12Y | BABR $468 \text{ fb}^{-1} E_{\text{cm}}^{\text{ee}}=10.6 \text{ GeV}$ |
| 2.6 ± 1.0 ± 0.5 | 6 | BARATE | 98E | ALEP 1991–1995 LEP runs |
| 2.3 ± 0.5 ± 0.3 | 42 | COAN | 96 | CLEO $E_{\text{cm}}^{\text{ee}} \approx 10.6 \text{ GeV}$ |

$\Gamma(\pi^- K_S^0 K_L^0 \nu_\tau)/\Gamma_{\text{total}}$

Γ_{49}/Γ

| VALUE (units 10^{-4}) | EVTS | DOCUMENT ID | TECN | COMMENT |
|--------------------------|------|-------------|------|---------|
|--------------------------|------|-------------|------|---------|

10.8±2.4 OUR FIT

| | | | | |
|---------------------|----|--------|-----|-------------------------|
| 10.1±2.3±1.3 | 68 | BARATE | 98E | ALEP 1991–1995 LEP runs |
|---------------------|----|--------|-----|-------------------------|

$\Gamma(\pi^- K_L^0 K_L^0 \nu_\tau)/\Gamma_{\text{total}}$

$\Gamma_{50}/\Gamma = \Gamma_{48}/\Gamma$

| VALUE (units 10^{-4}) | DOCUMENT ID |
|--------------------------|-------------|
|--------------------------|-------------|

2.33±0.07 OUR FIT

$\Gamma(\pi^- K^0 \bar{K}^0 \pi^0 \nu_\tau)/\Gamma_{\text{total}}$ VALUE (units 10^{-4}) **3.6 ± 1.2 OUR FIT**

• • • We use the following data for averages but not for fits. • • •

 3.1 ± 2.3 DOCUMENT ID $\Gamma_{51}/\Gamma = (\Gamma_{52} + \Gamma_{56} + \Gamma_{57})/\Gamma$ TECNCOMMENT

¹ BARATE 99R ALEP 1991–1995 LEP runs
¹ BARATE 99R combine BARATE 98E $\Gamma(\pi^- K_S^0 \bar{K}_S^0 \pi^0 \nu_\tau)/\Gamma_{\text{total}}$ and
 $\Gamma(\pi^- K_S^0 \bar{K}_L^0 \pi^0 \nu_\tau)/\Gamma_{\text{total}}$ measurements to obtain this value.

 $\Gamma(\pi^- K_S^0 \bar{K}_S^0 \pi^0 \nu_\tau)/\Gamma_{\text{total}}$ VALUE (units 10^{-5}) **1.82 ± 0.21 OUR FIT** **1.80 ± 0.21 OUR AVERAGE**

| | | | | | | |
|----------------------------|-----|------|-----|------|-----------------------|--|
| 2.00 \pm 0.22 \pm 0.20 | 303 | RYU | 14 | BELL | 669 fb^{-1} | $E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$ |
| 1.60 \pm 0.20 \pm 0.22 | 409 | LEES | 12Y | BABR | 468 fb^{-1} | $E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$ |

• • • We do not use the following data for averages, fits, limits, etc. • • •

<20 95 BARATE 98E ALEP 1991–1995 LEP runs

 Γ_{52}/Γ DOCUMENT IDTECNCOMMENT $\Gamma(K^* - K^0 \pi^0 \nu_\tau \rightarrow \pi^- K_S^0 \bar{K}_S^0 \pi^0 \nu_\tau)/\Gamma_{\text{total}}$ Γ_{53}/Γ VALUE (units 10^{-6}) **$10.8 \pm 1.4 \pm 1.5$** DOCUMENT IDTECNCOMMENTRYU 14 BELL 669 fb^{-1} $E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$ $\Gamma(f_1(1285) \pi^- \nu_\tau \rightarrow \pi^- K_S^0 \bar{K}_S^0 \pi^0 \nu_\tau)/\Gamma_{\text{total}}$ Γ_{54}/Γ VALUE (units 10^{-6}) **$6.8 \pm 1.3 \pm 0.7$** DOCUMENT IDTECNCOMMENTRYU 14 BELL 669 fb^{-1} $E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$ $\Gamma(f_1(1420) \pi^- \nu_\tau \rightarrow \pi^- K_S^0 \bar{K}_S^0 \pi^0 \nu_\tau)/\Gamma_{\text{total}}$ Γ_{55}/Γ VALUE (units 10^{-6}) **$2.4 \pm 0.5 \pm 0.6$** DOCUMENT IDTECNCOMMENTRYU 14 BELL 669 fb^{-1} $E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$ $\Gamma(\pi^- K_S^0 \bar{K}_L^0 \pi^0 \nu_\tau)/\Gamma_{\text{total}}$ Γ_{56}/Γ VALUE (units 10^{-4}) **3.2 ± 1.2 OUR FIT** **$3.1 \pm 1.1 \pm 0.5$** DOCUMENT IDTECNCOMMENT

BARATE 98E ALEP 1991–1995 LEP runs

 $\Gamma(\pi^- K_L^0 \bar{K}_L^0 \pi^0 \nu_\tau)/\Gamma_{\text{total}}$ $\Gamma_{57}/\Gamma = \Gamma_{52}/\Gamma$ VALUE (units 10^{-5}) **1.82 ± 0.21 OUR FIT**DOCUMENT ID $\Gamma(K^- K_S^0 \bar{K}_S^0 \nu_\tau)/\Gamma_{\text{total}}$ Γ_{58}/Γ VALUECL% **$<6.3 \times 10^{-7}$**

90

DOCUMENT IDTECNCOMMENTLEES 12Y BABR 468 fb^{-1} $E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$ $\Gamma(K^- K_S^0 \bar{K}_S^0 \pi^0 \nu_\tau)/\Gamma_{\text{total}}$ Γ_{59}/Γ VALUECL% **$<4.0 \times 10^{-7}$**

90

DOCUMENT IDTECNCOMMENTLEES 12Y BABR 468 fb^{-1} $E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$

$\Gamma(K^0 h^+ h^- h^- \geq 0 \text{ neutrals } \nu_\tau)/\Gamma_{\text{total}}$ Γ_{60}/Γ

| VALUE (%) | CL % | DOCUMENT ID | TECN | COMMENT |
|---|------|---------------|------|--|
| <0.17 | 95 | TSCHIRHART 88 | HRS | $E_{\text{cm}}^{\text{ee}} = 29 \text{ GeV}$ |
| $\bullet \bullet \bullet$ We do not use the following data for averages, fits, limits, etc. $\bullet \bullet \bullet$ | | | | |
| <0.27 | 90 | BELTRAMI | 85 | HRS $E_{\text{cm}}^{\text{ee}} = 29 \text{ GeV}$ |

 $\Gamma(K^0 h^+ h^- h^- \nu_\tau)/\Gamma_{\text{total}}$ Γ_{61}/Γ

| VALUE (units 10^{-4}) | EVTS | DOCUMENT ID | TECN | COMMENT |
|--|------|---------------------|------|-------------------------|
| 2.5±2.0 OUR FIT | | | | |
| 2.3±1.9±0.7 | 6 | ¹ BARATE | 98E | ALEP 1991–1995 LEP runs |
| ¹ BARATE 98E reconstruct K^0 's using $K_S^0 \rightarrow \pi^+ \pi^-$ decays. | | | | |

 $\Gamma(h^- h^- h^+ \geq 0 \text{ neutrals} \geq 0 K_L^0 \nu_\tau)/\Gamma_{\text{total}}$ Γ_{62}/Γ

$$\begin{aligned} \Gamma_{62}/\Gamma = & (0.34598\Gamma_{36} + 0.34598\Gamma_{38} + 0.34598\Gamma_{41} + 0.34598\Gamma_{43} + 0.4247\Gamma_{48} + \\ & 0.6920\Gamma_{49} + 0.4247\Gamma_{52} + 0.6920\Gamma_{56} + 0.6534\Gamma_{61} + \Gamma_{70} + \Gamma_{78} + \Gamma_{85} + \Gamma_{86} + \Gamma_{97} + \Gamma_{103} + \\ & \Gamma_{106} + \Gamma_{107} + 0.2810\Gamma_{148} + 0.2810\Gamma_{150} + 0.2810\Gamma_{152} + 0.2628\Gamma_{154} + 0.7259\Gamma_{168} + \\ & 0.9078\Gamma_{176} + 0.9078\Gamma_{177} + 0.9078\Gamma_{178})/\Gamma \end{aligned}$$

| VALUE (%) | EVTS | DOCUMENT ID | TECN | COMMENT |
|---|------|------------------------|------|--|
| 15.21± 0.06 OUR FIT | | | | |
| 14.8 ± 0.4 OUR AVERAGE | | | | |
| 14.4 ± 0.6 ± 0.3 | | ADEVA | 91F | L3 $E_{\text{cm}}^{\text{ee}} = 88.3\text{--}94.3 \text{ GeV}$ |
| 15.0 ± 0.4 ± 0.3 | | BEHREND | 89B | CELL $E_{\text{cm}}^{\text{ee}} = 14\text{--}47 \text{ GeV}$ |
| 15.1 ± 0.8 ± 0.6 | | AIHARA | 87B | TPC $E_{\text{cm}}^{\text{ee}} = 29 \text{ GeV}$ |
| $\bullet \bullet \bullet$ We do not use the following data for averages, fits, limits, etc. $\bullet \bullet \bullet$ | | | | |
| 13.5 ± 0.3 ± 0.3 | | ABACHI | 89B | HRS $E_{\text{cm}}^{\text{ee}} = 29 \text{ GeV}$ |
| 12.8 ± 1.0 ± 0.7 | | ¹ BURCHAT | 87 | MRK2 $E_{\text{cm}}^{\text{ee}} = 29 \text{ GeV}$ |
| 12.1 ± 0.5 ± 1.2 | | RUCKSTUHL | 86 | DLCO $E_{\text{cm}}^{\text{ee}} = 29 \text{ GeV}$ |
| 12.8 ± 0.5 ± 0.8 | 1420 | SCHMIDKE | 86 | MRK2 $E_{\text{cm}}^{\text{ee}} = 29 \text{ GeV}$ |
| 15.3 ± 1.1 ^{+1.3} _{-1.6} | 367 | ALTHOFF | 85 | TASS $E_{\text{cm}}^{\text{ee}} = 34.5 \text{ GeV}$ |
| 13.6 ± 0.5 ± 0.8 | | BARTEL | 85F | JADE $E_{\text{cm}}^{\text{ee}} = 34.6 \text{ GeV}$ |
| 12.2 ± 1.3 ± 3.9 | | ² BERGER | 85 | PLUT $E_{\text{cm}}^{\text{ee}} = 34.6 \text{ GeV}$ |
| 13.3 ± 0.3 ± 0.6 | | FERNANDEZ | 85 | MAC $E_{\text{cm}}^{\text{ee}} = 29 \text{ GeV}$ |
| 24 ± 6 | 35 | BRANDELIK | 80 | TASS $E_{\text{cm}}^{\text{ee}} = 30 \text{ GeV}$ |
| 32 ± 5 | 692 | ³ BACINO | 78B | DLCO $E_{\text{cm}}^{\text{ee}} = 3.1\text{--}7.4 \text{ GeV}$ |
| 35 ± 11 | | ³ BRANDELIK | 78 | DASP Assumes $V\text{--}A$ decay |
| 18 ± 6.5 | 33 | ³ JAROS | 78 | LGW $E_{\text{cm}}^{\text{ee}} > 6 \text{ GeV}$ |

¹ BURCHAT 87 value is not independent of SCHMIDKE 86 value.² Not independent of BERGER 85 $\Gamma(\mu^- \bar{\nu}_\mu \nu_\tau)/\Gamma_{\text{total}}$, $\Gamma(e^- \bar{\nu}_e \nu_\tau)/\Gamma_{\text{total}}$, $\Gamma(h^- \geq 1 \text{ neutrals} \nu_\tau)/\Gamma_{\text{total}}$, and $\Gamma(h^- \geq 0 K_L^0 \nu_\tau)/\Gamma_{\text{total}}$, and therefore not used in the fit.³ Low energy experiments are not in average or fit because the systematic errors in background subtraction are judged to be large.

$\Gamma(h^- h^- h^+ \geq 0 \text{ neutrals } \nu_\tau \text{ (ex. } K_S^0 \rightarrow \pi^+ \pi^- \text{)} / \Gamma_{\text{total}} \quad \Gamma_{63} / \Gamma$

$$\Gamma_{63} / \Gamma = (\Gamma_{70} + \Gamma_{78} + \Gamma_{85} + \Gamma_{86} + \Gamma_{97} + \Gamma_{103} + \Gamma_{106} + \Gamma_{107} + 0.2810 \Gamma_{148} + 0.2810 \Gamma_{150} + 0.2810 \Gamma_{152} + 0.489 \Gamma_{168} + 0.9078 \Gamma_{176} + 0.9078 \Gamma_{177} + 0.9078 \Gamma_{178}) / \Gamma$$

| VALUE (%) | EVTS | DOCUMENT ID | TECN | COMMENT |
|---|-------|----------------|------|---|
| 14.55 ± 0.06 OUR FIT | | | | |
| 14.61 ± 0.06 OUR AVERAGE | | | | |
| • • • We use the following data for averages but not for fits. • • • | | | | |
| 14.556 ± 0.105 ± 0.076 | 1 | ACHARD 01D | L3 | 1992–1995 LEP runs |
| 14.96 ± 0.09 ± 0.22 | 10.4k | AKERS 95Y | OPAL | 1991–1994 LEP runs |
| • • • We do not use the following data for averages, fits, limits, etc. • • • | | | | |
| 14.652 ± 0.067 ± 0.086 | | SCHAEL 05C | ALEP | 1991–1995 LEP runs |
| 14.569 ± 0.093 ± 0.048 | 23k | 2 ABREU 01M | DLPH | 1992–1995 LEP runs |
| 14.22 ± 0.10 ± 0.37 | | 3 BALEST 95C | CLEO | $E_{\text{cm}}^{ee} \approx 10.6 \text{ GeV}$ |
| • • • We do not use the following data for averages, fits, limits, etc. • • • | | | | |
| 15.26 ± 0.26 ± 0.22 | | ACTON 92H | OPAL | Repl. by AKERS 95Y |
| 13.3 ± 0.3 ± 0.8 | | 4 ALBRECHT 92D | ARG | $E_{\text{cm}}^{ee} = 9.4\text{--}10.6 \text{ GeV}$ |
| 14.35 ± 0.40 ± 0.45 | | DECAMP 92C | ALEP | 1989–1990 LEP runs |

¹ The correlation coefficients between this measurement and the ACHARD 01D measurements of $B(\tau \rightarrow \text{"1-prong"})$ and $B(\tau \rightarrow \text{"5-prong"})$ are -0.978 and -0.19 respectively.

² The correlation coefficients between this measurement and the ABREU 01M measurements of $B(\tau \rightarrow \text{1-prong})$ and $B(\tau \rightarrow \text{5-prong})$ are -0.98 and -0.08 respectively.

³ Not independent of BALEST 95C $B(h^- h^- h^+ \nu_\tau)$ and $B(h^- h^- h^+ \pi^0 \nu_\tau)$ values, and BORTOLETTO 93 $B(h^- h^- h^+ 2\pi^0 \nu_\tau) / B(h^- h^- h^+ \geq 0 \text{ neutrals } \nu_\tau)$ value.

⁴ This ALBRECHT 92D value is not independent of their $\Gamma(\mu^- \bar{\nu}_\mu \nu_\tau) \Gamma(e^- \bar{\nu}_e \nu_\tau) / \Gamma_{\text{total}}^2$ value.

 $\Gamma(h^- h^- h^+ \nu_\tau) / \Gamma_{\text{total}} \quad \Gamma_{64} / \Gamma$

$$\Gamma_{64} / \Gamma = (0.34598 \Gamma_{36} + 0.34598 \Gamma_{38} + \Gamma_{70} + \Gamma_{97} + \Gamma_{106} + 0.489 \Gamma_{168} + 0.0153 \Gamma_{176} + 0.0153 \Gamma_{177}) / \Gamma$$

| VALUE (%) | EVTS | DOCUMENT ID | TECN | COMMENT |
|---|-------|----------------|------|---|
| 9.80 ± 0.05 OUR FIT | | | | |
| • • • We use the following data for averages but not for fits. • • • | | | | |
| 7.6 ± 0.1 ± 0.5 | 7.5k | 1 ALBRECHT 96E | ARG | $E_{\text{cm}}^{ee} = 9.4\text{--}10.6 \text{ GeV}$ |
| • • • We do not use the following data for averages, fits, limits, etc. • • • | | | | |
| 9.92 ± 0.10 ± 0.09 | 11.2k | 2 BUSKULIC 96 | ALEP | Repl. by SCHAEL 05C |
| 9.49 ± 0.36 ± 0.63 | | DECAMP 92C | ALEP | Repl. by SCHAEL 05C |
| 8.7 ± 0.7 ± 0.3 | 694 | 3 BEHREND 90 | CELL | $E_{\text{cm}}^{ee} = 35 \text{ GeV}$ |
| 7.0 ± 0.3 ± 0.7 | 1566 | 4 BAND 87 | MAC | $E_{\text{cm}}^{ee} = 29 \text{ GeV}$ |
| 6.7 ± 0.8 ± 0.9 | | 5 BURCHAT 87 | MRK2 | $E_{\text{cm}}^{ee} = 29 \text{ GeV}$ |
| 6.4 ± 0.4 ± 0.9 | | 6 RUCKSTUHL 86 | DLCO | $E_{\text{cm}}^{ee} = 29 \text{ GeV}$ |
| 7.8 ± 0.5 ± 0.8 | 890 | SCHMIDKE 86 | MRK2 | $E_{\text{cm}}^{ee} = 29 \text{ GeV}$ |
| 8.4 ± 0.4 ± 0.7 | 1255 | 6 FERNANDEZ 85 | MAC | $E_{\text{cm}}^{ee} = 29 \text{ GeV}$ |
| 9.7 ± 2.0 ± 1.3 | | BEHREND 84 | CELL | $E_{\text{cm}}^{ee} = 14,22 \text{ GeV}$ |

¹ ALBRECHT 96E not independent of ALBRECHT 93C $\Gamma(h^- h^- h^+ \nu_\tau \text{ (ex. } K^0 \text{)} \times \Gamma(\text{particle}^- \geq 0 \text{ neutrals } \geq 0 K_L^0 \nu_\tau) / \Gamma_{\text{total}}^2$ value.

² BUSKULIC 96 quote $B(h^- h^- h^+ \nu_\tau \text{ (ex. } K^0 \text{)}) = 9.50 \pm 0.10 \pm 0.11$. We add 0.42 to remove their K^0 correction and reduce the systematic error accordingly.

³ BEHREND 90 subtract 0.3% to account for the $\tau^- \rightarrow K^*(892)^- \nu_\tau$ contribution to measured events.

⁴ BAND 87 subtract for charged kaon modes; not independent of FERNANDEZ 85 value.

⁵ BURCHAT 87 value is not independent of SCHMIDKE 86 value.

⁶ Value obtained by multiplying paper's $R = B(h^- h^- h^+ \nu_\tau)/B(3\text{-prong})$ by $B(3\text{-prong}) = 0.143$ and subtracting 0.3% for $K^*(892)$ background.

$\Gamma(h^- h^- h^+ \nu_\tau (\text{ex. } K^0))/\Gamma_{\text{total}}$

Γ_{65}/Γ

$$\Gamma_{65}/\Gamma = (\Gamma_{70} + \Gamma_{97} + \Gamma_{106} + 0.489\Gamma_{168} + 0.0153\Gamma_{176} + 0.0153\Gamma_{177})/\Gamma$$

| VALUE (%) | EVTS | DOCUMENT ID | TECN | COMMENT |
|-----------|------|-------------|------|---------|
|-----------|------|-------------|------|---------|

9.46 ± 0.05 OUR FIT

9.44 ± 0.14 OUR AVERAGE Error includes scale factor of 1.4. See the ideogram below.

$9.317 \pm 0.090 \pm 0.082$ 12.2k ¹ ABDALLAH 06A DLPH 1992–1995 LEP runs

$9.51 \pm 0.07 \pm 0.20$ 37.7k BALEST 95C CLEO $E_{\text{cm}}^{\text{ee}} \approx 10.6 \text{ GeV}$

• • • We use the following data for averages but not for fits. • • •

$9.87 \pm 0.10 \pm 0.24$ ² AKERS 95Y OPAL 1991–1994 LEP runs

• • • We do not use the following data for averages, fits, limits, etc. • • •

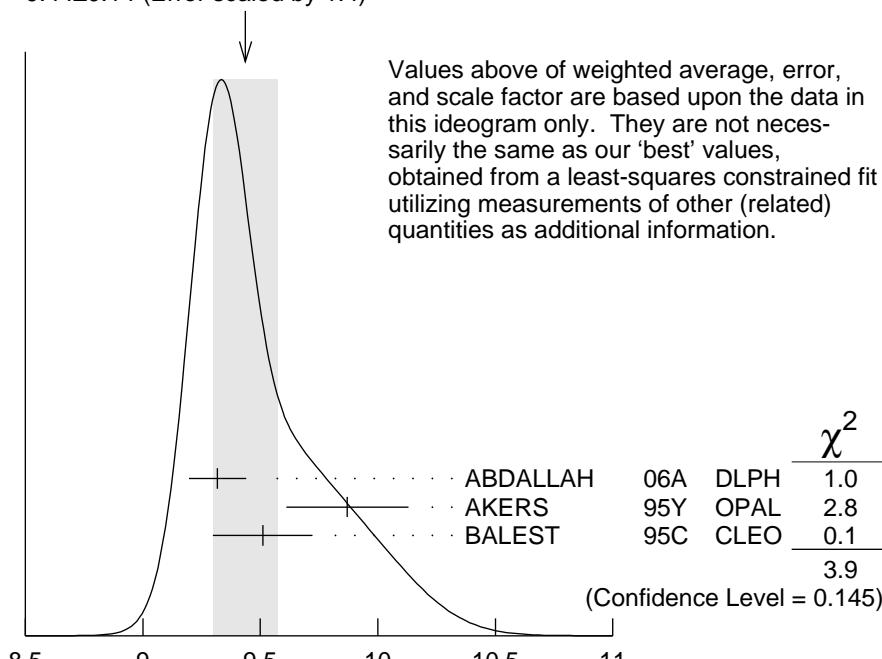
$9.50 \pm 0.10 \pm 0.11$ 11.2k ³ BUSKULIC 96 ALEP Repl. by SCHAEEL 05C

¹ See footnote to ABDALLAH 06A $\Gamma(\tau^- \rightarrow h^- \nu_\tau)/\Gamma_{\text{total}}$ measurement for correlations with other measurements.

² Not independent of AKERS 95Y $B(h^- h^- h^+ \geq 0 \text{ neutrals} \nu_\tau (\text{ex. } K_S^0 \rightarrow \pi^+ \pi^-))$ and $B(h^- h^- h^+ \nu_\tau (\text{ex. } K^0))/B(h^- h^- h^+ \geq 0 \text{ neutrals} \nu_\tau (\text{ex. } K_S^0 \rightarrow \pi^+ \pi^-))$ values.

³ Not independent of BUSKULIC 96 $B(h^- h^- h^+ \nu_\tau)$ value.

WEIGHTED AVERAGE
9.44±0.14 (Error scaled by 1.4)



$\Gamma(h^- h^- h^+ \nu_\tau (\text{ex. } K^0))/\Gamma_{\text{total}} (\%)$

$$\Gamma(h^- h^- h^+ \nu_\tau(\text{ex. } K^0)) / \Gamma(h^- h^- h^+ \geq 0 \text{ neutrals } \nu_\tau(\text{ex. } K_S^0 \rightarrow \pi^+ \pi^-)) \quad \Gamma_{65}/\Gamma_{63}$$

$\Gamma_{65}/\Gamma_{63} = (\Gamma_{70} + \Gamma_{97} + \Gamma_{106} + 0.489\Gamma_{168} + 0.0153\Gamma_{176} + 0.0153\Gamma_{177}) / (0.4247\Gamma_{52} +$
 $\Gamma_{70} + \Gamma_{78} + \Gamma_{85} + \Gamma_{89} + \Gamma_{97} + \Gamma_{103} + \Gamma_{106} + \Gamma_{107} + 0.2810\Gamma_{148} + 0.2292\Gamma_{149} +$
 $0.2810\Gamma_{150} + 0.2810\Gamma_{152} + 0.1131\Gamma_{154} + 0.3268\Gamma_{158} + 0.489\Gamma_{168} + 0.9078\Gamma_{176} +$
 $0.9078\Gamma_{177} + 0.9078\Gamma_{178} + 0.892\Gamma_{180})$

| VALUE (units 10^{-2}) | DOCUMENT ID | TECN | COMMENT |
|--|-------------|----------|--------------------|
| 64.98 ± 0.31 OUR FIT | | | |
| $66.0 \pm 0.4 \pm 1.4$ | AKERS | 95Y OPAL | 1991–1994 LEP runs |

$$\Gamma(h^- h^- h^+ \nu_\tau(\text{ex. } K^0, \omega)) / \Gamma_{\text{total}} \quad \Gamma_{66}/\Gamma$$

$$\Gamma_{66}/\Gamma = (\Gamma_{70} + \Gamma_{97} + \Gamma_{106} + 0.489\Gamma_{168}) / \Gamma$$

| VALUE (%) | DOCUMENT ID |
|---|-------------|
| 9.43 ± 0.05 OUR FIT | |

$$\Gamma(\pi^- \pi^+ \pi^- \nu_\tau) / \Gamma_{\text{total}} \quad \Gamma_{67}/\Gamma = (0.34598\Gamma_{36} + \Gamma_{70} + 0.0153\Gamma_{176}) / \Gamma$$

$$\Gamma(\pi^- \pi^+ \pi^- \nu_\tau) / \Gamma_{\text{total}} \quad \Gamma_{67}/\Gamma = (0.34598\Gamma_{36} + \Gamma_{70} + 0.0153\Gamma_{176}) / \Gamma$$

| VALUE (%) | EVTS | DOCUMENT ID | TECN | COMMENT |
|---|------|-------------|------|-------------------------------------|
| 9.02 ± 0.05 OUR FIT | | | | |
| 8.77 ± 0.13 OUR AVERAGE | | | | Error includes scale factor of 1.1. |

| | | | | | | | |
|-----------------|--------------------|------|---------------------|----|------|--|--|
| 8.42 ± 0.00 | $^{+0.26}_{-0.25}$ | 8.9M | ¹ LEE | 10 | BELL | 666 fb^{-1} | $E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$ |
| 8.83 ± 0.01 | ± 0.13 | 1.6M | ² AUBERT | 08 | BABR | 342 fb^{-1} | $E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$ |
| 9.13 ± 0.05 | ± 0.46 | 43k | ³ BRIERE | 03 | CLE3 | $E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$ | |

¹ Quoted statistical error is 0.003%. Correlation matrix for LEE 10 branching fractions:

- (1) $\Gamma(\tau^- \rightarrow \pi^- \pi^+ \pi^- \nu_\tau(\text{ex. } K^0)) / \Gamma_{\text{total}}$
- (2) $\Gamma(\tau^- \rightarrow K^- \pi^+ \pi^- \nu_\tau(\text{ex. } K^0)) / \Gamma_{\text{total}}$
- (3) $\Gamma(\tau^- \rightarrow K^- K^+ \pi^- \nu_\tau) / \Gamma_{\text{total}}$
- (4) $\Gamma(\tau^- \rightarrow K^- K^+ K^- \nu_\tau) / \Gamma_{\text{total}}$

| | | | |
|-----|--------|-------|--------|
| | (1) | (2) | (3) |
| (2) | 0.175 | | |
| (3) | 0.049 | 0.080 | |
| (4) | -0.053 | 0.035 | -0.008 |

² Correlation matrix for AUBERT 08 branching fractions:

- (1) $\Gamma(\tau^- \rightarrow \pi^- \pi^+ \pi^- \nu_\tau(\text{ex. } K^0)) / \Gamma_{\text{total}}$
- (2) $\Gamma(\tau^- \rightarrow K^- \pi^+ \pi^- \nu_\tau(\text{ex. } K^0)) / \Gamma_{\text{total}}$
- (3) $\Gamma(\tau^- \rightarrow K^- K^+ \pi^- \nu_\tau) / \Gamma_{\text{total}}$
- (4) $\Gamma(\tau^- \rightarrow K^- K^+ K^- \nu_\tau) / \Gamma_{\text{total}}$

| | | | |
|-----|-------|-------|-------|
| | (1) | (2) | (3) |
| (2) | 0.544 | | |
| (3) | 0.390 | 0.177 | |
| (4) | 0.031 | 0.093 | 0.087 |

³ 47% correlated with BRIERE 03 $\tau^- \rightarrow K^- \pi^+ \pi^- \nu_\tau$ and 71% correlated with $\tau^- \rightarrow K^- K^+ \pi^- \nu_\tau$ because of a common 5% normalization error.

$$\Gamma(\pi^-\pi^+\pi^-\nu_\tau(\text{ex. } K^0), \text{non-axial vector})/\Gamma(\pi^-\pi^+\pi^-\nu_\tau(\text{ex. } K^0)) \quad \Gamma_{69}/\Gamma_{68}$$

$$\Gamma_{69}/\Gamma_{68} = \Gamma_{69}/(\Gamma_{70} + 0.0153\Gamma_{175})$$

| VALUE | CL% | DOCUMENT ID | TECN | COMMENT |
|--------|-----|------------------|------|--------------------|
| <0.261 | 95 | 1 ACKERSTAFF 97R | OPAL | 1992–1994 LEP runs |

¹ Model-independent limit from structure function analysis on contribution to $B(\tau^- \rightarrow \pi^-\pi^+\pi^-\nu_\tau(\text{ex. } K^0))$ from non-axial vectors.

$$\Gamma(\pi^-\pi^+\pi^-\nu_\tau(\text{ex. } K^0, \omega))/\Gamma_{\text{total}} \quad \Gamma_{70}/\Gamma$$

| VALUE (%) | EVTS | DOCUMENT ID | TECN | COMMENT |
|---------------------------|------|-------------|----------|--------------------|
| 8.99 ±0.05 OUR FIT | | | | |
| 9.041±0.060±0.076 | 29k | 1 SCHAEL | 05C ALEP | 1991–1995 LEP runs |

¹ See footnote to SCHAEL 05C $\Gamma(\tau^- \rightarrow e^-\bar{\nu}_e\nu_\tau)/\Gamma_{\text{total}}$ measurement for correlations with other measurements.

$$\Gamma(h^- h^- h^+ \geq 1 \text{ neutrals } \nu_\tau)/\Gamma_{\text{total}} \quad \Gamma_{71}/\Gamma$$

$$\begin{aligned} \Gamma_{71}/\Gamma = & (0.34598\Gamma_{41} + 0.34598\Gamma_{43} + 0.4247\Gamma_{48} + 0.4247\Gamma_{52} + \Gamma_{78} + \Gamma_{85} + \Gamma_{86} + \\ & \Gamma_{103} + \Gamma_{107} + 0.2810\Gamma_{148} + 0.2810\Gamma_{150} + 0.2810\Gamma_{152} + 0.2926\Gamma_{154} + 0.892\Gamma_{176} + \\ & 0.892\Gamma_{177} + 0.9078\Gamma_{178})/\Gamma \end{aligned}$$

| VALUE (%) | EVTS | DOCUMENT ID | TECN | COMMENT |
|--------------------------|------|-------------|------|---------|
| 5.29±0.05 OUR FIT | | | | |

• • • We do not use the following data for averages, fits, limits, etc. • • •

| | | | | |
|-----------------|-----|---------------|---------|---|
| 5.6 ± 0.7 ± 0.3 | 352 | 1 BEHREND | 90 CELL | $E_{\text{cm}}^{\text{ee}} = 35 \text{ GeV}$ |
| 4.2 ± 0.5 ± 0.9 | 203 | 2 ALBRECHT | 87L ARG | $E_{\text{cm}}^{\text{ee}} = 10 \text{ GeV}$ |
| 6.1 ± 0.8 ± 0.9 | | 3 BURCHAT | 87 MRK2 | $E_{\text{cm}}^{\text{ee}} = 29 \text{ GeV}$ |
| 7.6 ± 0.4 ± 0.9 | | 4,5 RUCKSTUHL | 86 DLCO | $E_{\text{cm}}^{\text{ee}} = 29 \text{ GeV}$ |
| 4.7 ± 0.5 ± 0.8 | 530 | 6 SCHMIDKE | 86 MRK2 | $E_{\text{cm}}^{\text{ee}} = 29 \text{ GeV}$ |
| 5.6 ± 0.4 ± 0.7 | | 5 FERNANDEZ | 85 MAC | $E_{\text{cm}}^{\text{ee}} = 29 \text{ GeV}$ |
| 6.2 ± 2.3 ± 1.7 | | BEHREND | 84 CELL | $E_{\text{cm}}^{\text{ee}} = 14,22 \text{ GeV}$ |

¹ BEHREND 90 value is not independent of BEHREND 90 $B(3h\nu_\tau \geq 1 \text{ neutrals}) + B(5\text{-prong})$.

² ALBRECHT 87L measure the product of branching ratios $B(3\pi^\pm \pi^0 \nu_\tau) B((e\bar{\nu} \text{ or } \mu\bar{\nu} \text{ or } \pi \text{ or } K \text{ or } \rho)\nu_\tau) = 0.029$ and use the PDG 86 values for the second branching ratio which sum to 0.69 ± 0.03 to get the quoted value.

³ BURCHAT 87 value is not independent of SCHMIDKE 86 value.

⁴ Contributions from kaons and from $>1\pi^0$ are subtracted. Not independent of (3-prong + $0\pi^0$) and (3-prong + $\geq 0\pi^0$) values.

⁵ Value obtained using paper's $R = B(h^- h^- h^+ \nu_\tau)/B(3\text{-prong})$ and current $B(3\text{-prong}) = 0.143$.

⁶ Not independent of SCHMIDKE 86 $h^- h^- h^+ \nu_\tau$ and $h^- h^- h^+(\geq 0\pi^0)\nu_\tau$ values.

$$\Gamma(h^- h^- h^+ \geq 1\pi^0 \nu_\tau(\text{ex. } K^0))/\Gamma_{\text{total}} \quad \Gamma_{72}/\Gamma$$

$$\begin{aligned} \Gamma_{72}/\Gamma = & (\Gamma_{78} + \Gamma_{85} + \Gamma_{86} + \Gamma_{103} + \Gamma_{107} + 0.2292\Gamma_{148} + 0.2292\Gamma_{150} + 0.2292\Gamma_{152} + \\ & 0.892\Gamma_{176} + 0.892\Gamma_{177} + 0.9078\Gamma_{178})/\Gamma \end{aligned}$$

| VALUE (%) | EVTS | DOCUMENT ID | TECN | COMMENT |
|-------------------------------|------|-------------|------|---------|
| 5.09 ±0.05 OUR FIT | | | | |
| 5.10 ±0.12 OUR AVERAGE | | | | |

• • • We use the following data for averages but not for fits. • • •

$5.106 \pm 0.083 \pm 0.103$ 10.1k ¹ ABDALLAH 06A DLEPH 1992–1995 LEP runs
 $5.09 \pm 0.10 \pm 0.23$ ² AKERS 95Y OPAL 1991–1994 LEP runs

• • • We do not use the following data for averages, fits, limits, etc. • • •

$4.95 \pm 0.29 \pm 0.65$ 570 DECAMP 92C ALEP Repl. by SCHAEEL 05C

¹ See footnote to ABDALLAH 06A $\Gamma(\tau^- \rightarrow h^- \nu_\tau)/\Gamma_{\text{total}}$ measurement for correlations with other measurements.

² Not independent of AKERS 95Y $B(h^- h^- h^+ \geq 0 \text{ neutrals} \nu_\tau \text{ (ex. } K_S^0 \rightarrow \pi^+ \pi^-))$ and $B(h^- h^- h^+ \geq 0 \text{ neutrals} \nu_\tau \text{ (ex. } K^0))/B(h^- h^- h^+ \geq 0 \text{ neutrals} \nu_\tau \text{ (ex. } K_S^0 \rightarrow \pi^+ \pi^-))$ values.

$\Gamma(h^- h^- h^+ \pi^0 \nu_\tau)/\Gamma_{\text{total}}$

Γ_{73}/Γ

$$\Gamma_{73}/\Gamma = (0.34598\Gamma_{41} + 0.34598\Gamma_{43} + \Gamma_{78} + \Gamma_{103} + \Gamma_{107} + 0.2292\Gamma_{150} + 0.892\Gamma_{176} + 0.892\Gamma_{177} + 0.0153\Gamma_{178})/\Gamma$$

| VALUE (%) | EVTS | DOCUMENT ID | TECN | COMMENT |
|-----------|------|-------------|------|---------|
|-----------|------|-------------|------|---------|

4.76 ± 0.05 OUR FIT

• • • We do not use the following data for averages, fits, limits, etc. • • •

$4.45 \pm 0.09 \pm 0.07$ 6.1k ¹ BUSKULIC 96 ALEP Repl. by SCHAEEL 05C

¹ BUSKULIC 96 quote $B(h^- h^- h^+ \pi^0 \nu_\tau \text{ (ex. } K^0)) = 4.30 \pm 0.09 \pm 0.09$. We add 0.15 to remove their K^0 correction and reduce the systematic error accordingly.

$\Gamma(h^- h^- h^+ \pi^0 \nu_\tau \text{ (ex. } K^0))/\Gamma_{\text{total}}$

Γ_{74}/Γ

$$\Gamma_{74}/\Gamma = (\Gamma_{78} + \Gamma_{103} + \Gamma_{107} + 0.2292\Gamma_{150} + 0.892\Gamma_{176} + 0.892\Gamma_{177} + 0.0153\Gamma_{178})/\Gamma$$

| VALUE (%) | EVTS | DOCUMENT ID | TECN | COMMENT |
|-----------|------|-------------|------|---------|
|-----------|------|-------------|------|---------|

4.57 ± 0.05 OUR FIT

4.45 ± 0.14 OUR AVERAGE Error includes scale factor of 1.2.

$4.545 \pm 0.106 \pm 0.103$ 8.9k ¹ ABDALLAH 06A DLEPH 1992–1995 LEP runs

$4.23 \pm 0.06 \pm 0.22$ 7.2k BAEST 95C CLEO $E_{\text{cm}}^{\text{ee}} \approx 10.6 \text{ GeV}$

¹ See footnote to ABDALLAH 06A $\Gamma(\tau^- \rightarrow h^- \nu_\tau)/\Gamma_{\text{total}}$ measurement for correlations with other measurements.

$\Gamma(h^- h^- h^+ \pi^0 \nu_\tau \text{ (ex. } K^0, \omega))/\Gamma_{\text{total}}$

$$\Gamma_{75}/\Gamma = (\Gamma_{78} + \Gamma_{103} + \Gamma_{107} + 0.2292\Gamma_{150})/\Gamma$$

| VALUE (%) | DOCUMENT ID |
|-----------|-------------|
|-----------|-------------|

2.79 ± 0.07 OUR FIT

$\Gamma(\pi^- \pi^+ \pi^- \pi^0 \nu_\tau)/\Gamma_{\text{total}}$

Γ_{76}/Γ

$$\Gamma_{76}/\Gamma = (0.34598\Gamma_{41} + \Gamma_{78} + 0.892\Gamma_{176} + 0.0153\Gamma_{178})/\Gamma$$

| VALUE (%) | DOCUMENT ID |
|-----------|-------------|
|-----------|-------------|

4.62 ± 0.05 OUR FIT

$\Gamma(\pi^- \pi^+ \pi^- \pi^0 \nu_\tau \text{ (ex. } K^0))/\Gamma_{\text{total}}$

Γ_{77}/Γ

$$\Gamma_{77}/\Gamma = (\Gamma_{78} + 0.892\Gamma_{176} + 0.0153\Gamma_{178})/\Gamma$$

| VALUE (%) | EVTS | DOCUMENT ID | TECN | COMMENT |
|-----------|------|-------------|------|---------|
|-----------|------|-------------|------|---------|

4.49 ± 0.05 OUR FIT

4.55 ± 0.13 OUR AVERAGE Error includes scale factor of 1.6.

$4.598 \pm 0.057 \pm 0.064$ 16k ¹ SCHAEEL 05C ALEP 1991–1995 LEP runs

$4.19 \pm 0.10 \pm 0.21$ ² EDWARDS 00A CLEO $4.7 \text{ fb}^{-1} E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$

¹SCHAEL 05C quote $(4.590 \pm 0.057 \pm 0.064)\%$. We add 0.008% to remove their correction for $\tau^- \rightarrow \pi^- \pi^0 \nu_\tau \rightarrow \pi^- \pi^0 \pi^+ \pi^- \nu_\tau$ decays. See footnote to SCHAEL 05C $\Gamma(\tau^- \rightarrow e^- \bar{\nu}_e \nu_\tau)/\Gamma_{\text{total}}$ measurement for correlations with other measurements.
²EDWARDS 00A quote $(4.19 \pm 0.10) \times 10^{-2}$ with a 5% systematic error.

| $\Gamma(\pi^- \pi^+ \pi^- \pi^0 \nu_\tau (\text{ex. } K^0, \omega))/\Gamma_{\text{total}}$ | Γ_{78}/Γ |
|--|----------------------|
| VALUE (%) | DOCUMENT ID |
| 2.74 ± 0.07 OUR FIT | |

| $\Gamma(h^- \rho \pi^0 \nu_\tau)/\Gamma(h^- h^- h^+ \pi^0 \nu_\tau)$ | Γ_{79}/Γ_{73} | | | |
|---|---------------------------|-------------|---------|--|
| VALUE | EVTS | DOCUMENT ID | TECN | COMMENT |
| $\bullet \bullet \bullet$ We do not use the following data for averages, fits, limits, etc. $\bullet \bullet \bullet$ | | | | |
| $0.30 \pm 0.04 \pm 0.02$ | 393 | ALBRECHT | 91D ARG | $E_{\text{cm}}^{\text{ee}} = 9.4\text{--}10.6 \text{ GeV}$ |

| $\Gamma(h^- \rho^+ h^- \nu_\tau)/\Gamma(h^- h^- h^+ \pi^0 \nu_\tau)$ | Γ_{80}/Γ_{73} | | | |
|---|---------------------------|-------------|---------|--|
| VALUE | EVTS | DOCUMENT ID | TECN | COMMENT |
| $\bullet \bullet \bullet$ We do not use the following data for averages, fits, limits, etc. $\bullet \bullet \bullet$ | | | | |
| $0.10 \pm 0.03 \pm 0.04$ | 142 | ALBRECHT | 91D ARG | $E_{\text{cm}}^{\text{ee}} = 9.4\text{--}10.6 \text{ GeV}$ |

| $\Gamma(h^- \rho^- h^+ \nu_\tau)/\Gamma(h^- h^- h^+ \pi^0 \nu_\tau)$ | Γ_{81}/Γ_{73} | | | |
|---|---------------------------|-------------|---------|--|
| VALUE | EVTS | DOCUMENT ID | TECN | COMMENT |
| $\bullet \bullet \bullet$ We do not use the following data for averages, fits, limits, etc. $\bullet \bullet \bullet$ | | | | |
| $0.26 \pm 0.05 \pm 0.01$ | 370 | ALBRECHT | 91D ARG | $E_{\text{cm}}^{\text{ee}} = 9.4\text{--}10.6 \text{ GeV}$ |

| $\Gamma(h^- h^- h^+ \geq 2\pi^0 \nu_\tau (\text{ex. } K^0))/\Gamma_{\text{total}}$ | Γ_{82}/Γ | | | |
|---|----------------------|-----------------------|----------|--------------------|
| $\Gamma_{82}/\Gamma = (\Gamma_{85} + \Gamma_{86} + 0.2292\Gamma_{148} + 0.2292\Gamma_{152} + 0.892\Gamma_{178})/\Gamma$ | | | | |
| VALUE (%) | EVTS | DOCUMENT ID | TECN | COMMENT |
| 0.517 ± 0.031 OUR FIT | | | | |
| $0.561 \pm 0.068 \pm 0.095$ | 1.3k | ¹ ABDALLAH | 06A DLPH | 1992–1995 LEP runs |

¹See footnote to ABDALLAH 06A $\Gamma(\tau^- \rightarrow h^- \nu_\tau)/\Gamma_{\text{total}}$ measurement for correlations with other measurements.

| $\Gamma(h^- h^- h^+ 2\pi^0 \nu_\tau)/\Gamma_{\text{total}}$ | Γ_{83}/Γ |
|---|----------------------|
| $\Gamma_{83}/\Gamma = (0.4247\Gamma_{48} + \Gamma_{85} + 0.2292\Gamma_{148} + 0.2292\Gamma_{152} + 0.892\Gamma_{178})/\Gamma$ | |
| VALUE (%) | DOCUMENT ID |
| 0.505 ± 0.031 OUR FIT | |

| $\Gamma(h^- h^- h^+ 2\pi^0 \nu_\tau (\text{ex. } K^0))/\Gamma_{\text{total}}$ | Γ_{84}/Γ | | | |
|---|----------------------|---------------------|----------|--------------------|
| $\Gamma_{84}/\Gamma = (\Gamma_{85} + 0.2292\Gamma_{148} + 0.2292\Gamma_{152} + 0.892\Gamma_{178})/\Gamma$ | | | | |
| VALUE (%) | EVTS | DOCUMENT ID | TECN | COMMENT |
| 0.495 ± 0.031 OUR FIT | | | | |
| $0.435 \pm 0.030 \pm 0.035$ | 2.6k | ¹ SCHAEL | 05C ALEP | 1991–1995 LEP runs |

$\bullet \bullet \bullet$ We do not use the following data for averages, fits, limits, etc. $\bullet \bullet \bullet$

$0.50 \pm 0.07 \pm 0.07$ 1.8k BUSKULIC 96 ALEP Repl. by SCHAEL 05C

¹SCHAEL 05C quote $(0.392 \pm 0.030 \pm 0.035)\%$. We add 0.043% to remove their correction for $\tau^- \rightarrow \pi^- \eta \pi^0 \nu_\tau \rightarrow \pi^- \pi^+ \pi^- 2\pi^0 \nu_\tau$ and $\tau^- \rightarrow K^*(892)^- \eta \nu_\tau \rightarrow K^- \pi^+ \pi^- 2\pi^0 \nu_\tau$ decays. See footnote to SCHAEL 05C $\Gamma(\tau^- \rightarrow e^- \bar{\nu}_e \nu_\tau)/\Gamma_{\text{total}}$ measurement for correlations with other measurements.

$$\Gamma(h^- h^- h^+ 2\pi^0 \nu_\tau (\text{ex. } K^0)) / \Gamma(h^- h^- h^+ \geq 0 \text{ neutrals} \geq 0 K_L^0 \nu_\tau) \quad \Gamma_{84}/\Gamma_{62}$$

$$\begin{aligned} \Gamma_{84}/\Gamma_{62} = & (\Gamma_{85} + 0.2292\Gamma_{148} + 0.2292\Gamma_{152} + 0.892\Gamma_{178}) / (0.34598\Gamma_{36} + \\ & 0.34598\Gamma_{38} + 0.34598\Gamma_{41} + 0.34598\Gamma_{43} + 0.4247\Gamma_{48} + 0.6920\Gamma_{49} + 0.8494\Gamma_{52} + \\ & 0.6920\Gamma_{56} + 0.6534\Gamma_{61} + \Gamma_{70} + \Gamma_{78} + \Gamma_{85} + \Gamma_{89} + \Gamma_{97} + \Gamma_{103} + \Gamma_{106} + \Gamma_{107} + \\ & 0.2810\Gamma_{148} + 0.2292\Gamma_{149} + 0.2810\Gamma_{150} + 0.2810\Gamma_{152} + 0.3759\Gamma_{154} + 0.3268\Gamma_{158} + \\ & 0.7259\Gamma_{168} + 0.9078\Gamma_{176} + 0.9078\Gamma_{177} + 0.9078\Gamma_{178} + 0.892\Gamma_{180}) \end{aligned}$$

| <u>VALUE (units 10^{-2})</u> | <u>EVTS</u> | <u>DOCUMENT ID</u> | <u>TECN</u> | <u>COMMENT</u> |
|---|-------------|--------------------|-------------|---|
| 3.26 ± 0.20 OUR FIT | | | | |
| $3.4 \pm 0.2 \pm 0.3$ | 668 | BORTOLETTO93 | CLEO | $E_{\text{cm}}^{ee} \approx 10.6 \text{ GeV}$ |

$$\Gamma(h^- h^- h^+ 2\pi^0 \nu_\tau (\text{ex. } K^0, \omega, \eta)) / \Gamma_{\text{total}} \quad \Gamma_{85}/\Gamma$$

| <u>VALUE (units 10^{-4})</u> | <u>DOCUMENT ID</u> |
|---|--------------------|
| 10 ± 4 OUR FIT | |

$$\Gamma(h^- h^- h^+ 3\pi^0 \nu_\tau) / \Gamma_{\text{total}} \quad \Gamma_{86}/\Gamma = (0.4247\Gamma_{52} + \Gamma_{87} + 0.1131\Gamma_{154}) / \Gamma$$

| <u>VALUE (units 10^{-4})</u> | <u>CL%</u> | <u>EVTS</u> | <u>DOCUMENT ID</u> | <u>TECN</u> | <u>COMMENT</u> |
|---|------------|---------------|--------------------|---|----------------|
| 2.12 ± 0.30 OUR FIT | | | | | |
| $2.2 \pm 0.3 \pm 0.4$ | 139 | ANASTASSOV 01 | CLEO | $E_{\text{cm}}^{ee} = 10.6 \text{ GeV}$ | |
| • • • We do not use the following data for averages, fits, limits, etc. • • • | | | | | |

| | | | | | |
|--------------------------|-----|-----------------------|-----|------|-------------------------|
| < 4.9 | 95 | SCHAEL | 05C | ALEP | 1991-1995 LEP runs |
| $2.85 \pm 0.56 \pm 0.51$ | 57 | ANDERSON | 97 | CLEO | Repl. by ANAS-TASSOV 01 |
| $11 \pm 4 \pm 5$ | 440 | ¹ BUSKULIC | 96 | ALEP | Repl. by SCHAEL 05C |

¹ BUSKULIC 96 state their measurement is for $B(h^- h^- h^+ \geq 3\pi^0 \nu_\tau)$. We assume that $B(h^- h^- h^+ \geq 4\pi^0 \nu_\tau)$ is very small.

$$\Gamma(2\pi^- \pi^+ 3\pi^0 \nu_\tau (\text{ex. } K^0)) / \Gamma_{\text{total}} \quad \Gamma_{87}/\Gamma$$

$$\Gamma_{87}/\Gamma = (\Gamma_{89} + 0.2292\Gamma_{149} + 0.3268\Gamma_{158} + 0.892\Gamma_{180}) / \Gamma$$

| <u>VALUE (units 10^{-4})</u> | <u>DOCUMENT ID</u> | <u>TECN</u> | <u>COMMENT</u> |
|---|--------------------|-------------|----------------|
| 1.94 ± 0.30 OUR FIT | | | |

• • • We use the following data for averages but not for fits. • • •

| | | | | | |
|--|-------------------|-----|------|-----------------------|---|
| $2.07 \pm 0.18 \pm 0.37$ | ¹ LEES | 12X | BABR | 468 fb^{-1} | $E_{\text{cm}}^{ee} = 10.6 \text{ GeV}$ |
|--|-------------------|-----|------|-----------------------|---|

¹ Not independent of LEES 12X $\Gamma(\tau^- \rightarrow \eta \pi^- \pi^+ \pi^- \nu_\tau (\text{ex. } K^0)) / \Gamma$, $\Gamma(\tau^- \rightarrow \eta \pi^- \pi^0 \nu_\tau) / \Gamma$, $\Gamma(\tau^- \rightarrow \pi^- \omega 2\pi^0 \nu_\tau) / \Gamma$, and $\Gamma(\tau^- \rightarrow f_1(1285) \pi^- \nu_\tau \rightarrow \eta \pi^- \pi^+ \pi^- \nu_\tau) / \Gamma$ values.

$$\Gamma(2\pi^- \pi^+ 3\pi^0 \nu_\tau (\text{ex. } K^0, \eta, f_1(1285))) / \Gamma_{\text{total}} \quad \Gamma_{88}/\Gamma$$

| <u>VALUE (units 10^{-4})</u> | <u>DOCUMENT ID</u> | <u>TECN</u> | <u>COMMENT</u> |
|--|--------------------|-------------|----------------|
| $1.69 \pm 0.08 \pm 0.43$ | LEES | 12X | BABR |

$$\Gamma(2\pi^- \pi^+ 3\pi^0 \nu_\tau (\text{ex. } K^0, \eta, \omega, f_1(1285))) / \Gamma_{\text{total}} \quad \Gamma_{89}/\Gamma$$

| <u>VALUE (units 10^{-5})</u> | <u>DOCUMENT ID</u> | <u>TECN</u> | <u>COMMENT</u> |
|---|--------------------|-------------|----------------|
| 1.4 ± 2.7 OUR FIT | | | |

| | | | | | |
|---|-------------------|-----|------|-----------------------|---|
| $1.0 \pm 0.8 \pm 3.0$ | ¹ LEES | 12X | BABR | 468 fb^{-1} | $E_{\text{cm}}^{ee} = 10.6 \text{ GeV}$ |
|---|-------------------|-----|------|-----------------------|---|

¹ LEES 12X measurement corresponds to the lower limit of $< 5.8 \times 10^{-5}$ at 90% CL. ■

$$\Gamma(K^- h^+ h^- \geq 0 \text{ neutrals } \nu_\tau)/\Gamma_{\text{total}} \quad \Gamma_{90}/\Gamma$$

$$\Gamma_{90}/\Gamma = (0.34598\Gamma_{38} + 0.34598\Gamma_{43} + \Gamma_{97} + \Gamma_{103} + \Gamma_{106} + \Gamma_{107} + 0.2810\Gamma_{150} + 0.489\Gamma_{168} + 0.9078\Gamma_{177})/\Gamma$$

| <u>VALUE (%)</u> | <u>CL %</u> | <u>DOCUMENT ID</u> | <u>TECN</u> | <u>COMMENT</u> |
|----------------------------|-------------|--------------------|-------------|--|
| 0.629±0.014 OUR FIT | | | | |
| <0.6 | 90 | AIHARA | 84C TPC | $E_{\text{cm}}^{\text{ee}} = 29 \text{ GeV}$ |

$$\Gamma(K^- h^+ \pi^- \nu_\tau (\text{ex. } K^0))/\Gamma_{\text{total}} \quad \Gamma_{91}/\Gamma = (\Gamma_{97} + \Gamma_{106} + 0.0153\Gamma_{177})/\Gamma$$

| <u>VALUE (%)</u> | <u>DOCUMENT ID</u> |
|----------------------------|--------------------|
| 0.437±0.007 OUR FIT | |

$$\Gamma(K^- h^+ \pi^- \nu_\tau (\text{ex. } K^0))/\Gamma(\pi^- \pi^+ \pi^- \nu_\tau (\text{ex. } K^0)) \quad \Gamma_{91}/\Gamma_{68}$$

$$\Gamma_{91}/\Gamma_{68} = (\Gamma_{97} + \Gamma_{106} + 0.0153\Gamma_{177})/(\Gamma_{70} + 0.0153\Gamma_{176})$$

| <u>VALUE (%)</u> | <u>EVTS</u> | <u>DOCUMENT ID</u> | <u>TECN</u> | <u>COMMENT</u> |
|--------------------------|-------------|--------------------|-------------|---|
| 4.84±0.08 OUR FIT | | | | |
| 5.44±0.21±0.53 | 7.9k | RICHICHI | 99 | CLEO $E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$ |

$$\Gamma(K^- h^+ \pi^- \pi^0 \nu_\tau (\text{ex. } K^0))/\Gamma_{\text{total}} \quad \Gamma_{92}/\Gamma$$

$$\Gamma_{92}/\Gamma = (\Gamma_{103} + \Gamma_{107} + 0.2292\Gamma_{150} + 0.892\Gamma_{177})/\Gamma$$

| <u>VALUE (units 10^{-4})</u> | <u>DOCUMENT ID</u> |
|---|--------------------|
| 8.6±1.2 OUR FIT | |

$$\Gamma(K^- h^+ \pi^- \pi^0 \nu_\tau (\text{ex. } K^0))/\Gamma(\pi^- \pi^+ \pi^- \pi^0 \nu_\tau (\text{ex. } K^0)) \quad \Gamma_{92}/\Gamma_{77}$$

$$\Gamma_{92}/\Gamma_{77} = (\Gamma_{103} + \Gamma_{107} + 0.2292\Gamma_{150} + 0.892\Gamma_{177})/(\Gamma_{78} + 0.892\Gamma_{176} + 0.0153\Gamma_{178})$$

| <u>VALUE (%)</u> | <u>EVTS</u> | <u>DOCUMENT ID</u> | <u>TECN</u> | <u>COMMENT</u> |
|--------------------------|-------------|--------------------|-------------|---|
| 1.91±0.26 OUR FIT | | | | |
| 2.61±0.45±0.42 | 719 | RICHICHI | 99 | CLEO $E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$ |

$$\Gamma(K^- \pi^+ \pi^- \geq 0 \text{ neutrals } \nu_\tau)/\Gamma_{\text{total}} \quad \Gamma_{93}/\Gamma$$

$$\Gamma_{93}/\Gamma = (0.34598\Gamma_{38} + 0.34598\Gamma_{43} + \Gamma_{97} + \Gamma_{103} + 0.2810\Gamma_{150} + 0.9078\Gamma_{177})/\Gamma$$

| <u>VALUE (%)</u> | <u>EVTS</u> | <u>DOCUMENT ID</u> | <u>TECN</u> | <u>COMMENT</u> |
|----------------------------|-------------|--------------------|-------------|----------------|
| 0.477±0.014 OUR FIT | | | | |

$$0.58 \begin{array}{l} +0.15 \\ -0.13 \end{array} \pm 0.12 \quad 20 \quad ^1 \text{ BAUER} \quad 94 \quad \text{TPC} \quad E_{\text{cm}}^{\text{ee}} = 29 \text{ GeV}$$

• • • We do not use the following data for averages, fits, limits, etc. • • •

$$0.22 \begin{array}{l} +0.16 \\ -0.13 \end{array} \pm 0.05 \quad 9 \quad ^2 \text{ MILLS} \quad 85 \quad \text{DLCO} \quad E_{\text{cm}}^{\text{ee}} = 29 \text{ GeV}$$

¹ We multiply 0.58% by 0.20, the relative systematic error quoted by BAUER 94, to obtain the systematic error.

² Error correlated with MILLS 85 ($K K \pi \nu$) value. We multiply 0.22% by 0.23, the relative systematic error quoted by MILLS 85, to obtain the systematic error.

$$\Gamma(K^-\pi^+\pi^- \geq 0\pi^0\nu_\tau(\text{ex.}K^0))/\Gamma_{\text{total}} \quad \Gamma_{94}/\Gamma$$

$$\Gamma_{94}/\Gamma = (\Gamma_{97} + \Gamma_{103} + 0.2292\Gamma_{150} + 0.9078\Gamma_{177})/\Gamma$$

| VALUE (%) | DOCUMENT ID | TECN | COMMENT |
|-----------|-------------|------|---------|
|-----------|-------------|------|---------|

0.373±0.013 OUR FIT

0.30 ±0.05 OUR AVERAGE

• • • We use the following data for averages but not for fits. • • •

| | | | | |
|-------------------|---------------------|-----|------|--------------------|
| 0.343±0.073±0.031 | ABBIENDI | 00D | OPAL | 1990–1995 LEP runs |
| 0.275±0.064 | ¹ BARATE | 98 | ALEP | 1991–1995 LEP runs |

¹ Not independent of BARATE 98 $\Gamma(\tau^- \rightarrow K^-\pi^+\pi^-\nu_\tau)/\Gamma_{\text{total}}$ and $\Gamma(\tau^- \rightarrow K^-\pi^+\pi^-\pi^0\nu_\tau)/\Gamma_{\text{total}}$ values.

$$\Gamma(K^-\pi^+\pi^-\nu_\tau)/\Gamma_{\text{total}}$$

$$\Gamma_{95}/\Gamma = (0.34598\Gamma_{38} + \Gamma_{97} + 0.0153\Gamma_{177})/\Gamma$$

| VALUE (%) | DOCUMENT ID |
|-----------|-------------|
|-----------|-------------|

0.345±0.007 OUR FIT

$$\Gamma(K^-\pi^+\pi^-\nu_\tau(\text{ex.}K^0))/\Gamma_{\text{total}}$$

$$\Gamma_{96}/\Gamma = (\Gamma_{97} + 0.0153\Gamma_{177})/\Gamma$$

| VALUE (%) | EVTS | DOCUMENT ID | TECN | COMMENT |
|-----------|------|-------------|------|---------|
|-----------|------|-------------|------|---------|

0.293±0.007 OUR FIT

0.290±0.018 OUR AVERAGE

Error includes scale factor of 2.4. See the ideogram below.

| | | | | | | |
|---|------|---------------------|-----|------|--------------------------------------|--------------------------------------|
| 0.330±0.001 ^{+0.016} _{-0.017} | 794k | ¹ LEE | 10 | BELL | 666 fb ⁻¹ | $E_{\text{cm}}^{\text{ee}}=10.6$ GeV |
| 0.273±0.002±0.009 | 70k | ² AUBERT | 08 | BABR | 342 fb ⁻¹ | $E_{\text{cm}}^{\text{ee}}=10.6$ GeV |
| 0.415±0.053±0.040 | 269 | ABBIENDI | 04J | OPAL | 1991–1995 LEP runs | |
| 0.384±0.014±0.038 | 3.5k | ³ BRIERE | 03 | CLEO | $E_{\text{cm}}^{\text{ee}}=10.6$ GeV | |
| 0.214±0.037±0.029 | | BARATE | 98 | ALEP | 1991–1995 LEP runs | |

• • • We use the following data for averages but not for fits. • • •

0.346±0.023±0.056 158 ⁴ RICHICHI 99 CLEO $E_{\text{cm}}^{\text{ee}}=10.6$ GeV

• • • We do not use the following data for averages, fits, limits, etc. • • •

0.360±0.082±0.048 ABBIENDI 00D OPAL 1990–1995 LEP runs

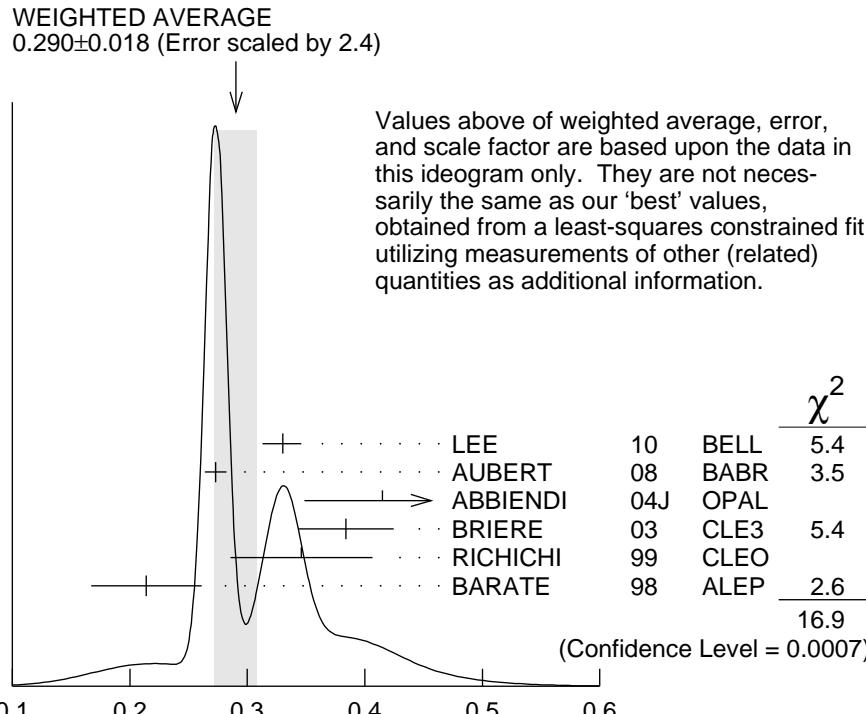
¹ See footnote to LEE 10 $\Gamma(\tau^- \rightarrow \pi^-\pi^+\pi^-\nu_\tau(\text{ex.}K^0))/\Gamma_{\text{total}}$ measurement for correlations with other measurements. Not independent of LEE 10 $\Gamma(\tau^- \rightarrow K^-\pi^+\pi^-\nu_\tau(\text{ex.}K^0))/\Gamma(\tau^- \rightarrow \pi^-\pi^+\pi^-\nu_\tau(\text{ex.}K^0))$ value.

² See footnote to AUBERT 08 $\Gamma(\tau^- \rightarrow \pi^-\pi^+\pi^-\nu_\tau(\text{ex.}K^0))/\Gamma_{\text{total}}$ measurement for correlations with other measurements.

³ 47% correlated with BRIERE 03 $\tau^- \rightarrow \pi^-\pi^+\pi^-\nu_\tau$ and 34% correlated with $\tau^- \rightarrow K^-K^+\pi^-\nu_\tau$ because of a common 5% normalization error.

⁴ Not independent of RICHICHI 99

$\Gamma(\tau^- \rightarrow K^-h^+\pi^-\nu_\tau(\text{ex.}K^0))/\Gamma(\tau^- \rightarrow \pi^-\pi^+\pi^-\nu_\tau(\text{ex.}K^0))$, $\Gamma(\tau^- \rightarrow K^-K^+\pi^-\nu_\tau)/\Gamma(\tau^- \rightarrow \pi^-\pi^+\pi^-\nu_\tau(\text{ex.}K^0))$ and BALEST 95C $\Gamma(\tau^- \rightarrow h^-h^-h^+\nu_\tau(\text{ex.}K^0))/\Gamma_{\text{total}}$ values.



$$\Gamma(K^-\pi^+\pi^-\nu_\tau(\text{ex. } K^0))/\Gamma_{\text{total}} (\%)$$

$$\begin{aligned} \Gamma(K^-\pi^+\pi^-\nu_\tau(\text{ex. } K^0))/\Gamma(\pi^-\pi^+\pi^-\nu_\tau(\text{ex. } K^0)) \\ \Gamma_{96}/\Gamma_{68} = (\Gamma_{97} + 0.0153\Gamma_{177})/(\Gamma_{70} + 0.0153\Gamma_{176}) \end{aligned}$$

| VALUE (units 10^{-2}) | EVTS | DOCUMENT ID | TECN | COMMENT |
|--------------------------|------|-------------|------|---------|
| 3.25±0.07 OUR FIT | | | | |

• • • We use the following data for averages but not for fits. • • •

$$3.92 \pm 0.02^{+0.15}_{-0.16} \quad 794k \quad ^1 \text{LEE} \quad 10 \quad \text{BELL} \quad 666 \text{ fb}^{-1} \quad E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$$

¹ Not independent of LEE 10 $\Gamma(\tau^- \rightarrow K^-\pi^+\pi^-\nu_\tau(\text{ex. } K^0))/\Gamma_{\text{total}}$ and $\Gamma(\tau^- \rightarrow \pi^-\pi^+\pi^-\nu_\tau(\text{ex. } K^0))/\Gamma_{\text{total}}$ values.

$$\Gamma(K^-\pi^+\pi^-\nu_\tau(\text{ex. } K^0,\omega))/\Gamma_{\text{total}} \quad \Gamma_{97}/\Gamma$$

| VALUE (units 10^{-3}) | DOCUMENT ID |
|--------------------------|-------------|
| 2.93±0.07 OUR FIT | |

$$\Gamma(K^-\rho^0\nu_\tau \rightarrow K^-\pi^+\pi^-\nu_\tau)/\Gamma(K^-\pi^+\pi^-\nu_\tau(\text{ex. } K^0)) \quad \Gamma_{98}/\Gamma_{96}$$

| VALUE | DOCUMENT ID | TECN | COMMENT |
|-----------------------|--------------------|----------|--|
| 0.48±0.14±0.10 | ¹ ASNER | 00B CLEO | $E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$ |

• • • We do not use the following data for averages, fits, limits, etc. • • •

$$0.39 \pm 0.14 \quad ^2 \text{BARATE} \quad 99R \quad \text{ALEP} \quad 1991\text{--}1995 \text{ LEP runs}$$

¹ ASNER 00B assume $\tau^- \rightarrow K^-\pi^+\pi^-\nu_\tau(\text{ex. } K^0)$ decays proceed only through $K\rho$ and $K^*\pi$ intermediate states. They assume the resonance structure of $\tau^- \rightarrow K^-\pi^+\pi^-\nu_\tau(\text{ex. } K^0)$ decays is dominated by $K_1(1270)^-$ and $K_1(1400)^-$ resonances, and assume $B(K_1(1270) \rightarrow K^*(892)\pi) = (16 \pm 5)\%$, $B(K_1(1270) \rightarrow K\rho) = (42 \pm 6)\%$, and $B(K_1(1400) \rightarrow K\rho) = 0$.

² BARATE 99R assume $\tau^- \rightarrow K^-\pi^+\pi^-\nu_\tau$ (ex. K^0) decays proceed only through $K\rho$ and $K^*\pi$ intermediate states. The quoted error is statistical only.

$$\Gamma(K^-\pi^+\pi^-\pi^0\nu_\tau)/\Gamma_{\text{total}} \quad \Gamma_{99}/\Gamma$$

$$\Gamma_{99}/\Gamma = (0.34598\Gamma_{43} + \Gamma_{103} + 0.2292\Gamma_{150} + 0.892\Gamma_{177})/\Gamma$$

| VALUE (units 10^{-4}) | DOCUMENT ID |
|--|-------------|
| 13.1 ± 1.2 OUR FIT | |

$$\Gamma(K^-\pi^+\pi^-\pi^0\nu_\tau(\text{ex. } K^0))/\Gamma_{\text{total}} \quad \Gamma_{100}/\Gamma$$

$$\Gamma_{100}/\Gamma = (\Gamma_{103} + 0.2292\Gamma_{150} + 0.892\Gamma_{177})/\Gamma$$

| VALUE (units 10^{-4}) | CL% | DOCUMENT ID | TECN | COMMENT |
|---|-----|-------------|------|---------|
| 7.9 ± 1.2 OUR FIT | | | | |

7.3±1.2 OUR AVERAGE

| | | | | |
|-----------------------|-------------------|----|------|---|
| $7.4 \pm 0.8 \pm 1.1$ | ¹ ARMS | 05 | CLE3 | $7.6 \text{ fb}^{-1}, E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$ |
| $6.1 \pm 3.9 \pm 1.8$ | BARATE | 98 | ALEP | 1991–1995 LEP runs |

• • • We use the following data for averages but not for fits. • • •

| | | | | |
|-----------------------|-----------------------|----|------|--|
| $7.5 \pm 2.6 \pm 1.8$ | ² RICHICHI | 99 | CLEO | $E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$ |
|-----------------------|-----------------------|----|------|--|

• • • We do not use the following data for averages, fits, limits, etc. • • •

| | | | | | |
|-----|----|----------|-----|------|--------------------|
| <17 | 95 | ABBIENDI | 00D | OPAL | 1990–1995 LEP runs |
|-----|----|----------|-----|------|--------------------|

¹ Not independent of ARMS 05 $\Gamma(\tau^- \rightarrow K^-\pi^+\pi^-\pi^0\nu_\tau(\text{ex. } K^0, \omega)) / \Gamma_{\text{total}}$ and $\Gamma(\tau^- \rightarrow K^-\omega\nu_\tau) / \Gamma_{\text{total}}$ values.

² Not independent of RICHICHI 99

$\Gamma(\tau^- \rightarrow K^-h^+\pi^-\nu_\tau(\text{ex. } K^0))/\Gamma(\tau^- \rightarrow \pi^-\pi^+\pi^-\nu_\tau(\text{ex. } K^0)), \Gamma(\tau^- \rightarrow K^-K^+\pi^-\nu_\tau)/\Gamma(\tau^- \rightarrow \pi^-\pi^+\pi^-\nu_\tau(\text{ex. } K^0))$ and BALEST 95C $\Gamma(\tau^- \rightarrow h^-h^-h^+\nu_\tau(\text{ex. } K^0))/\Gamma_{\text{total}}$ values.

$$\Gamma(K^-\pi^+\pi^-\pi^0\nu_\tau(\text{ex. } K^0, \eta))/\Gamma_{\text{total}} \quad \Gamma_{101}/\Gamma = (\Gamma_{103} + 0.892\Gamma_{177})/\Gamma$$

| VALUE (units 10^{-4}) | DOCUMENT ID |
|---|-------------|
| 7.6 ± 1.2 OUR FIT | |

$$\Gamma(K^-\pi^+\pi^-\pi^0\nu_\tau(\text{ex. } K^0, \omega))/\Gamma_{\text{total}} \quad \Gamma_{102}/\Gamma$$

| VALUE (units 10^{-4}) | EVTS | DOCUMENT ID | TECN | COMMENT |
|---|------|-------------|------|--|
| $3.7 \pm 0.5 \pm 0.8$ | 833 | ARMS | 05 | CLE3 $7.6 \text{ fb}^{-1}, E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$ |

$$\Gamma(K^-\pi^+\pi^-\pi^0\nu_\tau(\text{ex. } K^0, \omega, \eta))/\Gamma_{\text{total}} \quad \Gamma_{103}/\Gamma$$

| VALUE (units 10^{-4}) | DOCUMENT ID |
|---|-------------|
| 3.9 ± 1.4 OUR FIT | |

$$\Gamma(K^-\pi^+K^-\geq 0 \text{ neut. } \nu_\tau)/\Gamma_{\text{total}} \quad \Gamma_{104}/\Gamma$$

| VALUE (%) | CL% | DOCUMENT ID | TECN | COMMENT |
|-----------------|-----|-------------|------|--|
| <0.09 | 95 | BAUER | 94 | TPC $E_{\text{cm}}^{\text{ee}} = 29 \text{ GeV}$ |

| $\Gamma(K^- K^+ \pi^- \geq 0 \text{ neut. } \nu_\tau)/\Gamma_{\text{total}}$ | $\Gamma_{105}/\Gamma = (\Gamma_{106} + \Gamma_{107})/\Gamma$ | | | |
|---|--|---------------------|----------|---------------------------------------|
| VALUE (%) | EVTS | DOCUMENT ID | TECN | COMMENT |
| 0.1496 ± 0.0033 OUR FIT | | | | |
| 0.203 ± 0.031 OUR AVERAGE | | | | |
| 0.159 $\pm 0.053 \pm 0.020$ | | ABBIENDI | 00D OPAL | 1990–1995 LEP runs |
| 0.15 $\pm 0.09 \pm 0.03$ | 4 | ¹ BAUER | 94 TPC | $E_{\text{cm}}^{ee} = 29 \text{ GeV}$ |
| $\bullet \bullet \bullet$ We use the following data for averages but not for fits. $\bullet \bullet \bullet$ | | | | |
| 0.238 ± 0.042 | | ² BARATE | 98 ALEP | 1991–1995 LEP runs |
| ¹ We multiply 0.15% by 0.20, the relative systematic error quoted by BAUER 94, to obtain the systematic error. | | | | |
| ² Not independent of BARATE 98 $\Gamma(\tau^- \rightarrow K^- K^+ \pi^- \nu_\tau)/\Gamma_{\text{total}}$ and $\Gamma(\tau^- \rightarrow K^- K^+ \pi^- \pi^0 \nu_\tau)/\Gamma_{\text{total}}$ values. | | | | |

| $\Gamma(K^- K^+ \pi^- \nu_\tau)/\Gamma_{\text{total}}$ | Γ_{106}/Γ | | | |
|---|-----------------------|-----------------------|----------|---|
| VALUE (units 10^{-3}) | EVTS | DOCUMENT ID | TECN | COMMENT |
| 1.435 ± 0.027 OUR FIT | | | | |
| 1.43 ± 0.07 OUR AVERAGE | | | | Error includes scale factor of 2.4. See the ideogram below. |
| 1.55 $\pm 0.01 \pm 0.06$ | 108k | ¹ LEE | 10 BELL | $666 \text{ fb}^{-1} E_{\text{cm}}^{ee} = 10.6 \text{ GeV}$ |
| $1.346 \pm 0.010 \pm 0.036$ | 18k | ² AUBERT | 08 BABR | $342 \text{ fb}^{-1} E_{\text{cm}}^{ee} = 10.6 \text{ GeV}$ |
| 1.55 $\pm 0.06 \pm 0.09$ | 932 | ³ BRIERE | 03 CLE3 | $E_{\text{cm}}^{ee} = 10.6 \text{ GeV}$ |
| 1.63 $\pm 0.21 \pm 0.17$ | | BARATE | 98 ALEP | 1991–1995 LEP runs |
| $\bullet \bullet \bullet$ We use the following data for averages but not for fits. $\bullet \bullet \bullet$ | | | | |
| 0.87 $\pm 0.56 \pm 0.40$ | | ABBIENDI | 00D OPAL | 1990–1995 LEP runs |
| 1.45 $\pm 0.13 \pm 0.28$ | 2.3k | ⁴ RICHICHI | 99 CLEO | $E_{\text{cm}}^{ee} = 10.6 \text{ GeV}$ |
| $\bullet \bullet \bullet$ We do not use the following data for averages, fits, limits, etc. $\bullet \bullet \bullet$ | | | | |
| 2.2 $\pm 1.7 \pm 0.5$ | 9 | ⁵ MILLS | 85 DLCO | $E_{\text{cm}}^{ee} = 29 \text{ GeV}$ |

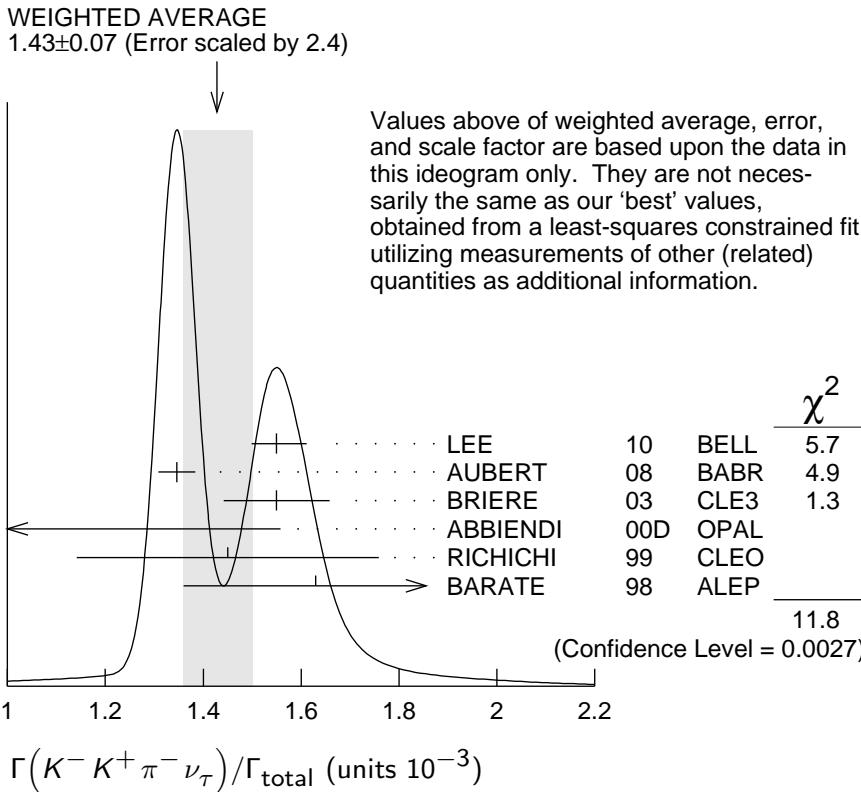
¹ See footnote to LEE 10 $\Gamma(\tau^- \rightarrow \pi^- \pi^+ \pi^- \nu_\tau (\text{ex. } K^0))/\Gamma_{\text{total}}$ measurement for correlations with other measurements. Not independent of LEE 10 $\Gamma(\tau^- \rightarrow K^- K^+ \pi^- \nu_\tau)/\Gamma(\tau^- \rightarrow \pi^- \pi^+ \pi^- \nu_\tau (\text{ex. } K^0))$ value.

² See footnote to AUBERT 08 $\Gamma(\tau^- \rightarrow \pi^- \pi^+ \pi^- \nu_\tau (\text{ex. } K^0))/\Gamma_{\text{total}}$ measurement for correlations with other measurements.

³ 71% correlated with BRIERE 03 $\tau^- \rightarrow \pi^- \pi^+ \pi^- \nu_\tau$ and 34% correlated with $\tau \rightarrow K^- \pi^+ \pi^- \nu_\tau$ because of a common 5% normalization error.

⁴ Not independent of RICHICHI 99 $\Gamma(\tau^- \rightarrow K^- K^+ \pi^- \nu_\tau)/\Gamma(\tau^- \rightarrow \pi^- \pi^+ \pi^- \nu_\tau (\text{ex. } K^0))$ and BALEST 95C $\Gamma(\tau^- \rightarrow h^- h^- h^+ \nu_\tau (\text{ex. } K^0))/\Gamma_{\text{total}}$ values.

⁵ Error correlated with MILLS 85 ($K \pi \pi \pi^0 \nu$) value. We multiply 0.22% by 0.23, the relative systematic error quoted by MILLS 85, to obtain the systematic error.



$$\Gamma(K^- K^+ \pi^- \nu_\tau) / \Gamma(\pi^- \pi^+ \pi^- \nu_\tau (\text{ex. } K^0)) \quad \Gamma_{106} / \Gamma_{68}$$

$$\Gamma_{106} / \Gamma_{68} = \Gamma_{106} / (\Gamma_{70} + 0.0153 \Gamma_{176})$$

| VALUE (%) | EVTS | DOCUMENT ID | TECN | COMMENT |
|-----------|------|-------------|------|---------|
|-----------|------|-------------|------|---------|

1.592±0.030 OUR FIT

1.83 ±0.05 OUR AVERAGE

$1.60 \pm 0.15 \pm 0.30$ 2.3k RICHICHI 99 CLEO $E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$

• • • We use the following data for averages but not for fits. • • •

$1.84 \pm 0.01 \pm 0.05$ 108k ¹ LEE 10 BELL $666 \text{ fb}^{-1} E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$

¹ Not independent of LEE 10 $\Gamma(\tau^- \rightarrow K^- K^+ \pi^- \nu_\tau) / \Gamma_{\text{total}}$ and $\Gamma(\tau^- \rightarrow \pi^- \pi^+ \pi^- \nu_\tau (\text{ex. } K^0)) / \Gamma_{\text{total}}$ values.

$$\Gamma(K^- K^+ \pi^0 \nu_\tau) / \Gamma_{\text{total}} \quad \Gamma_{107} / \Gamma$$

| VALUE (units 10^{-4}) | CL% | EVTS | DOCUMENT ID | TECN | COMMENT |
|--------------------------|-----|------|-------------|------|---------|
|--------------------------|-----|------|-------------|------|---------|

0.61±0.18 OUR FIT

0.60±0.18 OUR AVERAGE

$0.55 \pm 0.14 \pm 0.12$ 48 ARMS 05 CLE3 $7.6 \text{ fb}^{-1}, E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$

$7.5 \pm 2.9 \pm 1.5$ BARATE 98 ALEP 1991–1995 LEP runs

• • • We use the following data for averages but not for fits. • • •

$3.3 \pm 1.8 \pm 0.7$ 158 ¹ RICHICHI 99 CLEO $E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$

• • • We do not use the following data for averages, fits, limits, etc. • • •

<27 95 ABBIENDI 00D OPAL 1990–1995 LEP runs

¹ Not independent of RICHICHI 99

$\Gamma(\tau^- \rightarrow K^- K^+ \pi^- \nu_\tau) / \Gamma(\tau^- \rightarrow \pi^- \pi^+ \pi^- \nu_\tau (\text{ex. } K^0))$ and BAEST 95C $\Gamma(\tau^- \rightarrow h^- h^- h^+ \nu_\tau (\text{ex. } K^0)) / \Gamma_{\text{total}}$ values.

$$\frac{\Gamma(K^- K^+ \pi^- \pi^0 \nu_\tau)/\Gamma(\pi^- \pi^+ \pi^- \pi^0 \nu_\tau (\text{ex.} K^0))}{\Gamma_{107}/\Gamma_{77}} = \Gamma_{107}/(\Gamma_{78} + 0.892\Gamma_{176} + 0.0153\Gamma_{178})$$

Γ_{107}/Γ_{77}

| <u>VALUE (%)</u> | <u>EVTS</u> | <u>DOCUMENT ID</u> | <u>TECN</u> | <u>COMMENT</u> |
|--------------------------|-------------|-----------------------|-------------|---|
| 0.14±0.04 OUR FIT | | | | |
| 0.79±0.44±0.16 | 158 | ¹ RICHICHI | 99 | CLEO $E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$ |

¹ RICHICHI 99 also quote a 95%CL upper limit of 0.0157 for this measurement.

$$\frac{\Gamma(K^- K^+ K^- \nu_\tau)/\Gamma_{\text{total}}}{\Gamma_{108}/\Gamma = 0.489\Gamma_{168}/\Gamma}$$

| <u>VALUE (units 10⁻⁵)</u> | <u>CL%</u> | <u>EVTS</u> | <u>DOCUMENT ID</u> | <u>TECN</u> | <u>COMMENT</u> |
|--------------------------------------|------------|-------------|--------------------|-------------|-------------------------------------|
| 2.2 ±0.8 OUR FIT | | | | | Error includes scale factor of 5.4. |
| 2.1 ±0.8 OUR AVERAGE | | | | | Error includes scale factor of 5.4. |

| | | | | | |
|---------------------------------|------|---------------------|----|------|--|
| $3.29 \pm 0.17^{+0.19}_{-0.20}$ | 3.2k | ¹ LEE | 10 | BELL | $666 \text{ fb}^{-1} E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$ |
| $1.58 \pm 0.13 \pm 0.12$ | 275 | ² AUBERT | 08 | BABR | $342 \text{ fb}^{-1} E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$ |

• • • We do not use the following data for averages, fits, limits, etc. • • •

| | | | | | |
|-------|----|--------|----|------|--|
| < 3.7 | 90 | BRIERE | 03 | CLE3 | $E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$ |
| < 19 | 90 | BARATE | 98 | ALEP | 1991–1995 LEP runs |

¹ See footnote to LEE 10 $\Gamma(\tau^- \rightarrow \pi^- \pi^+ \pi^- \nu_\tau (\text{ex.} K^0))/\Gamma_{\text{total}}$ measurement for correlations with other measurements. Not independent of LEE 10 $\Gamma(\tau^- \rightarrow K^- K^+ K^- \nu_\tau)/\Gamma(\tau^- \rightarrow \pi^- \pi^+ \pi^- \nu_\tau (\text{ex.} K^0))$ value.

² See footnote to AUBERT 08 $\Gamma(\tau^- \rightarrow \pi^- \pi^+ \pi^- \nu_\tau (\text{ex.} K^0))/\Gamma_{\text{total}}$ measurement for correlations with other measurements.

$$\frac{\Gamma(K^- K^+ K^- \nu_\tau)/\Gamma(\pi^- \pi^+ \pi^- \nu_\tau (\text{ex.} K^0))}{\Gamma_{108}/\Gamma_{68}}$$

| <u>VALUE (units 10⁻⁴)</u> | <u>EVTS</u> | <u>DOCUMENT ID</u> | <u>TECN</u> | <u>COMMENT</u> |
|---|-------------|--------------------|-------------|----------------|
| • • • We do not use the following data for averages, fits, limits, etc. • • • | | | | |

| | | | | | |
|---------------------------------|------|------------------|----|------|--|
| $3.90 \pm 0.02^{+0.22}_{-0.23}$ | 3.2k | ¹ LEE | 10 | BELL | $666 \text{ fb}^{-1} E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$ |
|---------------------------------|------|------------------|----|------|--|

¹ Not independent of LEE 10 $\Gamma(\tau^- \rightarrow K^- K^+ K^- \nu_\tau)/\Gamma_{\text{total}}$ and $\Gamma(\tau^- \rightarrow \pi^- \pi^+ \pi^- \nu_\tau (\text{ex.} K^0))/\Gamma_{\text{total}}$ values.

$$\frac{\Gamma(K^- K^+ K^- \nu_\tau (\text{ex. } \phi))/\Gamma_{\text{total}}}{\Gamma_{109}/\Gamma}$$

| <u>VALUE</u> | <u>CL%</u> | <u>DOCUMENT ID</u> | <u>TECN</u> | <u>COMMENT</u> |
|------------------------|------------|--------------------|-------------|---|
| $< 2.5 \times 10^{-6}$ | 90 | AUBERT | 08 | BABR $342 \text{ fb}^{-1} E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$ |

$$\frac{\Gamma(K^- K^+ K^- \pi^0 \nu_\tau)/\Gamma_{\text{total}}}{\Gamma_{110}/\Gamma}$$

| <u>VALUE</u> | <u>CL%</u> | <u>DOCUMENT ID</u> | <u>TECN</u> | <u>COMMENT</u> |
|------------------------|------------|--------------------|-------------|--|
| $< 4.8 \times 10^{-6}$ | 90 | ARMS | 05 | CLE3 $7.6 \text{ fb}^{-1}, E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$ |

$$\frac{\Gamma(\pi^- K^+ \pi^- \geq 0 \text{ neut. } \nu_\tau)/\Gamma_{\text{total}}}{\Gamma_{111}/\Gamma}$$

| <u>VALUE (%)</u> | <u>CL%</u> | <u>DOCUMENT ID</u> | <u>TECN</u> | <u>COMMENT</u> |
|------------------|------------|--------------------|-------------|--|
| < 0.25 | 95 | BAUER | 94 | TPC $E_{\text{cm}}^{\text{ee}} = 29 \text{ GeV}$ |

$$\frac{\Gamma(e^- e^- e^+ \bar{\nu}_e \nu_\tau)/\Gamma_{\text{total}}}{\Gamma_{112}/\Gamma}$$

| <u>VALUE (units 10⁻⁵)</u> | <u>EVTS</u> | <u>DOCUMENT ID</u> | <u>TECN</u> | <u>COMMENT</u> |
|--------------------------------------|-------------|--------------------|-------------|---|
| 2.8±1.4±0.4 | 5 | ALAM | 96 | CLEO $E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$ |

| $\Gamma(\mu^- e^- e^+ \bar{\nu}_\mu \nu_\tau)/\Gamma_{\text{total}}$ | Γ_{113}/Γ | | | |
|--|-----------------------|--------------------|-------------|--|
| <u>VALUE (units 10^{-5})</u> | <u>CL%</u> | <u>DOCUMENT ID</u> | <u>TECN</u> | <u>COMMENT</u> |
| <3.6 | 90 | ALAM | 96 | CLEO $E_{\text{cm}}^{ee} = 10.6 \text{ GeV}$ |

| $\Gamma(3h^- 2h^+ \geq 0 \text{ neutrals } \nu_\tau \text{ (ex. } K_S^0 \rightarrow \pi^- \pi^+ \text{) ("5-prong")})/\Gamma_{\text{total}}$ | Γ_{114}/Γ |
|--|-----------------------|
| $\Gamma_{114}/\Gamma = (\Gamma_{115} + \Gamma_{121})/\Gamma$ | |

| <u>VALUE (%)</u> | <u>EVTS</u> | <u>DOCUMENT ID</u> | <u>TECN</u> | <u>COMMENT</u> |
|------------------------------|-------------|--------------------|-------------|----------------|
| 0.099 ± 0.004 OUR FIT | | | | |

0.107 ± 0.007 OUR AVERAGE Error includes scale factor of 1.1.

| | | | | | |
|-----------------------------|-----|---------------------|-----|------|---|
| $0.170 \pm 0.022 \pm 0.026$ | | ¹ ACHARD | 01D | L3 | 1992–1995 LEP runs |
| $0.097 \pm 0.005 \pm 0.011$ | 419 | GIBAUT | 94B | CLEO | $E_{\text{cm}}^{ee} = 10.6 \text{ GeV}$ |
| 0.102 ± 0.029 | 13 | BYLSMA | 87 | HRS | $E_{\text{cm}}^{ee} = 29 \text{ GeV}$ |

• • • We use the following data for averages but not for fits. • • •

| | | | | | |
|-----------------------------|-----|-------------------------|-----|------|--------------------|
| $0.093 \pm 0.009 \pm 0.012$ | | SCHAEL | 05C | ALEP | 1991–1995 LEP runs |
| $0.115 \pm 0.013 \pm 0.006$ | 112 | ² ABREU | 01M | DLPH | 1992–1995 LEP runs |
| $0.119 \pm 0.013 \pm 0.008$ | 119 | ³ ACKERSTAFF | 99E | OPAL | 1991–1995 LEP runs |

• • • We do not use the following data for averages, fits, limits, etc. • • •

| | | | | | |
|---|----|----------|-----|------|--|
| $0.26 \pm 0.06 \pm 0.05$ | | ACTON | 92H | OPAL | $E_{\text{cm}}^{ee} = 88.2\text{--}94.2 \text{ GeV}$ |
| $0.10 \begin{array}{l} +0.05 \\ -0.04 \end{array} \pm 0.03$ | | DECAMP | 92C | ALEP | 1989–1990 LEP runs |
| $0.16 \pm 0.13 \pm 0.04$ | | BEHREND | 89B | CELL | $E_{\text{cm}}^{ee} = 14\text{--}47 \text{ GeV}$ |
| $0.3 \pm 0.1 \pm 0.2$ | | BARTEL | 85F | JADE | $E_{\text{cm}}^{ee} = 34.6 \text{ GeV}$ |
| 0.13 ± 0.04 | 10 | BELTRAMI | 85 | HRS | Repl. by BYLSMA 87 |
| $0.16 \pm 0.08 \pm 0.04$ | 4 | BURCHAT | 85 | MRK2 | $E_{\text{cm}}^{ee} = 29 \text{ GeV}$ |
| 1.0 ± 0.4 | 10 | BEHREND | 82 | CELL | Repl. by BEHREND 89B |

¹ The correlation coefficients between this measurement and the ACHARD 01D measurements of $B(\tau \rightarrow \text{"1-prong"})$ and $B(\tau \rightarrow \text{"3-prong"})$ are -0.082 and -0.19 respectively.

² The correlation coefficients between this measurement and the ABREU 01M measurements of $B(\tau \rightarrow \text{1-prong})$ and $B(\tau \rightarrow \text{3-prong})$ are -0.08 and -0.08 respectively.

³ Not independent of ACKERSTAFF 99E $B(\tau^- \rightarrow 3h^- 2h^+ \nu_\tau \text{ (ex. } K^0\text{)})$ and $B(\tau^- \rightarrow 3h^- 2h^+ \pi^0 \nu_\tau \text{ (ex. } K^0\text{)})$ measurements.

| $\Gamma(3h^- 2h^+ \nu_\tau \text{ (ex. } K^0\text{)})/\Gamma_{\text{total}}$ | $\Gamma_{115}/\Gamma = (\Gamma_{116} + \Gamma_{118} + 0.0153\Gamma_{183})/\Gamma$ | | | |
|--|---|--------------------|-------------|----------------|
| <u>VALUE (units 10^{-4})</u> | <u>EVTS</u> | <u>DOCUMENT ID</u> | <u>TECN</u> | <u>COMMENT</u> |
| 8.22 ± 0.32 OUR FIT | | | | |

8.32 ± 0.35 OUR AVERAGE

| | | | | | |
|-----------------------|-----|-----------------------|-----|------|---|
| $9.7 \pm 1.5 \pm 0.5$ | 96 | ¹ ABDALLAH | 06A | DLPH | 1992–1995 LEP runs |
| $7.2 \pm 0.9 \pm 1.2$ | 165 | ² SCHAEL | 05C | ALEP | 1991–1995 LEP runs |
| $9.1 \pm 1.4 \pm 0.6$ | 97 | ACKERSTAFF | 99E | OPAL | 1991–1995 LEP runs |
| $7.7 \pm 0.5 \pm 0.9$ | 295 | GIBAUT | 94B | CLEO | $E_{\text{cm}}^{ee} = 10.6 \text{ GeV}$ |
| $6.4 \pm 2.3 \pm 1.0$ | 12 | ALBRECHT | 88B | ARG | $E_{\text{cm}}^{ee} = 10 \text{ GeV}$ |
| 5.1 ± 2.0 | 7 | BYLSMA | 87 | HRS | $E_{\text{cm}}^{ee} = 29 \text{ GeV}$ |

• • • We use the following data for averages but not for fits. • • •

$8.56 \pm 0.05 \pm 0.42$ 34k AUBERT,B 05W BABR 232 fb^{-1} , $E_{\text{cm}}^{ee} = 10.6 \text{ GeV}$

• • • We do not use the following data for averages, fits, limits, etc. • • •

| | | | | | |
|-----------------------|----|-----------------------|----|------|---------------------|
| $8.0 \pm 1.1 \pm 1.3$ | 58 | BUSKULIC | 96 | ALEP | Repl. by SCHAEL 05C |
| 6.7 ± 3.0 | 5 | ³ BELTRAMI | 85 | HRS | Repl. by BYLSMA 87 |

¹ See footnote to ABDALLAH 06A $\Gamma(\tau^- \rightarrow h^- \nu_\tau)/\Gamma_{\text{total}}$ measurement for correlations with other measurements.

² See footnote to SCHAEEL 05C $\Gamma(\tau^- \rightarrow e^- \bar{\nu}_e \nu_\tau)/\Gamma_{\text{total}}$ measurement for correlations with other measurements.

³ The error quoted is statistical only.

$\Gamma(3\pi^- 2\pi^+ \nu_\tau (\text{ex. } K^0, \omega)) / \Gamma_{\text{total}}$

$\Gamma_{116}/\Gamma = (\Gamma_{117} + \Gamma_{171})/\Gamma$

| VALUE (units 10^{-4}) | DOCUMENT ID | TECN | COMMENT |
|--------------------------|-------------|------|---------|
|--------------------------|-------------|------|---------|

8.21 ± 0.31 OUR FIT

• • • We use the following data for averages but not for fits. • • •

$8.33 \pm 0.04 \pm 0.43$

¹ LEES 12X BABR 468 fb^{-1} $E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$

¹ Not independent of LEES 12X $\Gamma(\tau^- \rightarrow f_1(1285) \pi^- \nu_\tau \rightarrow 3\pi^- 2\pi^+ \nu_\tau)/\Gamma$ and $\Gamma(\tau^- \rightarrow 3\pi^- 2\pi^+ \nu_\tau (\text{ex. } K^0, \omega, f_1(1285))/\Gamma$ values.

$\Gamma(3\pi^- 2\pi^+ \nu_\tau (\text{ex. } K^0, \omega, f_1(1285))) / \Gamma_{\text{total}}$

Γ_{117}/Γ

| VALUE (units 10^{-4}) | EVTS | DOCUMENT ID | TECN | COMMENT |
|--------------------------|------|-------------|------|---------|
|--------------------------|------|-------------|------|---------|

7.69 ± 0.30 OUR FIT

$7.68 \pm 0.04 \pm 0.40$

69k LEES 12X BABR 468 fb^{-1} $E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$

$\Gamma(K^- 2\pi^- 2\pi^+ \nu_\tau (\text{ex. } K^0)) / \Gamma_{\text{total}}$

Γ_{118}/Γ

| VALUE (units 10^{-6}) | DOCUMENT ID | TECN | COMMENT |
|--------------------------|-------------|------|---------|
|--------------------------|-------------|------|---------|

0.6 ± 1.2 OUR FIT

$0.6 \pm 0.5 \pm 1.1$

¹ LEES 12X BABR 468 fb^{-1} $E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$

¹ LEES 12X measurement corresponds to the lower limit of $< 2.4 \times 10^{-6}$ at 90% CL. |

$\Gamma(K^+ 3\pi^- \pi^+ \nu_\tau) / \Gamma_{\text{total}}$

Γ_{119}/Γ

| VALUE | CL% | DOCUMENT ID | TECN | COMMENT |
|------------------------|-----|-------------|------|---|
| $< 5.0 \times 10^{-6}$ | 90 | LEES | 12X | BABR 468 fb^{-1} $E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$ |

$\Gamma(K^+ K^- 2\pi^- \pi^+ \nu_\tau) / \Gamma_{\text{total}}$

Γ_{120}/Γ

| VALUE | CL% | DOCUMENT ID | TECN | COMMENT |
|------------------------|-----|-------------|------|---|
| $< 4.5 \times 10^{-7}$ | 90 | LEES | 12X | BABR 468 fb^{-1} $E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$ |

$\Gamma(3h^- 2h^+ \pi^0 \nu_\tau (\text{ex. } K^0)) / \Gamma_{\text{total}}$

$\Gamma_{121}/\Gamma = (\Gamma_{122} + \Gamma_{125})/\Gamma$

| VALUE (units 10^{-4}) | EVTS | DOCUMENT ID | TECN | COMMENT |
|--------------------------|------|-------------|------|---------|
|--------------------------|------|-------------|------|---------|

1.64 ± 0.11 OUR FIT

1.74 ± 0.27 OUR AVERAGE

| | | | | |
|--|-----|---------------------------|------|--|
| $1.6 \pm 1.2 \pm 0.6$ | 13 | ¹ ABDALLAH 06A | DLPH | 1992–1995 LEP runs |
| $2.1 \pm 0.7 \pm 0.9$ | 95 | ² SCHAEEL 05C | ALEP | 1991–1995 LEP runs |
| $1.7 \pm 0.2 \pm 0.2$ | 231 | ANASTASSOV 01 | CLEO | $E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$ |
| $2.7 \pm 1.8 \pm 0.9$ | 23 | ACKERSTAFF 99E | OPAL | 1991–1995 LEP runs |
| • • • We do not use the following data for averages, fits, limits, etc. • • • | | | | |
| 1.8 $\pm 0.7 \pm 1.2$ | 18 | BUSKULIC 96 | ALEP | Repl. by SCHAEEL 05C |
| 1.9 $\pm 0.4 \pm 0.4$ | 31 | GIBAUT 94B | CLEO | Repl. by ANASTASSOV 01 |
| 5.1 ± 2.2 | 6 | BYLSMA 87 | HRS | $E_{\text{cm}}^{\text{ee}} = 29 \text{ GeV}$ |
| 6.7 ± 3.0 | 5 | ³ BELTRAMI 85 | HRS | Repl. by BYLSMA 87 |

¹ See footnote to ABDALLAH 06A $\Gamma(\tau^- \rightarrow h^- \nu_\tau)/\Gamma_{\text{total}}$ measurement for correlations with other measurements.

² SCHael 05C quote $(1.4 \pm 0.7 \pm 0.9) \times 10^{-4}$. We add 0.7×10^{-4} to remove their correction for $\tau^- \rightarrow \eta \pi^- \pi^+ \pi^- \nu_\tau \rightarrow 3\pi^- 2\pi^+ \pi^0 \nu_\tau$ and $\tau^- \rightarrow K^*(892)^- \eta \nu_\tau \rightarrow 3\pi^- 2\pi^+ \pi^0 \nu_\tau$ decays. See footnote to SCHael 05C $\Gamma(\tau^- \rightarrow e^- \bar{\nu}_e \nu_\tau)/\Gamma_{\text{total}}$ measurement for correlations with other measurements.

³ The error quoted is statistical only.

$$\Gamma(3\pi^- 2\pi^+ \pi^0 \nu_\tau (\text{ex. } K^0)) / \Gamma_{\text{total}} \quad \Gamma_{122} / \Gamma$$

$$\Gamma_{122} / \Gamma = (\Gamma_{124} + 0.2292 \Gamma_{158} + 0.892 \Gamma_{183}) / \Gamma$$

| VALUE (units 10^{-4}) | DOCUMENT ID | TECN | COMMENT |
|----------------------------|-------------|------|---------|
| 1.62 ± 0.11 OUR FIT | | | |

• • • We use the following data for averages but not for fits. • • •

$$1.65 \pm 0.05 \pm 0.09 \quad ^1 \text{LEES} \quad 12X \quad \text{BABR} \quad 468 \text{ fb}^{-1} \quad E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$$

¹ Not independent of LEES 12X measurements of $\Gamma(\tau^- \rightarrow 2\pi^- \pi^+ \omega \nu_\tau (\text{ex. } K^0)) / \Gamma$, $\Gamma(\tau^- \rightarrow \eta \pi^- \pi^+ \pi^- \nu_\tau (\text{ex. } K^0)) / \Gamma$, and $\Gamma(\tau^- \rightarrow 3\pi^- 2\pi^+ \pi^0 \nu_\tau (\text{ex. } K^0, \eta, \omega, f_1(1285))) / \Gamma$.

$$\Gamma(3\pi^- 2\pi^+ \pi^0 \nu_\tau (\text{ex. } K^0, \eta, f_1(1285))) / \Gamma_{\text{total}} \quad \Gamma_{123} / \Gamma$$

| VALUE (units 10^{-4}) | DOCUMENT ID | TECN | COMMENT |
|---------------------------|-------------------|------|---|
| 1.11 ± 0.04 ± 0.09 | ¹ LEES | 12X | BABR $468 \text{ fb}^{-1} E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$ |

¹ Not independent of LEES 12X $\Gamma(\tau^- \rightarrow 2\pi^- \pi^+ \omega \nu_\tau (\text{ex. } K^0)) / \Gamma$ and $\Gamma(\tau^- \rightarrow 3\pi^- 2\pi^+ \pi^0 \nu_\tau (\text{ex. } K^0, \eta, \omega, f_1(1285))) / \Gamma$ values.

$$\Gamma(3\pi^- 2\pi^+ \pi^0 \nu_\tau (\text{ex. } K^0, \eta, \omega, f_1(1285))) / \Gamma_{\text{total}} \quad \Gamma_{124} / \Gamma$$

| VALUE (units 10^{-4}) | EVTS | DOCUMENT ID | TECN | COMMENT |
|----------------------------|------|-------------|------|---|
| 0.38 ± 0.09 OUR FIT | | | | |
| 0.36 ± 0.03 ± 0.09 | 7.3k | LEES | 12X | BABR $468 \text{ fb}^{-1} E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$ |

$$\Gamma(K^- 2\pi^- 2\pi^+ \pi^0 \nu_\tau (\text{ex. } K^0)) / \Gamma_{\text{total}} \quad \Gamma_{125} / \Gamma$$

| VALUE (units 10^{-6}) | DOCUMENT ID | TECN | COMMENT |
|--------------------------|-------------|------|---------|
| 1.1 ± 0.6 OUR FIT | | | |

$$1.1 \pm 0.4 \pm 0.4 \quad ^1 \text{LEES} \quad 12X \quad \text{BABR} \quad 468 \text{ fb}^{-1} \quad E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$$

¹ LEES 12X measurement corresponds to the lower limit of $< 1.9 \times 10^{-6}$ at 90% CL. ■

$$\Gamma(K^+ 3\pi^- \pi^+ \pi^0 \nu_\tau) / \Gamma_{\text{total}} \quad \Gamma_{126} / \Gamma$$

| VALUE | CL% | DOCUMENT ID | TECN | COMMENT |
|----------------------|-----|-------------|------|---|
| $< 8 \times 10^{-7}$ | 90 | LEES | 12X | BABR $468 \text{ fb}^{-1} E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$ |

$$\Gamma(3h^- 2h^+ 2\pi^0 \nu_\tau) / \Gamma_{\text{total}} \quad \Gamma_{127} / \Gamma$$

| VALUE | CL% | DOCUMENT ID | TECN | COMMENT |
|------------------------|-----|-------------|------|--|
| $< 3.4 \times 10^{-6}$ | 90 | AUBERT,B 06 | BABR | $232 \text{ fb}^{-1} E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$ |

• • • We do not use the following data for averages, fits, limits, etc. • • •

$$< 1.1 \times 10^{-4} \quad 90 \quad \text{GIBAUT} \quad 94B \quad \text{CLEO} \quad E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$$

| $\Gamma((5\pi)^-\nu_\tau)/\Gamma_{\text{total}}$ | Γ_{128}/Γ |
|--|--------------------------|
| $\Gamma_{128}/\Gamma = (\Gamma_{30} + \frac{1}{2}\Gamma_{45} + \Gamma_{48} + \frac{1}{2}\Gamma_{61} + \Gamma_{85} + \Gamma_{115} + 0.5559\Gamma_{148} + 0.892\Gamma_{178})/\Gamma$ | |
| VALUE (%) | DOCUMENT ID TECN COMMENT |

0.78±0.05 OUR FIT

• • • We use the following data for averages but not for fits. • • •

| | |
|-----------------------|--|
| 0.61±0.06±0.08 | 1 GIBAUT 94B CLEO $E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$ |
|-----------------------|--|

¹ Not independent of GIBAUT 94B $B(3h^- 2h^+\nu_\tau)$, PROCARIO 93 $B(h^- 4\pi^0\nu_\tau)$, and BORTOLETTO 93 $B(2h^- h^+ 2\pi^0\nu_\tau)/B(\text{"3prong"})$ measurements. Result is corrected for η contributions.

| $\Gamma(4h^- 3h^+ \geq 0 \text{ neutrals } \nu_\tau \text{ ("7-prong")})/\Gamma_{\text{total}}$ | Γ_{129}/Γ |
|---|-----------------------|
| VALUE DOCUMENT ID TECN COMMENT | |

| | |
|----------------------------------|--|
| <3.0 × 10⁻⁷ | 90 AUBERT,B 05F BABR $232 \text{ fb}^{-1}, E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$ |
|----------------------------------|--|

• • • We do not use the following data for averages, fits, limits, etc. • • •

| | |
|-------------------------|--|
| <1.8 × 10 ⁻⁵ | 95 ACKERSTAFF 97J OPAL 1990–1995 LEP runs |
| <2.4 × 10 ⁻⁶ | 90 EDWARDS 97B CLEO $E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$ |
| <2.9 × 10 ⁻⁴ | 90 BYLSMA 87 HRS $E_{\text{cm}}^{\text{ee}} = 29 \text{ GeV}$ |

| $\Gamma(4h^- 3h^+ \nu_\tau)/\Gamma_{\text{total}}$ | Γ_{130}/Γ |
|--|-----------------------|
| VALUE DOCUMENT ID TECN COMMENT | |

| | |
|----------------------------------|--|
| <4.3 × 10⁻⁷ | 90 AUBERT,B 05F BABR $232 \text{ fb}^{-1}, E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$ |
|----------------------------------|--|

| $\Gamma(4h^- 3h^+ \pi^0 \nu_\tau)/\Gamma_{\text{total}}$ | Γ_{131}/Γ |
|--|-----------------------|
| VALUE DOCUMENT ID TECN COMMENT | |

| | |
|----------------------------------|--|
| <2.5 × 10⁻⁷ | 90 AUBERT,B 05F BABR $232 \text{ fb}^{-1}, E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$ |
|----------------------------------|--|

| $\Gamma(X^-(S=-1)\nu_\tau)/\Gamma_{\text{total}}$ | Γ_{132}/Γ |
|--|-----------------------|
| $\Gamma_{132}/\Gamma = (\Gamma_{10} + \Gamma_{16} + \Gamma_{23} + \Gamma_{28} + \Gamma_{36} + \Gamma_{41} + \Gamma_{45} + \Gamma_{61} + \Gamma_{97} + \Gamma_{103} + \Gamma_{118} + \Gamma_{125} + \Gamma_{150} + \Gamma_{152} + \Gamma_{154} + 0.8312\Gamma_{168} + \Gamma_{177})/\Gamma$ | |

| | | | |
|-----------|-------------|------|---------|
| VALUE (%) | DOCUMENT ID | TECN | COMMENT |
|-----------|-------------|------|---------|

2.92±0.04 OUR FIT

• • • We use the following data for averages but not for fits. • • •

| | |
|------------------|--------------------------------------|
| 2.87±0.12 | 1 BARATE 99R ALEP 1991–1995 LEP runs |
|------------------|--------------------------------------|

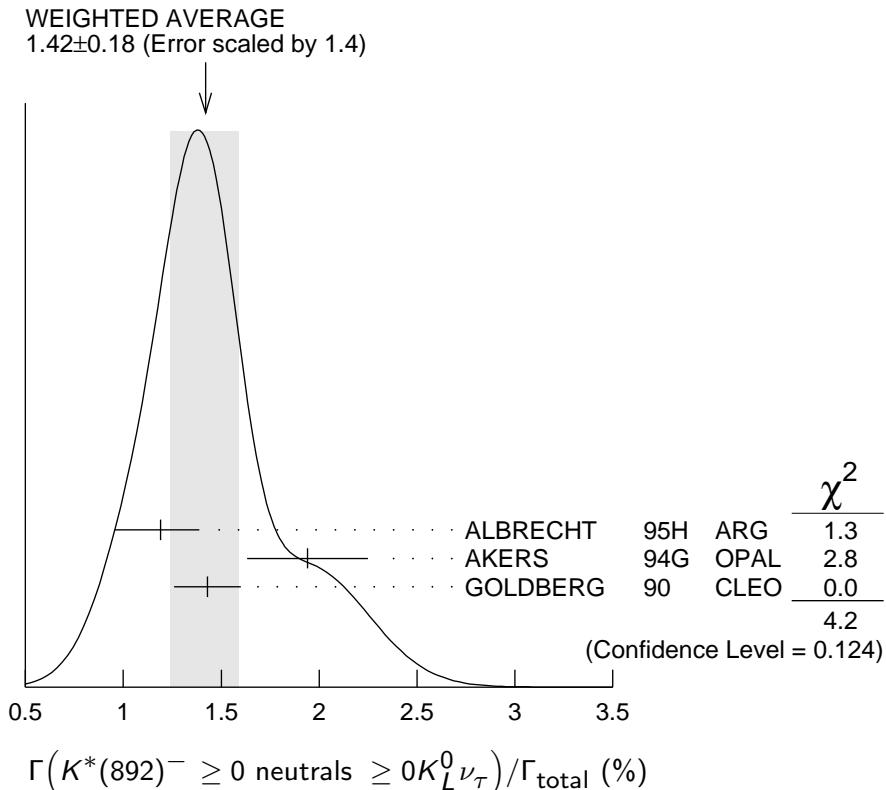
¹ BARATE 99R perform a combined analysis of all ALEPH LEP 1 data on τ branching fraction measurements for decay modes having total strangeness equal to -1 .

| $\Gamma(K^*(892)^-\geq 0 \text{ neutrals} \geq 0 K_L^0 \nu_\tau)/\Gamma_{\text{total}}$ | Γ_{133}/Γ |
|---|-----------------------|
| VALUE (%) EVTS DOCUMENT ID TECN COMMENT | |

1.42±0.18 OUR AVERAGE Error includes scale factor of 1.4. See the ideogram below.

| | |
|---|--|
| 1.19±0.15 ^{+0.13} _{-0.18} | 104 ALBRECHT 95H ARG $E_{\text{cm}}^{\text{ee}} = 9.4\text{--}10.6 \text{ GeV}$ |
| 1.94±0.27±0.15 | 74 ¹ AKERS 94G OPAL $E_{\text{cm}}^{\text{ee}} = 88\text{--}94 \text{ GeV}$ |
| 1.43±0.11±0.13 | 475 ² GOLDBERG 90 CLEO $E_{\text{cm}}^{\text{ee}} = 9.4\text{--}10.9 \text{ GeV}$ |

- ¹ AKERS 94G reject events in which a K_S^0 accompanies the $K^*(892)^-$. We do not correct for them.
² GOLDBERG 90 estimates that 10% of observed $K^*(892)$ are accompanied by a π^0 .

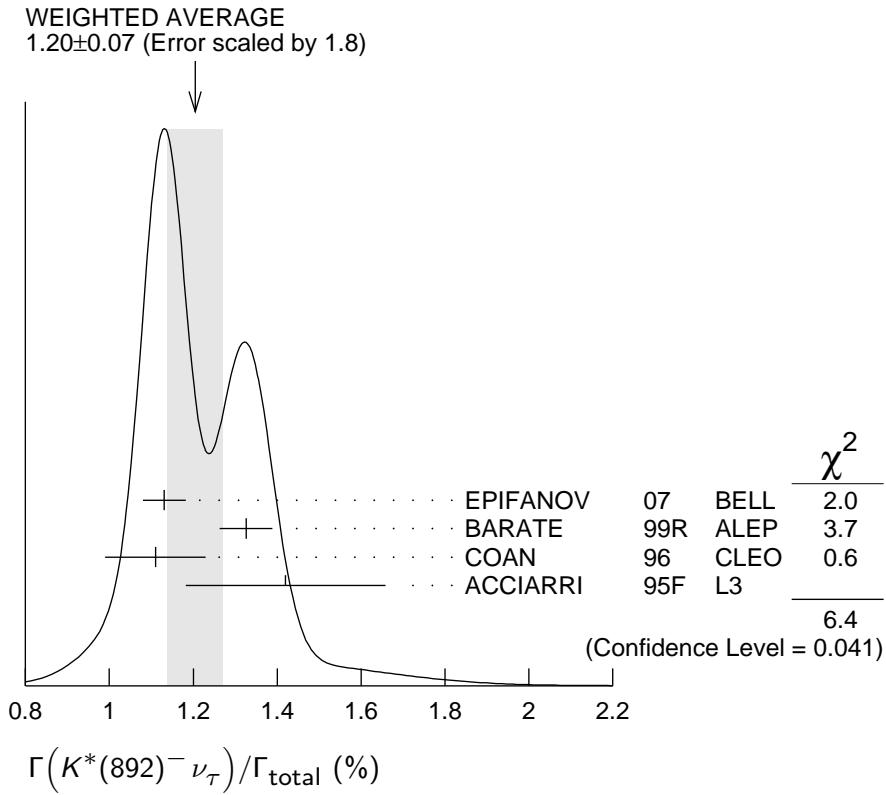


$\Gamma(K^*(892)^- / \nu_\tau) / \Gamma_{\text{total}}$ Γ_{134} / Γ

| VALUE (%) | EVTS | DOCUMENT ID | TECN | COMMENT |
|---|------|-----------------|--|---|
| 1.20 ± 0.07 OUR AVERAGE | | | | Error includes scale factor of 1.8. See the ideogram below. |
| 1.131 ± 0.006 ± 0.051 | 49k | 1 EPIFANOV 07 | BELL 351 fb ⁻¹ | $E_{\text{cm}}^{ee} = 10.6 \text{ GeV}$ |
| 1.326 ± 0.063 | | BARATE 99R | ALEP 1991–1995 LEP runs | |
| 1.11 ± 0.12 | | 2 COAN 96 | CLEO $E_{\text{cm}}^{ee} \approx 10.6 \text{ GeV}$ | |
| 1.42 ± 0.22 ± 0.09 | | 3 ACCIARRI 95F | L3 1991–1993 LEP runs | |
| • • • We do not use the following data for averages, fits, limits, etc. • • • | | | | |
| 1.39 ± 0.09 ± 0.10 | | 4 BUSKULIC 96 | ALEP | Repl. by BARATE 99R |
| 1.45 ± 0.13 ± 0.11 | 273 | 5 BUSKULIC 94F | ALEP | Repl. by BUSKULIC 96 |
| 1.23 ± 0.21 ± 0.11 | 54 | 6 ALBRECHT 88L | ARG | $E_{\text{cm}}^{ee} = 10 \text{ GeV}$ |
| 1.9 ± 0.3 ± 0.4 | 44 | 7 TSCHIRHART 88 | HRS | $E_{\text{cm}}^{ee} = 29 \text{ GeV}$ |
| 1.5 ± 0.4 ± 0.4 | 15 | 8 AIHARA 87C | TPC | $E_{\text{cm}}^{ee} = 29 \text{ GeV}$ |
| 1.3 ± 0.3 ± 0.3 | 31 | YELTON 86 | MRK2 | $E_{\text{cm}}^{ee} = 29 \text{ GeV}$ |
| 1.7 ± 0.7 | 11 | DORFAN 81 | MRK2 | $E_{\text{cm}}^{ee} = 4.2\text{--}6.7 \text{ GeV}$ |

¹ EPIFANOV 07 quote $B(\tau^- \rightarrow K^*(892)^- \nu_\tau) B(K^*(892)^- \rightarrow K_S^0 \pi^-) = (3.77 \pm 0.02(\text{stat}) \pm 0.12(\text{syst}) \pm 0.12(\text{mod})) \times 10^{-3}$. We add the systematic and model uncertainties in quadrature and divide by $B(K^*(892)^- \rightarrow K_S^0 \pi^-) = 0.3333$.

- ² Not independent of COAN 96 $B(\pi^-\bar{K}^0\nu_\tau)$ and BATTLE 94 $B(K^-\pi^0\nu_\tau)$ measurements. $K\pi$ final states are consistent with and assumed to originate from $K^*(892)^-$ production.
- ³ This result is obtained from their $B(\pi^-\bar{K}^0\nu_\tau)$ assuming all those decays originate in $K^*(892)^-$ decays.
- ⁴ Not independent of BUSKULIC 96 $B(\pi^-\bar{K}^0\nu_\tau)$ and $B(K^-\pi^0\nu_\tau)$ measurements.
- ⁵ BUSKULIC 94F obtain this result from BUSKULIC 94F $B(\bar{K}^0\pi^-\nu_\tau)$ and BUSKULIC 94E $B(K^-\pi^0\nu_\tau)$ assuming all of those decays originate in $K^*(892)^-$ decays.
- ⁶ The authors divide by $\Gamma_2/\Gamma = 0.865$ to obtain this result.
- ⁷ Not independent of TSCHIRHART 88 $\Gamma(\tau^- \rightarrow h^-\bar{K}^0 \geq 0 \text{ neutrals} \geq 0 K_L^0\nu_\tau) / \Gamma$.
- ⁸ Decay π^- identified in this experiment, is assumed in the others.



$$\Gamma(K^*(892)^-\nu_\tau)/\Gamma(\pi^-\pi^0\nu_\tau)$$

$$\Gamma_{134}/\Gamma_{14}$$

| VALUE | DOCUMENT ID | TECN | COMMENT |
|--------------------|-------------|------|-----------------|
| 0.075±0.027 | 1 ABREU 94K | DLPH | LEP 1992 Z data |

¹ ABREU 94K quote $B(\tau^- \rightarrow K^*(892)^-\nu_\tau)B(K^*(892)^-\rightarrow K^-\pi^0)/B(\tau^- \rightarrow \rho^-\nu_\tau) = 0.025 \pm 0.009$. We divide by $B(K^*(892)^-\rightarrow K^-\pi^0) = 0.333$ to obtain this result.

$$\Gamma(K^*(892)^-\nu_\tau \rightarrow \pi^-\bar{K}^0\nu_\tau)/\Gamma(\pi^-\bar{K}^0\nu_\tau)$$

$$\Gamma_{135}/\Gamma_{36}$$

| VALUE | EVTS | DOCUMENT ID | TECN | COMMENT |
|--------------------|------|-------------|------|---|
| 0.933±0.027 | 49k | EPIFANOV 07 | BELL | $351 \text{ fb}^{-1} E_{\text{cm}}^{ee} = 10.6 \text{ GeV}$ |

$$\Gamma(K^*(892)^0 K^- \geq 0 \text{ neutrals} \nu_\tau)/\Gamma_{\text{total}}$$

$$\Gamma_{136}/\Gamma$$

| VALUE (%) | EVTS | DOCUMENT ID | TECN | COMMENT |
|-----------------------|------|-------------|------|---|
| 0.32±0.08±0.12 | 119 | GOLDBERG 90 | CLEO | $E_{\text{cm}}^{ee} = 9.4\text{--}10.9 \text{ GeV}$ |

| $\Gamma(K^*(892)^0 K^- \nu_\tau)/\Gamma_{\text{total}}$ | | Γ_{137}/Γ | | | |
|---|------|-----------------------|------|---|--|
| VALUE (%) | EVTS | DOCUMENT ID | TECN | COMMENT | |
| 0.21 ± 0.04 OUR AVERAGE | | | | | |
| 0.21 ± 0.048 | | 1 BARATE 98 | ALEP | 1991–1995 LEP runs | |
| 0.20 ± 0.05 ± 0.04 | 47 | ALBRECHT 95H | ARG | $E_{\text{cm}}^{ee} = 9.4\text{--}10.6 \text{ GeV}$ | |
| 1 BARATE 98 measure the $K^- (\rho^0 \rightarrow \pi^+ \pi^-)$ fraction in $\tau^- \rightarrow K^- \pi^+ \pi^- \nu_\tau$ decays to be $(35 \pm 11)\%$ and derive this result from their measurement of $\Gamma(\tau^- \rightarrow K^- \pi^+ \pi^- \nu_\tau)/\Gamma_{\text{total}}$ assuming the intermediate states are all $K^- \rho$ and $K^- K^*(892)^0$. | | | | | |

| $\Gamma(\bar{K}^*(892)^0 \pi^- \geq 0 \text{ neutrals } \nu_\tau)/\Gamma_{\text{total}}$ | | Γ_{138}/Γ | | | |
|--|------|-----------------------|------|---|--|
| VALUE (%) | EVTS | DOCUMENT ID | TECN | COMMENT | |
| 0.38 ± 0.11 ± 0.13 | | | | | |
| 0.38 ± 0.11 ± 0.13 | 105 | GOLDBERG 90 | CLEO | $E_{\text{cm}}^{ee} = 9.4\text{--}10.9 \text{ GeV}$ | |

| $\Gamma(\bar{K}^*(892)^0 \pi^- \nu_\tau)/\Gamma_{\text{total}}$ | | Γ_{139}/Γ | | | |
|--|------|-----------------------|------|---|--|
| VALUE (%) | EVTS | DOCUMENT ID | TECN | COMMENT | |
| 0.22 ± 0.05 OUR AVERAGE | | | | | |
| 0.209 ± 0.058 | | 1 BARATE 98 | ALEP | 1991–1995 LEP runs | |
| 0.25 ± 0.10 ± 0.05 | 27 | ALBRECHT 95H | ARG | $E_{\text{cm}}^{ee} = 9.4\text{--}10.6 \text{ GeV}$ | |
| 1 BARATE 98 measure the $K^- K^*(892)^0$ fraction in $\tau^- \rightarrow K^- K^+ \pi^- \nu_\tau$ decays to be $(87 \pm 13)\%$ and derive this result from their measurement of $\Gamma(\tau^- \rightarrow K^- K^+ \pi^- \nu_\tau)/\Gamma_{\text{total}}$. | | | | | |

| $\Gamma((\bar{K}^*(892)\pi)^- \nu_\tau \rightarrow \pi^- \bar{K}^0 \pi^0 \nu_\tau)/\Gamma_{\text{total}}$ | | Γ_{140}/Γ | | | |
|---|------|-----------------------|------|--------------------|--|
| VALUE (%) | EVTS | DOCUMENT ID | TECN | COMMENT | |
| 0.10 ± 0.04 OUR AVERAGE | | | | | |
| 0.097 ± 0.044 ± 0.036 | | 1 BARATE 99K | ALEP | 1991–1995 LEP runs | |
| 0.106 ± 0.037 ± 0.032 | | 2 BARATE 98E | ALEP | 1991–1995 LEP runs | |
| 1 BARATE 99K measure K^0 's by detecting K_L^0 's in their hadron calorimeter. They determine the $\bar{K}^0 \rho^-$ fraction in $\tau^- \rightarrow \pi^- \bar{K}^0 \pi^0 \nu_\tau$ decays to be $(0.72 \pm 0.12 \pm 0.10)$ and multiply their $B(\pi^- \bar{K}^0 \pi^0 \nu_\tau)$ measurement by one minus this fraction to obtain the quoted result. | | | | | |
| 2 BARATE 98E reconstruct K^0 's using $K_S^0 \rightarrow \pi^+ \pi^-$ decays. They determine the $\bar{K}^0 \rho^-$ fraction in $\tau^- \rightarrow \pi^- \bar{K}^0 \pi^0 \nu_\tau$ decays to be $(0.64 \pm 0.09 \pm 0.10)$ and multiply their $B(\pi^- \bar{K}^0 \pi^0 \nu_\tau)$ measurement by one minus this fraction to obtain the quoted result. | | | | | |

| $\Gamma(K_1(1270)^- \nu_\tau)/\Gamma_{\text{total}}$ | | Γ_{141}/Γ | | | |
|--|------|-----------------------|------|---------------------------------------|--|
| VALUE (%) | EVTS | DOCUMENT ID | TECN | COMMENT | |
| 0.47 ± 0.11 OUR AVERAGE | | | | | |
| 0.48 ± 0.11 | | BARATE 99R | ALEP | 1991–1995 LEP runs | |
| 0.41 ± 0.41 ± 0.35 ± 0.10 | 5 | 1 BAUER 94 | TPC | $E_{\text{cm}}^{ee} = 29 \text{ GeV}$ | |

1 We multiply 0.41% by 0.25, the relative systematic error quoted by BAUER 94, to obtain the systematic error.

| $\Gamma(K_1(1400)^-\nu_\tau)/\Gamma_{\text{total}}$ | Γ_{142}/Γ | | | |
|---|-----------------------|-------------------------------------|-------------|--|
| <u>VALUE (%)</u> | <u>EVTS</u> | <u>DOCUMENT ID</u> | <u>TECN</u> | <u>COMMENT</u> |
| 0.17±0.26 OUR AVERAGE | | Error includes scale factor of 1.7. | | |
| 0.05±0.17 | | BARATE 99R | ALEP | 1991–1995 LEP runs |
| $0.76^{+0.40}_{-0.33} \pm 0.20$ | 11 | ¹ BAUER 94 | TPC | $E_{\text{cm}}^{\text{ee}} = 29 \text{ GeV}$ |

¹ We multiply 0.76% by 0.25, the relative systematic error quoted by BAUER 94, to obtain the systematic error.

| $[\Gamma(K_1(1270)^-\nu_\tau) + \Gamma(K_1(1400)^-\nu_\tau)]/\Gamma_{\text{total}}$ | $(\Gamma_{141}+\Gamma_{142})/\Gamma$ | | | |
|---|--------------------------------------|-----------------------|-------------|--|
| <u>VALUE (%)</u> | <u>EVTS</u> | <u>DOCUMENT ID</u> | <u>TECN</u> | <u>COMMENT</u> |
| 1.17^{+0.41}_{-0.37}±0.29 | 16 | ¹ BAUER 94 | TPC | $E_{\text{cm}}^{\text{ee}} = 29 \text{ GeV}$ |

¹ We multiply 1.17% by 0.25, the relative systematic error quoted by BAUER 94, to obtain the systematic error. Not independent of BAUER 94 $B(K_1(1270)^-\nu_\tau)$ and BAUER 94 $B(K_1(1400)^-\nu_\tau)$ measurements.

| $\Gamma(K_1(1270)^-\nu_\tau)/[\Gamma(K_1(1270)^-\nu_\tau) + \Gamma(K_1(1400)^-\nu_\tau)]$ | $\Gamma_{141}/(\Gamma_{141}+\Gamma_{142})$ | | |
|---|--|-------------|----------------|
| <u>VALUE</u> | <u>DOCUMENT ID</u> | <u>TECN</u> | <u>COMMENT</u> |
| 0.69±0.15 OUR AVERAGE | | | |

| | | | | |
|----------------|---------------------------|-----|------|--|
| 0.71±0.16±0.11 | ¹ ABBIENDI 00D | 00D | OPAL | 1990–1995 LEP runs |
| 0.66±0.19±0.13 | ² ASNER 00B | 00B | CLEO | $E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$ |

¹ ABBIENDI 00D assume the resonance structure of $\tau^- \rightarrow K^-\pi^+\pi^-\nu_\tau$ decays is dominated by the $K_1(1270)^-$ and $K_1(1400)^-$ resonances.

² ASNER 00B assume the resonance structure of $\tau^- \rightarrow K^-\pi^+\pi^-\nu_\tau$ (ex. K^0) decays is dominated by $K_1(1270)^-$ and $K_1(1400)^-$ resonances.

| $\Gamma(K^*(1410)^-\nu_\tau)/\Gamma_{\text{total}}$ | Γ_{143}/Γ | | |
|---|-----------------------|-------------|--------------------|
| <u>VALUE (units 10^{-3})</u> | <u>DOCUMENT ID</u> | <u>TECN</u> | <u>COMMENT</u> |
| 1.5^{+1.4}_{-1.0} | BARATE 99R | ALEP | 1991–1995 LEP runs |

| $\Gamma(K_0^*(1430)^-\nu_\tau)/\Gamma_{\text{total}}$ | Γ_{144}/Γ | | | |
|---|-----------------------|--------------------|-------------|--------------------|
| <u>VALUE (units 10^{-3})</u> | <u>CL%</u> | <u>DOCUMENT ID</u> | <u>TECN</u> | <u>COMMENT</u> |
| <0.5 | 95 | BARATE 99R | ALEP | 1991–1995 LEP runs |

| $\Gamma(K_2^*(1430)^-\nu_\tau)/\Gamma_{\text{total}}$ | Γ_{145}/Γ | | | | |
|---|-----------------------|-------------|--------------------|-------------|--|
| <u>VALUE (%)</u> | <u>CL%</u> | <u>EVTS</u> | <u>DOCUMENT ID</u> | <u>TECN</u> | <u>COMMENT</u> |
| <0.3 | 95 | | TSCHIRHART 88 | HRS | $E_{\text{cm}}^{\text{ee}} = 29 \text{ GeV}$ |

• • • We do not use the following data for averages, fits, limits, etc. • • •

| | | | | | |
|-------|----|---------------------------|-----------|--------------------|---|
| <0.33 | 95 | ¹ ACCIARRI 95F | L3 | 1991–1993 LEP runs | |
| <0.9 | 95 | 0 | DORFAN 81 | MRK2 | $E_{\text{cm}}^{\text{ee}} = 4.2\text{--}6.7 \text{ GeV}$ |

¹ ACCIARRI 95F quote $B(\tau^- \rightarrow K^*(1430)^- \rightarrow \pi^-\bar{K}^0\nu_\tau) < 0.11\%$. We divide by $B(K^*(1430)^- \rightarrow \pi^-\bar{K}^0) = 0.33$ to obtain the limit shown.

| $\Gamma(a_0(980)^-\geq 0 \text{ neutrals } \nu_\tau)/\Gamma_{\text{total}} \times B(a_0(980)\rightarrow K^0\bar{K}^-)$ | $\Gamma_{146}/\Gamma \times B$ | | | |
|--|--------------------------------|--------------------|-------------|--|
| <u>VALUE (units 10^{-4})</u> | <u>CL%</u> | <u>DOCUMENT ID</u> | <u>TECN</u> | <u>COMMENT</u> |
| <2.8 | 90 | GOLDBERG 90 | CLEO | $E_{\text{cm}}^{\text{ee}} = 9.4\text{--}10.9 \text{ GeV}$ |

$\Gamma(\eta\pi^-\nu_\tau)/\Gamma_{\text{total}}$ Γ_{147}/Γ

| VALUE (units 10^{-4}) | CL% | EVTS | DOCUMENT ID | TECN | COMMENT |
|--------------------------|-----|------|------------------------------|------|--|
| < 0.99 | 95 | | ¹ DEL-AMO-SA..11E | BABR | $470 \text{ fb}^{-1} E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$ |

• • • We do not use the following data for averages, fits, limits, etc. • • •

| | | | | | |
|-----------------------|----|---|----------|-----|--|
| < 6.2 | 95 | | BUSKULIC | 97C | ALEP 1991–1994 LEP runs |
| < 1.4 | 95 | 0 | BARTEL | 96 | CLEO $E_{\text{cm}}^{\text{ee}} \approx 10.6 \text{ GeV}$ |
| < 3.4 | 95 | | ARTUSO | 92 | CLEO $E_{\text{cm}}^{\text{ee}} \approx 10.6 \text{ GeV}$ |
| < 90 | 95 | | ALBRECHT | 88M | ARG $E_{\text{cm}}^{\text{ee}} \approx 10 \text{ GeV}$ |
| <140 | 90 | | BEHREND | 88 | CELL $E_{\text{cm}}^{\text{ee}} = 14\text{--}46.8 \text{ GeV}$ |
| <180 | 95 | | BARINGER | 87 | CLEO $E_{\text{cm}}^{\text{ee}} = 10.5 \text{ GeV}$ |
| <250 | 90 | 0 | COFFMAN | 87 | MRK3 $E_{\text{cm}}^{\text{ee}} = 3.77 \text{ GeV}$ |
| 510 $\pm 100 \pm 120$ | 65 | | DERRICK | 87 | HRS $E_{\text{cm}}^{\text{ee}} = 29 \text{ GeV}$ |
| <100 | 95 | | GAN | 87B | MRK2 $E_{\text{cm}}^{\text{ee}} = 29 \text{ GeV}$ |

¹ DEL-AMO-SANCHEZ 11E also quote $B(\tau^- \rightarrow \eta\pi^-\nu_\tau) = (3.4 \pm 3.4 \pm 2.1) \times 10^{-5}$.

 $\Gamma(\eta\pi^-\pi^0\nu_\tau)/\Gamma_{\text{total}}$ Γ_{148}/Γ

| VALUE (units 10^{-3}) | CL% | EVTS | DOCUMENT ID | TECN | COMMENT |
|---|-----|------|-------------|------|---------|
| 1.39 ± 0.07 OUR FIT | | | | | |

1.38 ± 0.09 OUR AVERAGE Error includes scale factor of 1.2.

| | | | | | |
|---|------|------------------|-----|------|--|
| 1.35 $\pm 0.03 \pm 0.07$ | 6.0k | INAMI | 09 | BELL | $490 \text{ fb}^{-1} E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$ |
| 1.8 $\pm 0.4 \pm 0.2$ | | BUSKULIC | 97C | ALEP | 1991–1994 LEP runs |
| 1.7 $\pm 0.2 \pm 0.2$ | 125 | ARTUSO | 92 | CLEO | $E_{\text{cm}}^{\text{ee}} \approx 10.6 \text{ GeV}$ |
| • • • We do not use the following data for averages, fits, limits, etc. • • • | | | | | |
| < 11.0 | 95 | ALBRECHT | 88M | ARG | $E_{\text{cm}}^{\text{ee}} \approx 10 \text{ GeV}$ |
| < 21.0 | 95 | BARINGER | 87 | CLEO | $E_{\text{cm}}^{\text{ee}} = 10.5 \text{ GeV}$ |
| 42.0 $\pm 7.0 \pm 16.0$ | | ¹ GAN | 87 | MRK2 | $E_{\text{cm}}^{\text{ee}} = 29 \text{ GeV}$ |

¹ Highly correlated with GAN 87 $\Gamma(\pi^- 3\pi^0\nu_\tau)/\Gamma(\text{total})$ value.

 $\Gamma(\eta\pi^-\pi^0\pi^0\nu_\tau)/\Gamma_{\text{total}}$ Γ_{149}/Γ

| VALUE (units 10^{-4}) | CL% | EVTS | DOCUMENT ID | TECN | COMMENT |
|---|-----|------|-------------|------|---------|
| 1.9 ± 0.4 OUR FIT | | | | | |
| 1.81 ± 0.31 OUR AVERAGE | | | | | |

| | | | | | |
|--|-----|------|-----|------|--|
| 2.01 $\pm 0.34 \pm 0.22$ | 381 | LEES | 12X | BABR | $468 \text{ fb}^{-1} E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$ |
| • • • We use the following data for averages but not for fits. • • • | | | | | |

1.5 ± 0.5 30 ¹ ANASTASSOV 01 CLEO $E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$

• • • We do not use the following data for averages, fits, limits, etc. • • •

| | | | | | |
|-----------------------|----|-----------------------|-----|------|--|
| 1.4 $\pm 0.6 \pm 0.3$ | 15 | ² BERGFELD | 97 | CLEO | Repl. by ANAS-TASSOV 01 |
| < 4.3 | 95 | ARTUSO | 92 | CLEO | $E_{\text{cm}}^{\text{ee}} \approx 10.6 \text{ GeV}$ |
| <120 | 95 | ALBRECHT | 88M | ARG | $E_{\text{cm}}^{\text{ee}} \approx 10 \text{ GeV}$ |

¹ Weighted average of BERGFELD 97 and ANASTASSOV 01 value of $(1.5 \pm 0.6 \pm 0.3) \times 10^{-4}$ obtained using η 's reconstructed from $\eta \rightarrow \pi^+\pi^-\pi^0$ decays.

² BERGFELD 97 reconstruct η 's using $\eta \rightarrow \gamma\gamma$ decays.

$\Gamma(\eta K^- \nu_\tau)/\Gamma_{\text{total}}$ Γ_{150}/Γ

| <u>VALUE (units 10^{-4})</u> | <u>CL%</u> | <u>EVTS</u> | <u>DOCUMENT ID</u> | <u>TECN</u> | <u>COMMENT</u> |
|---|------------|------------------|--------------------|-----------------------|--|
| 1.55 ± 0.08 OUR FIT | | | | | |
| 1.54 ± 0.08 OUR AVERAGE | | | | | |
| 1.42 $\pm 0.11 \pm 0.07$ | 690 | DEL-AMO-SA...11E | BABR | 470 fb^{-1} | $E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$ |
| 1.58 $\pm 0.05 \pm 0.09$ | 1.6k | INAMI | 09 | BELL | 490 fb^{-1} |
| $2.9^{+1.3}_{-1.2} \pm 0.7$ | | BUSKULIC | 97C | ALEP | 1991–1994 LEP runs |
| $2.6 \pm 0.5 \pm 0.5$ | 85 | BARTEL | 96 | CLEO | $E_{\text{cm}}^{\text{ee}} \approx 10.6 \text{ GeV}$ |
| $\bullet \bullet \bullet$ We do not use the following data for averages, fits, limits, etc. $\bullet \bullet \bullet$ | | | | | |
| < 4.7 | 95 | ARTUSO | 92 | CLEO | $E_{\text{cm}}^{\text{ee}} \approx 10.6 \text{ GeV}$ |

 $\Gamma(\eta K^*(892)^- \nu_\tau)/\Gamma_{\text{total}}$ Γ_{151}/Γ

| <u>VALUE (units 10^{-4})</u> | <u>EVTS</u> | <u>DOCUMENT ID</u> | <u>TECN</u> | <u>COMMENT</u> |
|---|-------------|--------------------|-------------|----------------|
| 1.38 ± 0.15 OUR AVERAGE | | | | |
| 1.34 $\pm 0.12 \pm 0.09$ | 245 | ¹ INAMI | 09 | BELL |
| 2.90 $\pm 0.80 \pm 0.42$ | 25 | BISHAI | 99 | CLEO |

¹ Not independent of INAMI 09 $B(\tau^- \rightarrow \eta K^- \pi^0 \nu_\tau)$ and $B(\tau^- \rightarrow \eta \bar{K}^0 \pi^- \nu_\tau)$ values.

 $\Gamma(\eta K^- \pi^0 \nu_\tau)/\Gamma_{\text{total}}$ Γ_{152}/Γ

| <u>VALUE (units 10^{-4})</u> | <u>EVTS</u> | <u>DOCUMENT ID</u> | <u>TECN</u> | <u>COMMENT</u> |
|---|-------------|--------------------|-------------|----------------|
| 0.48 ± 0.12 OUR FIT | | | | |
| 0.48 ± 0.12 OUR AVERAGE | | | | |
| 0.46 $\pm 0.11 \pm 0.04$ | 270 | INAMI | 09 | BELL |
| 1.77 $\pm 0.56 \pm 0.71$ | 36 | BISHAI | 99 | CLEO |

 $\Gamma(\eta K^- \pi^0 (\text{non-}K^*(892)) \nu_\tau)/\Gamma_{\text{total}}$ Γ_{153}/Γ

| <u>VALUE</u> | <u>CL%</u> | <u>DOCUMENT ID</u> | <u>TECN</u> | <u>COMMENT</u> |
|------------------------|------------|--------------------|-------------|----------------|
| $< 3.5 \times 10^{-5}$ | 90 | INAMI | 09 | BELL |

 $\Gamma(\eta \bar{K}^0 \pi^- \nu_\tau)/\Gamma_{\text{total}}$ Γ_{154}/Γ

| <u>VALUE (units 10^{-4})</u> | <u>EVTS</u> | <u>DOCUMENT ID</u> | <u>TECN</u> | <u>COMMENT</u> |
|---|-------------|---------------------|-------------|----------------|
| 0.94 ± 0.15 OUR FIT | | | | |
| 0.93 ± 0.15 OUR AVERAGE | | | | |
| 0.88 $\pm 0.14 \pm 0.06$ | 161 | ¹ INAMI | 09 | BELL |
| 2.20 $\pm 0.70 \pm 0.22$ | 15 | ² BISHAI | 99 | CLEO |

¹ We multiply the INAMI 09 measurement $B(\tau^- \rightarrow \eta K_S^0 \pi^- \nu_\tau) = (0.44 \pm 0.07 \pm 0.03) \times 10^{-4}$ by 2 to obtain the listed value.

² We multiply the BISHAI 99 measurement $B(\tau^- \rightarrow \eta K_S^0 \pi^- \nu_\tau) = (1.10 \pm 0.35 \pm 0.11) \times 10^{-4}$ by 2 to obtain the listed value.

 $\Gamma(\eta \bar{K}^0 \pi^- \pi^0 \nu_\tau)/\Gamma_{\text{total}}$ Γ_{155}/Γ

| <u>VALUE</u> | <u>CL%</u> | <u>DOCUMENT ID</u> | <u>TECN</u> | <u>COMMENT</u> |
|------------------------|------------|--------------------|-------------|----------------|
| $< 5.0 \times 10^{-5}$ | 90 | ¹ INAMI | 09 | BELL |

¹ We multiply the INAMI 09 measurement $B(\tau^- \rightarrow \eta K_S^0 \pi^- \pi^0 \nu_\tau) < 2.5 \times 10^{-5}$ by 2 to obtain the listed value.

| $\Gamma(\eta K^- K^0 \nu_\tau)/\Gamma_{\text{total}}$ | Γ_{156}/Γ | | | |
|---|-----------------------|--------------------|-------------|---|
| <u>VALUE</u> | <u>CL%</u> | <u>DOCUMENT ID</u> | <u>TECN</u> | <u>COMMENT</u> |
| $<9.0 \times 10^{-6}$ | 90 | 1 INAMI | 09 | BELL $490 \text{ fb}^{-1} E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$ |

¹ We multiply the INAMI 09 measurement $B(\tau^- \rightarrow \eta K^- K_S^0 \nu_\tau) < 4.5 \times 10^{-6}$ by 2 to obtain the listed value.

| $\Gamma(\eta \pi^+ \pi^- \pi^- \geq 0 \text{ neutrals } \nu_\tau)/\Gamma_{\text{total}}$ | Γ_{157}/Γ | | | |
|--|-----------------------|--------------------|-------------|--|
| <u>VALUE (%)</u> | <u>CL%</u> | <u>DOCUMENT ID</u> | <u>TECN</u> | <u>COMMENT</u> |
| <0.3 | 90 | ABACHI | 87B HRS | $E_{\text{cm}}^{\text{ee}} = 29 \text{ GeV}$ |

| $\Gamma(\eta \pi^- \pi^+ \pi^- \nu_\tau (\text{ex. } K^0))/\Gamma_{\text{total}}$ | Γ_{158}/Γ | | | |
|---|-----------------------|--------------------|-------------|----------------|
| <u>VALUE (units 10^{-4})</u> | <u>EVTS</u> | <u>DOCUMENT ID</u> | <u>TECN</u> | <u>COMMENT</u> |

2.19±0.13 OUR FIT

2.23±0.12 OUR AVERAGE

| | | | | |
|--------------------------|------|-------------------|----------|--------------------------------------|
| $2.10 \pm 0.09 \pm 0.13$ | 2.9k | ¹ LEES | 12x BABR | $\eta \rightarrow \gamma\gamma$ |
| $2.37 \pm 0.12 \pm 0.18$ | 1.4k | ¹ LEES | 12x BABR | $\eta \rightarrow \pi^+ \pi^- \pi^0$ |
| $2.54 \pm 0.27 \pm 0.25$ | 315 | ¹ LEES | 12x BABR | $\eta \rightarrow 3\pi^0$ |

• • • We use the following data for averages but not for fits. • • •

2.3 ± 0.5 170 ² ANASTASSOV 01 CLEO $E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$

• • • We do not use the following data for averages, fits, limits, etc. • • •

1.60 ± 0.05 ± 0.11 1.8 k AUBERT 08AE BABR Repl. by LEES 12x

3.4 $^{+0.6}_{-0.5}$ ± 0.6 89 ³ BERGFELD 97 CLEO Repl. by ANASTASSOV 01

¹ LEES 12x uses 468 fb^{-1} of data taken at $E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$. It gives the average of the three measurements listed here as $(2.25 \pm 0.07 \pm 0.12) \times 10^{-4}$.

² Weighted average of BERGFELD 97 and ANASTASSOV 01 measurements using η 's reconstructed from $\eta \rightarrow \pi^+ \pi^- \pi^0$ and $\eta \rightarrow 3\pi^0$ decays.

³ BERGFELD 97 reconstruct η 's using $\eta \rightarrow \gamma\gamma$ and $\eta \rightarrow 3\pi^0$ decays.

| $\Gamma(\eta \pi^- \pi^+ \pi^- \nu_\tau (\text{ex. } K^0, f_1(1285)))/\Gamma_{\text{total}}$ | Γ_{159}/Γ | | |
|--|-----------------------|-------------|--|
| <u>VALUE (units 10^{-4})</u> | <u>DOCUMENT ID</u> | <u>TECN</u> | <u>COMMENT</u> |
| $0.99 \pm 0.09 \pm 0.13$ | ¹ LEES | 12x BABR | $468 \text{ fb}^{-1} E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$ |

¹ LEES 12X obtain this result by subtracting their $B(\tau^- \rightarrow f_1(1285) \pi^- \nu_\tau \rightarrow \eta \pi^- \pi^+ \pi^- \nu_\tau)$ measurement from their $B(\tau^- \rightarrow \eta \pi^- \pi^+ \pi^- \nu_\tau (\text{ex. } K^0))$ measurement.

| $\Gamma(\eta a_1(1260)^- \nu_\tau \rightarrow \eta \pi^- \rho^0 \nu_\tau)/\Gamma_{\text{total}}$ | Γ_{160}/Γ | | | |
|--|-----------------------|--------------------|-------------|---|
| <u>VALUE</u> | <u>CL%</u> | <u>DOCUMENT ID</u> | <u>TECN</u> | <u>COMMENT</u> |
| $<3.9 \times 10^{-4}$ | 90 | BERGFELD | 97 | CLEO $E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$ |

| $\Gamma(\eta \eta \pi^- \nu_\tau)/\Gamma_{\text{total}}$ | Γ_{161}/Γ | | | |
|--|-----------------------|--------------------|-------------|----------------|
| <u>VALUE</u> | <u>CL%</u> | <u>DOCUMENT ID</u> | <u>TECN</u> | <u>COMMENT</u> |

| | | | | |
|-----------------------|----|-------|----|---|
| $<7.4 \times 10^{-6}$ | 90 | INAMI | 09 | BELL $490 \text{ fb}^{-1} E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$ |
|-----------------------|----|-------|----|---|

• • • We do not use the following data for averages, fits, limits, etc. • • •

| | | | | |
|-----------------------|----|----------|---------|---|
| $<1.1 \times 10^{-4}$ | 95 | ARTUSO | 92 | CLEO $E_{\text{cm}}^{\text{ee}} \approx 10.6 \text{ GeV}$ |
| $<8.3 \times 10^{-3}$ | 95 | ALBRECHT | 88M ARG | $E_{\text{cm}}^{\text{ee}} \approx 10 \text{ GeV}$ |

| $\Gamma(\eta\eta\pi^-\nu_\tau)/\Gamma_{\text{total}}$ | | | | Γ_{162}/Γ |
|---|------------|--------------------|-------------|---|
| <u>VALUE</u> (units 10^{-4}) | <u>CL%</u> | <u>DOCUMENT ID</u> | <u>TECN</u> | <u>COMMENT</u> |
| < 2.0 | 95 | ARTUSO | 92 | CLEO $E_{\text{cm}}^{\text{ee}} \approx 10.6 \text{ GeV}$ |
| $\bullet \bullet \bullet$ We do not use the following data for averages, fits, limits, etc. $\bullet \bullet \bullet$ | | | | |
| <90 | 95 | ALBRECHT | 88M | ARG $E_{\text{cm}}^{\text{ee}} \approx 10 \text{ GeV}$ |

| $\Gamma(\eta\eta K^-\nu_\tau)/\Gamma_{\text{total}}$ | | | | Γ_{163}/Γ |
|--|------------|--------------------|-------------|---|
| <u>VALUE</u> | <u>CL%</u> | <u>DOCUMENT ID</u> | <u>TECN</u> | <u>COMMENT</u> |
| $< 3.0 \times 10^{-6}$ | 90 | INAMI | 09 | BELL $490 \text{ fb}^{-1} E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$ |

| $\Gamma(\eta'(958)\pi^-\nu_\tau)/\Gamma_{\text{total}}$ | | | | Γ_{164}/Γ |
|---|------------|--------------------|-------------|--|
| <u>VALUE</u> | <u>CL%</u> | <u>DOCUMENT ID</u> | <u>TECN</u> | <u>COMMENT</u> |
| $< 4.0 \times 10^{-6}$ | 90 | LEES | 12X | BABR $468 \text{ fb}^{-1} E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$ |
| $\bullet \bullet \bullet$ We do not use the following data for averages, fits, limits, etc. $\bullet \bullet \bullet$ | | | | |
| $< 7.2 \times 10^{-6}$ | 90 | AUBERT | 08AE | BABR $384 \text{ fb}^{-1}, E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$ |
| $< 7.4 \times 10^{-5}$ | 90 | BERGFELD | 97 | CLEO $E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$ |

| $\Gamma(\eta'(958)\pi^-\pi^0\nu_\tau)/\Gamma_{\text{total}}$ | | | | Γ_{165}/Γ |
|---|------------|--------------------|-------------|---|
| <u>VALUE</u> | <u>CL%</u> | <u>DOCUMENT ID</u> | <u>TECN</u> | <u>COMMENT</u> |
| $< 1.2 \times 10^{-5}$ | 90 | LEES | 12X | BABR $468 \text{ fb}^{-1} E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$ |
| $\bullet \bullet \bullet$ We do not use the following data for averages, fits, limits, etc. $\bullet \bullet \bullet$ | | | | |
| $< 8.0 \times 10^{-5}$ | 90 | BERGFELD | 97 | CLEO $E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$ |

| $\Gamma(\eta'(958)K^-\nu_\tau)/\Gamma_{\text{total}}$ | | | | Γ_{166}/Γ |
|---|------------|--------------------|-------------|---|
| <u>VALUE</u> | <u>CL%</u> | <u>DOCUMENT ID</u> | <u>TECN</u> | <u>COMMENT</u> |
| $< 2.4 \times 10^{-6}$ | 90 | LEES | 12X | BABR $468 \text{ fb}^{-1} E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$ |

| $\Gamma(\phi\pi^-\nu_\tau)/\Gamma_{\text{total}}$ | | | | Γ_{167}/Γ | |
|---|------------|-------------|--------------------|-----------------------|--|
| <u>VALUE</u> (units 10^{-5}) | <u>CL%</u> | <u>EVTS</u> | <u>DOCUMENT ID</u> | <u>TECN</u> | <u>COMMENT</u> |
| $3.42 \pm 0.55 \pm 0.25$ | 344 | AUBERT | 08 | BABR | $342 \text{ fb}^{-1} E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$ |
| $\bullet \bullet \bullet$ We do not use the following data for averages, fits, limits, etc. $\bullet \bullet \bullet$ | | | | | |
| < 20 | 90 | 1 AVERY | 97 | CLEO | $E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$ |
| < 35 | 90 | ALBRECHT | 95H | ARG | $E_{\text{cm}}^{\text{ee}} = 9.4\text{--}10.6 \text{ GeV}$ |

¹ AVERY 97 limit varies from $(1.2\text{--}2.0) \times 10^{-4}$ depending on decay model assumptions.

| $\Gamma(\phi K^-\nu_\tau)/\Gamma_{\text{total}}$ | | | | Γ_{168}/Γ | |
|---|------------|-------------|--------------------|-----------------------|--|
| <u>VALUE</u> (units 10^{-5}) | <u>CL%</u> | <u>EVTS</u> | <u>DOCUMENT ID</u> | <u>TECN</u> | <u>COMMENT</u> |
| 4.4 ± 1.6 OUR FIT | | | | | |
| 3.70 ± 0.33 OUR AVERAGE | | | | | Error includes scale factor of 1.3. |
| $\bullet \bullet \bullet$ We use the following data for averages but not for fits. $\bullet \bullet \bullet$ | | | | | |
| 3.39 ± 0.20 ± 0.28 | 274 | AUBERT | 08 | BABR | $342 \text{ fb}^{-1} E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$ |
| 4.05 ± 0.25 ± 0.26 | 551 | INAMI | 06 | BELL | $401 \text{ fb}^{-1} E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$ |
| $\bullet \bullet \bullet$ We do not use the following data for averages, fits, limits, etc. $\bullet \bullet \bullet$ | | | | | |
| <6.7 | 90 | 1 AVERY | 97 | CLEO | $E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$ |
| ¹ AVERY 97 limit varies from $(5.4\text{--}6.7) \times 10^{-5}$ depending on decay model assumptions. | | | | | |

$\Gamma(f_1(1285)\pi^-\nu_\tau)/\Gamma_{\text{total}}$ Γ_{169}/Γ

| <u>VALUE</u> (units 10^{-4}) | <u>EVTS</u> | <u>DOCUMENT ID</u> | <u>TECN</u> | <u>COMMENT</u> |
|--|-------------|-----------------------|-------------|--|
| 3.9 ± 0.5 OUR AVERAGE | | | | Error includes scale factor of 1.9. |
| 4.73 ± 0.28 ± 0.45 | 3.7k | ¹ LEES | 12X BABR | $468 \text{ fb}^{-1} E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$ |
| 3.60 ± 0.18 ± 0.23 | 2.5k | ² LEES | 12X BABR | $468 \text{ fb}^{-1} E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$ |
| • • • We do not use the following data for averages, fits, limits, etc. • • • | | | | |
| 3.19 ± 0.18 ± 1.00 | 1.3 k | ³ AUBERT | 08AE BABR | Repl. by LEES 12X |
| 3.9 ± 0.7 ± 0.5 | 1.4 k | ⁴ AUBERT,B | 05W BABR | Repl. by LEES 12X |
| 5.8 $^{+1.4}_{-1.3}$ ± 1.8 | 54 | ⁵ BERGFELD | 97 CLEO | $E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$ |

¹ LEES 12X obtain this value by dividing their $B(\tau^- \rightarrow f_1(1285)\pi^-\nu_\tau \rightarrow 3\pi^- 2\pi^+\nu_\tau)$ measurement by the PDG 12 value of $B(f_1(1285) \rightarrow 2\pi^+ 2\pi^-) = 0.111^{+0.007}_{-0.006}$.

² LEES 12X obtain this value by dividing their $B(\tau^- \rightarrow f_1(1285)\pi^-\nu_\tau \rightarrow \eta\pi^- \pi^+ \pi^- \nu_\tau)$ measurement by 2/3 of the PDG 12 value of $B(f_1(1285) \rightarrow \eta\pi\pi) = 0.524^{+0.019}_{-0.021}$.

³ AUBERT 08AE obtain this value by dividing their $B(\tau^- \rightarrow f_1(1285)\pi^-\nu_\tau \rightarrow \eta\pi^- \pi^+ \pi^- \nu_\tau)$ measurement by the PDG 06 value of $B(f_1(1285) \rightarrow \eta\pi^- \pi^+) = 0.35 \pm 0.11$. The quote $(3.19 \pm 0.18 \pm 0.16 \pm 0.99) \times 10^{-4}$ where the final error is due to the uncertainty on $B(f_1(1285) \rightarrow \eta\pi^- \pi^+)$. We combine the two systematic errors in quadrature.

⁴ AUBERT,B 05W use the $f_1(1285) \rightarrow 2\pi^+ 2\pi^-$ decay mode and the PDG 04 value of $B(f_1(1285) \rightarrow 2\pi^+ 2\pi^-) = 0.110^{+0.007}_{-0.006}$.

⁵ BERGFELD 97 use the $f_1(1285) \rightarrow \eta\pi^+ \pi^-$ decay mode.

 $\Gamma(f_1(1285)\pi^-\nu_\tau \rightarrow \eta\pi^- \pi^+ \pi^- \nu_\tau)/\Gamma_{\text{total}}$ Γ_{170}/Γ

| <u>VALUE</u> (units 10^{-4}) | <u>EVTS</u> | <u>DOCUMENT ID</u> | <u>TECN</u> | <u>COMMENT</u> |
|---------------------------------|-------------|--------------------|-------------|---|
| 1.18 ± 0.07 OUR AVERAGE | | | | Error includes scale factor of 1.3. |
| 1.26 ± 0.06 ± 0.06 | 2.5k | LEES | 12X BABR | $468 \text{ fb}^{-1} E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$ |
| 1.11 ± 0.06 ± 0.05 | 1.3 k | AUBERT | 08AE BABR | $384 \text{ fb}^{-1}, E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$ |

 $\Gamma(f_1(1285)\pi^-\nu_\tau \rightarrow \eta\pi^- \pi^+ \pi^- \nu_\tau)/\Gamma(\eta\pi^- \pi^+ \pi^- \nu_\tau (\text{ex. } K^0))$ $\Gamma_{170}/\Gamma_{158}$

| <u>VALUE</u> | <u>DOCUMENT ID</u> | <u>TECN</u> | <u>COMMENT</u> |
|---------------------------|---------------------|-------------|---|
| 0.69 ± 0.01 ± 0.05 | ¹ AUBERT | 08AE BABR | $384 \text{ fb}^{-1}, E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$ |

• • • We do not use the following data for averages, fits, limits, etc. • • •

0.55 ± 0.14 BERGFELD 97 CLEO $E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$

¹ Not independent of AUBERT 08AE $B(\tau^- \rightarrow f_1(1285)\pi^-\nu_\tau \rightarrow \eta\pi^- \pi^+ \pi^- \nu_\tau)$ and $B(\tau^- \rightarrow \eta\pi^- \pi^+ \pi^- \nu_\tau (\text{ex. } K^0))$ values.

 $\Gamma(f_1(1285)\pi^-\nu_\tau \rightarrow 3\pi^- 2\pi^+\nu_\tau)/\Gamma_{\text{total}}$ Γ_{171}/Γ

| <u>VALUE</u> (units 10^{-4}) | <u>EVTS</u> | <u>DOCUMENT ID</u> | <u>TECN</u> | <u>COMMENT</u> |
|---------------------------------|-------------|--------------------|-------------|--|
| 0.52 ± 0.04 OUR FIT | | | | |
| 0.520 ± 0.031 ± 0.037 | 3.7k | LEES | 12X BABR | $468 \text{ fb}^{-1} E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$ |

 $\Gamma(\pi(1300)^-\nu_\tau \rightarrow (\rho\pi)^-\nu_\tau \rightarrow (3\pi)^-\nu_\tau)/\Gamma_{\text{total}}$ Γ_{172}/Γ

| <u>VALUE</u> | <u>CL%</u> | <u>DOCUMENT ID</u> | <u>TECN</u> | <u>COMMENT</u> |
|----------------------------------|------------|--------------------|-------------|---|
| <1.0 × 10⁻⁴ | 90 | ASNER | 00 | CLEO $E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$ |

| $\Gamma(\pi(1300)^-\nu_\tau \rightarrow ((\pi\pi)_{S\text{-wave}}\pi)^-\nu_\tau \rightarrow (3\pi)^-\nu_\tau)/\Gamma_{\text{total}}$ | Γ_{173}/Γ | | | |
|--|-----------------------|--------------------|-------------|---|
| <u>VALUE</u> | <u>CL%</u> | <u>DOCUMENT ID</u> | <u>TECN</u> | <u>COMMENT</u> |
| $<1.9 \times 10^{-4}$ | 90 | ASNER | 00 | CLEO $E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$ |

| $\Gamma(h^-\omega \geq 0 \text{ neutrals } \nu_\tau)/\Gamma_{\text{total}}$ | Γ_{174}/Γ | | | |
|---|-----------------------|--------------------|-------------|----------------|
| <u>VALUE (%)</u> | <u>EVTS</u> | <u>DOCUMENT ID</u> | <u>TECN</u> | <u>COMMENT</u> |
| 2.40 ± 0.08 OUR FIT | | | | |

• • • We use the following data for averages but not for fits. • • •

| | | | | |
|------------------------|------|----------|---------|--|
| $1.65 \pm 0.3 \pm 0.2$ | 1513 | ALBRECHT | 88M ARG | $E_{\text{cm}}^{\text{ee}} \approx 10 \text{ GeV}$ |
|------------------------|------|----------|---------|--|

| $\Gamma(h^-\omega\nu_\tau)/\Gamma_{\text{total}}$ | $\Gamma_{175}/\Gamma = (\Gamma_{176} + \Gamma_{177})/\Gamma$ | | | |
|---|--|--------------------|-------------|----------------|
| <u>VALUE (%)</u> | <u>EVTS</u> | <u>DOCUMENT ID</u> | <u>TECN</u> | <u>COMMENT</u> |

 1.99 ± 0.06 OUR FIT **1.92 ± 0.07 OUR AVERAGE**

| | | | | |
|--------------------------|------|----------|----------|--|
| $1.91 \pm 0.07 \pm 0.06$ | 5803 | BUSKULIC | 97C ALEP | 1991–1994 LEP runs |
| $1.60 \pm 0.27 \pm 0.41$ | 139 | BARINGER | 87 CLEO | $E_{\text{cm}}^{\text{ee}} = 10.5 \text{ GeV}$ |

• • • We use the following data for averages but not for fits. • • •

| | | | | |
|--------------------------|------|---------------------|----------|--|
| $1.95 \pm 0.07 \pm 0.11$ | 2223 | ¹ BALEST | 95C CLEO | $E_{\text{cm}}^{\text{ee}} \approx 10.6 \text{ GeV}$ |
|--------------------------|------|---------------------|----------|--|

¹ Not independent of BALEST 95C $B(\tau^- \rightarrow h^-\omega\nu_\tau)/B(\tau^- \rightarrow h^-h^+\pi^0\nu_\tau)$ value.

| $[\Gamma(\pi^-\omega\nu_\tau) + \Gamma(K^-\omega\nu_\tau)]/\Gamma(h^-h^-h^+\pi^0\nu_\tau \text{ (ex. } K^0\text{)})$ | $(\Gamma_{176} + \Gamma_{177})/\Gamma_{74}$ | | | |
|---|---|--------------------|-------------|----------------|
| $(\Gamma_{176} + \Gamma_{177})/\Gamma_{74} = (\Gamma_{176} + \Gamma_{177})/(\Gamma_{78} + \Gamma_{103} + \Gamma_{107} + 0.2292\Gamma_{150} + 0.892\Gamma_{176} + 0.892\Gamma_{177} + 0.0153\Gamma_{178})$ | | | | |
| <u>VALUE (units 10^{-2})</u> | <u>EVTS</u> | <u>DOCUMENT ID</u> | <u>TECN</u> | <u>COMMENT</u> |

 43.5 ± 1.4 OUR FIT **45.3 ± 1.9 OUR AVERAGE**

| | | | | |
|------------------------|------|-----------------------|----------|--|
| 43.1 ± 3.3 | 2350 | ¹ BUSKULIC | 96 ALEP | LEP 1991–1993 data |
| $46.4 \pm 1.6 \pm 1.7$ | 2223 | ² BALEST | 95C CLEO | $E_{\text{cm}}^{\text{ee}} \approx 10.6 \text{ GeV}$ |

• • • We do not use the following data for averages, fits, limits, etc. • • •

| | | | | |
|------------------|-----|-----------------------|---------|--|
| $37 \pm 5 \pm 2$ | 458 | ³ ALBRECHT | 91D ARG | $E_{\text{cm}}^{\text{ee}} = 9.4\text{--}10.6 \text{ GeV}$ |
|------------------|-----|-----------------------|---------|--|

¹ BUSKULIC 96 quote the fraction of $\tau^- \rightarrow h^-h^-h^+\pi^0\nu_\tau$ (ex. K^0) decays which originate in a $h^-\omega$ final state $= 0.383 \pm 0.029$. We divide this by the $\omega(782) \rightarrow \pi^+\pi^-\pi^0$ branching fraction (0.888).² BALEST 95C quote the fraction of $\tau^- \rightarrow h^-h^-h^+\pi^0\nu_\tau$ (ex. K^0) decays which originate in a $h^-\omega$ final state equals $0.412 \pm 0.014 \pm 0.015$. We divide this by the $\omega(782) \rightarrow \pi^+\pi^-\pi^0$ branching fraction (0.888).³ ALBRECHT 91D quote the fraction of $\tau^- \rightarrow h^-h^-h^+\pi^0\nu_\tau$ decays which originate in a $\pi^-\omega$ final state equals $0.33 \pm 0.04 \pm 0.02$. We divide this by the $\omega(782) \rightarrow \pi^+\pi^-\pi^0$ branching fraction (0.888).

| $\Gamma(\pi^-\omega\nu_\tau)/\Gamma_{\text{total}}$ | Γ_{176}/Γ |
|---|-----------------------|
| <u>VALUE (%)</u> | <u>DOCUMENT ID</u> |
| 1.95 ± 0.06 OUR FIT | |

$\Gamma(K^-\omega\nu_\tau)/\Gamma_{\text{total}}$ Γ_{177}/Γ

| <u>VALUE (units 10^{-4})</u> | <u>EVTS</u> | <u>DOCUMENT ID</u> | <u>TECN</u> | <u>COMMENT</u> |
|---|-------------|--------------------|-------------|---|
| 4.1±0.9 OUR FIT | | | | |
| 4.1±0.6±0.7 | 500 | ARMS | 05 | CLE3 7.6 fb^{-1} , $E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$ |

 $\Gamma(h^-\omega\pi^0\nu_\tau)/\Gamma_{\text{total}}$ Γ_{178}/Γ

| <u>VALUE (%)</u> | <u>EVTS</u> | <u>DOCUMENT ID</u> | <u>TECN</u> | <u>COMMENT</u> |
|--------------------------|-------------|--------------------|-------------|----------------------------|
| 0.41±0.04 OUR FIT | | | | |
| 0.43±0.06±0.05 | 7283 | BUSKULIC | 97C | ALEP 1991–1994 LEP runs |

 $\Gamma(h^-\omega\pi^0\nu_\tau)/\Gamma(h^-h^-h^+\geq 0 K_L^0\nu_\tau)$ Γ_{178}/Γ_{62}

$$\Gamma_{178}/\Gamma_{62} = \Gamma_{178}/(\Gamma_{36} + 0.34598\Gamma_{38} + 0.34598\Gamma_{41} + 0.34598\Gamma_{43} + 0.4247\Gamma_{48} + 0.6920\Gamma_{49} + 0.8494\Gamma_{52} + 0.6920\Gamma_{56} + 0.6534\Gamma_{61} + \Gamma_{70} + \Gamma_{78} + \Gamma_{85} + \Gamma_{89} + \Gamma_{97} + \Gamma_{103} + \Gamma_{106} + \Gamma_{107} + 0.2810\Gamma_{148} + 0.2292\Gamma_{149} + 0.2810\Gamma_{150} + 0.2810\Gamma_{152} + 0.3759\Gamma_{154} + 0.3268\Gamma_{158} + 0.7259\Gamma_{168} + 0.9078\Gamma_{176} + 0.9078\Gamma_{177} + 0.9078\Gamma_{178} + 0.892\Gamma_{180})$$

| <u>VALUE</u> | <u>EVTS</u> | <u>DOCUMENT ID</u> | <u>TECN</u> | <u>COMMENT</u> |
|--|-------------|--------------------|-------------|----------------|
| (2.69 ±0.28) × 10⁻² OUR FIT | | | | |

• • • We use the following data for averages but not for fits. • • •

0.028±0.003±0.003 430 ¹ BORTOLETTO 93 CLEO $E_{\text{cm}}^{\text{ee}} \approx 10.6 \text{ GeV}$

¹ Not independent of BORTOLETTO 93 $\Gamma(\tau^- \rightarrow h^-\omega\pi^0\nu_\tau)/\Gamma(\tau^- \rightarrow h^-h^-h^+2\pi^0\nu_\tau \text{ (ex. } K^0\text{)})$ value.

 $\Gamma(h^-\omega\pi^0\nu_\tau)/\Gamma(h^-h^-h^+2\pi^0\nu_\tau \text{ (ex. } K^0\text{)})$ Γ_{178}/Γ_{84}

$$\Gamma_{178}/\Gamma_{84} = \Gamma_{178}/(\Gamma_{85} + 0.2292\Gamma_{148} + 0.2292\Gamma_{152} + 0.892\Gamma_{178})$$

| <u>VALUE (units 10^{-2})</u> | <u>DOCUMENT ID</u> | <u>TECN</u> | <u>COMMENT</u> |
|---|--------------------|-------------|--|
| 82±8 OUR FIT | | | |
| 81±6±6 | BORTOLETTO93 | CLEO | $E_{\text{cm}}^{\text{ee}} \approx 10.6 \text{ GeV}$ |

 $\Gamma(h^-\omega 2\pi^0\nu_\tau)/\Gamma_{\text{total}}$ Γ_{179}/Γ

| <u>VALUE (units 10^{-4})</u> | <u>EVTS</u> | <u>DOCUMENT ID</u> | <u>TECN</u> | <u>COMMENT</u> |
|---|-------------|--------------------|-------------|--|
| 1.4 ±0.4 ±0.3 | 53 | ANASTASSOV 01 | CLEO | $E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$ |
| • • • We do not use the following data for averages, fits, limits, etc. • • • | | | | |

$1.89^{+0.74}_{-0.67} \pm 0.40$ 19 ANDERSON 97 CLEO Repl. by ANASTASSOV 01

 $\Gamma(\pi^-\omega 2\pi^0\nu_\tau)/\Gamma_{\text{total}}$ Γ_{180}/Γ

| <u>VALUE (units 10^{-4})</u> | <u>EVTS</u> | <u>DOCUMENT ID</u> | <u>TECN</u> | <u>COMMENT</u> |
|---|-------------|--------------------|-------------|---|
| 0.71±0.16 OUR FIT | | | | |
| 0.73±0.12±0.12 | 1.1k | LEES | 12x | BABR 468 fb^{-1} , $E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$ |

 $\Gamma(h^-2\omega\nu_\tau)/\Gamma_{\text{total}}$ Γ_{181}/Γ

| <u>VALUE</u> | <u>CL%</u> | <u>DOCUMENT ID</u> | <u>TECN</u> | <u>COMMENT</u> |
|----------------------------------|------------|--------------------|-------------|---|
| <5.4 × 10⁻⁷ | 90 | AUBERT,B | 06 | BABR 232 fb^{-1} , $E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$ |

 $\Gamma(2h^-h^+\omega\nu_\tau)/\Gamma_{\text{total}}$ Γ_{182}/Γ

| <u>VALUE (units 10^{-4})</u> | <u>EVTS</u> | <u>DOCUMENT ID</u> | <u>TECN</u> | <u>COMMENT</u> |
|---|-------------|--------------------|-------------|--|
| 1.2±0.2±0.1 | 110 | ANASTASSOV 01 | CLEO | $E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$ |

$\Gamma(2\pi^-\pi^+\omega\nu_\tau(\text{ex.}K^0))/\Gamma_{\text{total}}$ Γ_{183}/Γ

| <u>VALUE (units 10^{-4})</u> | <u>EVTS</u> | <u>DOCUMENT ID</u> | <u>TECN</u> | <u>COMMENT</u> |
|--|-------------|--------------------|-------------|--|
| 0.84 ± 0.06 OUR FIT | | | | |
| $0.84 \pm 0.04 \pm 0.06$ | 2.4k | LEES | 12x BABR | 468 fb^{-1} $E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$ |

 $\Gamma(e^-\gamma)/\Gamma_{\text{total}}$ Γ_{184}/Γ

Test of lepton family number conservation.

| <u>VALUE</u> | <u>CL%</u> | <u>DOCUMENT ID</u> | <u>TECN</u> | <u>COMMENT</u> |
|--|------------|--------------------|-------------|---|
| $< 3.3 \times 10^{-8}$ | 90 | AUBERT | 10B BABR | 516 fb^{-1} , $E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$ |
| • • • We do not use the following data for averages, fits, limits, etc. • • • | | | | |
| $< 1.2 \times 10^{-7}$ | 90 | HAYASAKA | 08 BELL | 535 fb^{-1} , $E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$ |
| $< 1.1 \times 10^{-7}$ | 90 | AUBERT | 06C BABR | 232 fb^{-1} , $E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$ |
| $< 3.9 \times 10^{-7}$ | 90 | HAYASAKA | 05 BELL | 86.7 fb^{-1} , $E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$ |
| $< 2.7 \times 10^{-6}$ | 90 | EDWARDS | 97 CLEO | |
| $< 1.1 \times 10^{-4}$ | 90 | ABREU | 95U DLPH | 1990–1993 LEP runs |
| $< 1.2 \times 10^{-4}$ | 90 | ALBRECHT | 92K ARG | $E_{\text{cm}}^{\text{ee}} = 10 \text{ GeV}$ |
| $< 2.0 \times 10^{-4}$ | 90 | KEH | 88 CBAL | $E_{\text{cm}}^{\text{ee}} = 10 \text{ GeV}$ |
| $< 6.4 \times 10^{-4}$ | 90 | HAYES | 82 MRK2 | $E_{\text{cm}}^{\text{ee}} = 3.8\text{--}6.8 \text{ GeV}$ |

 $\Gamma(\mu^-\gamma)/\Gamma_{\text{total}}$ Γ_{185}/Γ

Test of lepton family number conservation.

| <u>VALUE</u> | <u>CL%</u> | <u>DOCUMENT ID</u> | <u>TECN</u> | <u>COMMENT</u> |
|--|------------|--------------------|-------------|---|
| $< 4.4 \times 10^{-8}$ | 90 | AUBERT | 10B BABR | 516 fb^{-1} , $E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$ |
| • • • We do not use the following data for averages, fits, limits, etc. • • • | | | | |
| $< 4.5 \times 10^{-8}$ | 90 | HAYASAKA | 08 BELL | 535 fb^{-1} , $E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$ |
| $< 6.8 \times 10^{-8}$ | 90 | AUBERT,B | 05A BABR | 232 fb^{-1} , $E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$ |
| $< 3.1 \times 10^{-7}$ | 90 | ABE | 04B BELL | 86.3 fb^{-1} , $E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$ |
| $< 1.1 \times 10^{-6}$ | 90 | AHMED | 00 CLEO | $E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$ |
| $< 3.0 \times 10^{-6}$ | 90 | EDWARDS | 97 CLEO | |
| $< 6.2 \times 10^{-5}$ | 90 | ABREU | 95U DLPH | 1990–1993 LEP runs |
| $< 0.42 \times 10^{-5}$ | 90 | BEAN | 93 CLEO | $E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$ |
| $< 3.4 \times 10^{-5}$ | 90 | ALBRECHT | 92K ARG | $E_{\text{cm}}^{\text{ee}} = 10 \text{ GeV}$ |
| $< 55 \times 10^{-5}$ | 90 | HAYES | 82 MRK2 | $E_{\text{cm}}^{\text{ee}} = 3.8\text{--}6.8 \text{ GeV}$ |

 $\Gamma(e^-\pi^0)/\Gamma_{\text{total}}$ Γ_{186}/Γ

Test of lepton family number conservation.

| <u>VALUE</u> | <u>CL%</u> | <u>DOCUMENT ID</u> | <u>TECN</u> | <u>COMMENT</u> |
|--|------------|--------------------|-------------|--|
| $< 8.0 \times 10^{-8}$ | 90 | MIYAZAKI | 07 BELL | 401 fb^{-1} , $E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$ |
| • • • We do not use the following data for averages, fits, limits, etc. • • • | | | | |
| $< 1.3 \times 10^{-7}$ | 90 | AUBERT | 07I BABR | 339 fb^{-1} , $E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$ |
| $< 1.9 \times 10^{-7}$ | 90 | ENARI | 05 BELL | 154 fb^{-1} , $E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$ |
| $< 3.7 \times 10^{-6}$ | 90 | BONVICINI | 97 CLEO | $E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$ |
| $< 17 \times 10^{-5}$ | 90 | ALBRECHT | 92K ARG | $E_{\text{cm}}^{\text{ee}} = 10 \text{ GeV}$ |
| $< 14 \times 10^{-5}$ | 90 | KEH | 88 CBAL | $E_{\text{cm}}^{\text{ee}} = 10 \text{ GeV}$ |
| $< 210 \times 10^{-5}$ | 90 | HAYES | 82 MRK2 | $E_{\text{cm}}^{\text{ee}} = 3.8\text{--}6.8 \text{ GeV}$ |

$\Gamma(\mu^- \pi^0)/\Gamma_{\text{total}}$ Γ_{187}/Γ

Test of lepton family number conservation.

| VALUE | CL% | DOCUMENT ID | TECN | COMMENT |
|---|-----|-------------|------|---|
| $< 1.1 \times 10^{-7}$ | 90 | AUBERT | 07I | BABR 339 fb^{-1} , $E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$ |
| $\bullet \bullet \bullet$ We do not use the following data for averages, fits, limits, etc. $\bullet \bullet \bullet$ | | | | |
| $< 1.2 \times 10^{-7}$ | 90 | MIYAZAKI | 07 | BELL 401 fb^{-1} , $E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$ |
| $< 4.1 \times 10^{-7}$ | 90 | ENARI | 05 | BELL 154 fb^{-1} , $E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$ |
| $< 4.0 \times 10^{-6}$ | 90 | BONVICINI | 97 | CLEO $E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$ |
| $< 4.4 \times 10^{-5}$ | 90 | ALBRECHT | 92K | ARG $E_{\text{cm}}^{\text{ee}} = 10 \text{ GeV}$ |
| $< 82 \times 10^{-5}$ | 90 | HAYES | 82 | MRK2 $E_{\text{cm}}^{\text{ee}} = 3.8\text{--}6.8 \text{ GeV}$ |

 $\Gamma(e^- K_S^0)/\Gamma_{\text{total}}$ Γ_{188}/Γ

Test of lepton family number conservation.

| VALUE | CL% | DOCUMENT ID | TECN | COMMENT |
|---|-----|-------------|------|---|
| $< 2.6 \times 10^{-8}$ | 90 | MIYAZAKI | 10A | BELL 671 fb^{-1} , $E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$ |
| $\bullet \bullet \bullet$ We do not use the following data for averages, fits, limits, etc. $\bullet \bullet \bullet$ | | | | |
| $< 3.3 \times 10^{-8}$ | 90 | AUBERT | 09D | BABR 469 fb^{-1} , $E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$ |
| $< 5.6 \times 10^{-8}$ | 90 | MIYAZAKI | 06A | BELL 281 fb^{-1} , $E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$ |
| $< 9.1 \times 10^{-7}$ | 90 | CHEN | 02C | CLEO $E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$ |
| $< 1.3 \times 10^{-3}$ | 90 | HAYES | 82 | MRK2 $E_{\text{cm}}^{\text{ee}} = 3.8\text{--}6.8 \text{ GeV}$ |

 $\Gamma(\mu^- K_S^0)/\Gamma_{\text{total}}$ Γ_{189}/Γ

Test of lepton family number conservation.

| VALUE | CL% | DOCUMENT ID | TECN | COMMENT |
|---|-----|-------------|------|---|
| $< 2.3 \times 10^{-8}$ | 90 | MIYAZAKI | 10A | BELL 671 fb^{-1} , $E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$ |
| $\bullet \bullet \bullet$ We do not use the following data for averages, fits, limits, etc. $\bullet \bullet \bullet$ | | | | |
| $< 4.0 \times 10^{-8}$ | 90 | AUBERT | 09D | BABR 469 fb^{-1} , $E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$ |
| $< 4.9 \times 10^{-8}$ | 90 | MIYAZAKI | 06A | BELL 281 fb^{-1} , $E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$ |
| $< 9.5 \times 10^{-7}$ | 90 | CHEN | 02C | CLEO $E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$ |
| $< 1.0 \times 10^{-3}$ | 90 | HAYES | 82 | MRK2 $E_{\text{cm}}^{\text{ee}} = 3.8\text{--}6.8 \text{ GeV}$ |

 $\Gamma(e^- \eta)/\Gamma_{\text{total}}$ Γ_{190}/Γ

Test of lepton family number conservation.

| VALUE | CL% | DOCUMENT ID | TECN | COMMENT |
|---|-----|-------------|------|---|
| $< 9.2 \times 10^{-8}$ | 90 | MIYAZAKI | 07 | BELL 401 fb^{-1} , $E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$ |
| $\bullet \bullet \bullet$ We do not use the following data for averages, fits, limits, etc. $\bullet \bullet \bullet$ | | | | |
| $< 1.6 \times 10^{-7}$ | 90 | AUBERT | 07I | BABR 339 fb^{-1} , $E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$ |
| $< 2.4 \times 10^{-7}$ | 90 | ENARI | 05 | BELL 154 fb^{-1} , $E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$ |
| $< 8.2 \times 10^{-6}$ | 90 | BONVICINI | 97 | CLEO $E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$ |
| $< 6.3 \times 10^{-5}$ | 90 | ALBRECHT | 92K | ARG $E_{\text{cm}}^{\text{ee}} = 10 \text{ GeV}$ |
| $< 24 \times 10^{-5}$ | 90 | KEH | 88 | CBAL $E_{\text{cm}}^{\text{ee}} = 10 \text{ GeV}$ |

$\Gamma(\mu^- \eta)/\Gamma_{\text{total}}$ Γ_{191}/Γ

Test of lepton family number conservation.

| <u>VALUE</u> | <u>CL%</u> | <u>DOCUMENT ID</u> | <u>TECN</u> | <u>COMMENT</u> |
|---|------------|--------------------|-------------|---|
| $< 6.5 \times 10^{-8}$ | 90 | MIYAZAKI | 07 | BELL $401 \text{ fb}^{-1}, E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$ |
| $\bullet \bullet \bullet$ We do not use the following data for averages, fits, limits, etc. $\bullet \bullet \bullet$ | | | | |
| $< 1.5 \times 10^{-7}$ | 90 | AUBERT | 07I | BABR $339 \text{ fb}^{-1}, E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$ |
| $< 1.5 \times 10^{-7}$ | 90 | ENARI | 05 | BELL $154 \text{ fb}^{-1}, E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$ |
| $< 3.4 \times 10^{-7}$ | 90 | ENARI | 04 | BELL $84.3 \text{ fb}^{-1}, E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$ |
| $< 9.6 \times 10^{-6}$ | 90 | BONVICINI | 97 | CLEO $E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$ |
| $< 7.3 \times 10^{-5}$ | 90 | ALBRECHT | 92K | ARG $E_{\text{cm}}^{\text{ee}} = 10 \text{ GeV}$ |

 $\Gamma(e^- \rho^0)/\Gamma_{\text{total}}$ Γ_{192}/Γ

Test of lepton family number conservation.

| <u>VALUE</u> | <u>CL%</u> | <u>DOCUMENT ID</u> | <u>TECN</u> | <u>COMMENT</u> |
|---|------------|---------------------|-------------|---|
| $< 1.8 \times 10^{-8}$ | 90 | MIYAZAKI | 11 | BELL $854 \text{ fb}^{-1} E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$ |
| $\bullet \bullet \bullet$ We do not use the following data for averages, fits, limits, etc. $\bullet \bullet \bullet$ | | | | |
| $< 4.6 \times 10^{-8}$ | 90 | AUBERT | 09W | BABR $451 \text{ fb}^{-1} E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$ |
| $< 6.3 \times 10^{-8}$ | 90 | NISHIO | 08 | BELL $543 \text{ fb}^{-1} E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$ |
| $< 6.5 \times 10^{-7}$ | 90 | YUSA | 06 | BELL $158 \text{ fb}^{-1} E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$ |
| $< 2.0 \times 10^{-6}$ | 90 | BLISS | 98 | CLEO $E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$ |
| $< 4.2 \times 10^{-6}$ | 90 | ¹ BARTEL | 94 | CLEO Repl. by BLISS 98 |
| $< 1.9 \times 10^{-5}$ | 90 | ALBRECHT | 92K | ARG $E_{\text{cm}}^{\text{ee}} = 10 \text{ GeV}$ |
| $< 37 \times 10^{-5}$ | 90 | HAYES | 82 | MRK2 $E_{\text{cm}}^{\text{ee}} = 3.8\text{--}6.8 \text{ GeV}$ |

¹ BARTEL 94 assume phase space decays. $\Gamma(\mu^- \rho^0)/\Gamma_{\text{total}}$ Γ_{193}/Γ

Test of lepton family number conservation.

| <u>VALUE</u> | <u>CL%</u> | <u>DOCUMENT ID</u> | <u>TECN</u> | <u>COMMENT</u> |
|---|------------|---------------------|-------------|---|
| $< 1.2 \times 10^{-8}$ | 90 | MIYAZAKI | 11 | BELL $854 \text{ fb}^{-1} E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$ |
| $\bullet \bullet \bullet$ We do not use the following data for averages, fits, limits, etc. $\bullet \bullet \bullet$ | | | | |
| $< 2.6 \times 10^{-8}$ | 90 | AUBERT | 09W | BABR $451 \text{ fb}^{-1} E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$ |
| $< 6.8 \times 10^{-8}$ | 90 | NISHIO | 08 | BELL $543 \text{ fb}^{-1} E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$ |
| $< 2.0 \times 10^{-7}$ | 90 | YUSA | 06 | BELL $158 \text{ fb}^{-1} E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$ |
| $< 6.3 \times 10^{-6}$ | 90 | BLISS | 98 | CLEO $E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$ |
| $< 5.7 \times 10^{-6}$ | 90 | ¹ BARTEL | 94 | CLEO Repl. by BLISS 98 |
| $< 2.9 \times 10^{-5}$ | 90 | ALBRECHT | 92K | ARG $E_{\text{cm}}^{\text{ee}} = 10 \text{ GeV}$ |
| $< 44 \times 10^{-5}$ | 90 | HAYES | 82 | MRK2 $E_{\text{cm}}^{\text{ee}} = 3.8\text{--}6.8 \text{ GeV}$ |

¹ BARTEL 94 assume phase space decays. $\Gamma(e^- \omega)/\Gamma_{\text{total}}$ Γ_{194}/Γ

| <u>VALUE</u> | <u>CL%</u> | <u>DOCUMENT ID</u> | <u>TECN</u> | <u>COMMENT</u> |
|---|------------|--------------------|-------------|---|
| $< 4.8 \times 10^{-8}$ | 90 | MIYAZAKI | 11 | BELL $854 \text{ fb}^{-1} E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$ |
| $\bullet \bullet \bullet$ We do not use the following data for averages, fits, limits, etc. $\bullet \bullet \bullet$ | | | | |
| $< 1.1 \times 10^{-7}$ | 90 | AUBERT | 08K | BABR $384 \text{ fb}^{-1} E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$ |
| $< 1.8 \times 10^{-7}$ | 90 | NISHIO | 08 | BELL $543 \text{ fb}^{-1} E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$ |

$\Gamma(\mu^- \omega)/\Gamma_{\text{total}}$ Γ_{195}/Γ

| <u>VALUE</u> | <u>CL%</u> | <u>DOCUMENT ID</u> | <u>TECN</u> | <u>COMMENT</u> |
|---|------------|--------------------|-------------|---|
| $<4.7 \times 10^{-8}$ | 90 | MIYAZAKI | 11 | BELL $854 \text{ fb}^{-1} E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$ |
| $\bullet \bullet \bullet$ We do not use the following data for averages, fits, limits, etc. $\bullet \bullet \bullet$ | | | | |
| $<1.0 \times 10^{-7}$ | 90 | AUBERT | 08W | BABR $384 \text{ fb}^{-1} E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$ |
| $<8.9 \times 10^{-8}$ | 90 | NISHIO | 08 | BELL $543 \text{ fb}^{-1} E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$ |

 $\Gamma(e^- K^*(892)^0)/\Gamma_{\text{total}}$ Γ_{196}/Γ

Test of lepton family number conservation.

| <u>VALUE</u> | <u>CL%</u> | <u>DOCUMENT ID</u> | <u>TECN</u> | <u>COMMENT</u> |
|---|------------|---------------------|-------------|---|
| $<3.2 \times 10^{-8}$ | 90 | MIYAZAKI | 11 | BELL $854 \text{ fb}^{-1} E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$ |
| $\bullet \bullet \bullet$ We do not use the following data for averages, fits, limits, etc. $\bullet \bullet \bullet$ | | | | |
| $<5.9 \times 10^{-8}$ | 90 | AUBERT | 09W | BABR $451 \text{ fb}^{-1} E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$ |
| $<7.8 \times 10^{-8}$ | 90 | NISHIO | 08 | BELL $543 \text{ fb}^{-1} E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$ |
| $<3.0 \times 10^{-7}$ | 90 | YUSA | 06 | BELL $158 \text{ fb}^{-1} E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$ |
| $<5.1 \times 10^{-6}$ | 90 | BLISS | 98 | CLEO $E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$ |
| $<6.3 \times 10^{-6}$ | 90 | ¹ BARTEL | 94 | CLEO Repl. by BLISS 98 |
| $<3.8 \times 10^{-5}$ | 90 | ALBRECHT | 92K | ARG $E_{\text{cm}}^{\text{ee}} = 10 \text{ GeV}$ |

¹ BARTEL 94 assume phase space decays. $\Gamma(\mu^- K^*(892)^0)/\Gamma_{\text{total}}$ Γ_{197}/Γ

Test of lepton family number conservation.

| <u>VALUE</u> | <u>CL%</u> | <u>DOCUMENT ID</u> | <u>TECN</u> | <u>COMMENT</u> |
|---|------------|---------------------|-------------|---|
| $<5.9 \times 10^{-8}$ | 90 | NISHIO | 08 | BELL $543 \text{ fb}^{-1} E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$ |
| $\bullet \bullet \bullet$ We do not use the following data for averages, fits, limits, etc. $\bullet \bullet \bullet$ | | | | |
| $<7.2 \times 10^{-8}$ | 90 | MIYAZAKI | 11 | BELL $854 \text{ fb}^{-1} E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$ |
| $<1.7 \times 10^{-7}$ | 90 | AUBERT | 09W | BABR $451 \text{ fb}^{-1} E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$ |
| $<3.9 \times 10^{-7}$ | 90 | YUSA | 06 | BELL $158 \text{ fb}^{-1} E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$ |
| $<7.5 \times 10^{-6}$ | 90 | BLISS | 98 | CLEO $E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$ |
| $<9.4 \times 10^{-6}$ | 90 | ¹ BARTEL | 94 | CLEO Repl. by BLISS 98 |
| $<4.5 \times 10^{-5}$ | 90 | ALBRECHT | 92K | ARG $E_{\text{cm}}^{\text{ee}} = 10 \text{ GeV}$ |

¹ BARTEL 94 assume phase space decays. $\Gamma(e^- \bar{K}^*(892)^0)/\Gamma_{\text{total}}$ Γ_{198}/Γ

Test of lepton family number conservation.

| <u>VALUE</u> | <u>CL%</u> | <u>DOCUMENT ID</u> | <u>TECN</u> | <u>COMMENT</u> |
|---|------------|---------------------|-------------|---|
| $<3.4 \times 10^{-8}$ | 90 | MIYAZAKI | 11 | BELL $854 \text{ fb}^{-1} E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$ |
| $\bullet \bullet \bullet$ We do not use the following data for averages, fits, limits, etc. $\bullet \bullet \bullet$ | | | | |
| $<4.6 \times 10^{-8}$ | 90 | AUBERT | 09W | BABR $451 \text{ fb}^{-1} E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$ |
| $<7.7 \times 10^{-8}$ | 90 | NISHIO | 08 | BELL $543 \text{ fb}^{-1} E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$ |
| $<4.0 \times 10^{-7}$ | 90 | YUSA | 06 | BELL $158 \text{ fb}^{-1} E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$ |
| $<7.4 \times 10^{-6}$ | 90 | BLISS | 98 | CLEO $E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$ |
| $<1.1 \times 10^{-5}$ | 90 | ¹ BARTEL | 94 | CLEO Repl. by BLISS 98 |

¹ BARTEL 94 assume phase space decays.

$\Gamma(\mu^- \bar{K}^*(892)^0)/\Gamma_{\text{total}}$ Γ_{199}/Γ

Test of lepton family number conservation.

| <u>VALUE</u> | <u>CL%</u> | <u>DOCUMENT ID</u> | <u>TECN</u> | <u>COMMENT</u> |
|---|------------|---------------------|-------------|---|
| $<7.0 \times 10^{-8}$ | 90 | MIYAZAKI | 11 | BELL 854 fb^{-1} $E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$ |
| $\bullet \bullet \bullet$ We do not use the following data for averages, fits, limits, etc. $\bullet \bullet \bullet$ | | | | |
| $<7.3 \times 10^{-8}$ | 90 | AUBERT | 09W | BABR 451 fb^{-1} $E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$ |
| $<1.0 \times 10^{-7}$ | 90 | NISHIO | 08 | BELL 543 fb^{-1} $E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$ |
| $<4.0 \times 10^{-7}$ | 90 | YUSA | 06 | BELL 158 fb^{-1} $E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$ |
| $<7.5 \times 10^{-6}$ | 90 | BLISS | 98 | CLEO $E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$ |
| $<8.7 \times 10^{-6}$ | 90 | ¹ BARTEL | 94 | CLEO Repl. by BLISS 98 |

¹ BARTEL 94 assume phase space decays. $\Gamma(e^- \eta'(958))/\Gamma_{\text{total}}$ Γ_{200}/Γ

| <u>VALUE</u> | <u>CL%</u> | <u>DOCUMENT ID</u> | <u>TECN</u> | <u>COMMENT</u> |
|---|------------|--------------------|-------------|---|
| $< 1.6 \times 10^{-7}$ | 90 | MIYAZAKI | 07 | BELL 401 fb^{-1} , $E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$ |
| $\bullet \bullet \bullet$ We do not use the following data for averages, fits, limits, etc. $\bullet \bullet \bullet$ | | | | |
| $< 2.4 \times 10^{-7}$ | 90 | AUBERT | 07I | BABR 339 fb^{-1} , $E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$ |
| $< 10. \times 10^{-7}$ | 90 | ENARI | 05 | BELL 154 fb^{-1} , $E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$ |

 $\Gamma(\mu^- \eta'(958))/\Gamma_{\text{total}}$ Γ_{201}/Γ

| <u>VALUE</u> | <u>CL%</u> | <u>DOCUMENT ID</u> | <u>TECN</u> | <u>COMMENT</u> |
|---|------------|--------------------|-------------|---|
| $< 1.3 \times 10^{-7}$ | 90 | MIYAZAKI | 07 | BELL 401 fb^{-1} , $E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$ |
| $\bullet \bullet \bullet$ We do not use the following data for averages, fits, limits, etc. $\bullet \bullet \bullet$ | | | | |
| $< 1.4 \times 10^{-7}$ | 90 | AUBERT | 07I | BABR 339 fb^{-1} , $E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$ |
| $< 4.7 \times 10^{-7}$ | 90 | ENARI | 05 | BELL 154 fb^{-1} , $E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$ |

 $\Gamma(e^- f_0(980) \rightarrow e^- \pi^+ \pi^-)/\Gamma_{\text{total}}$ Γ_{202}/Γ

| <u>VALUE</u> | <u>CL%</u> | <u>DOCUMENT ID</u> | <u>TECN</u> | <u>COMMENT</u> |
|------------------------|------------|--------------------|-------------|---|
| $< 3.2 \times 10^{-8}$ | 90 | MIYAZAKI | 09 | BELL 671 fb^{-1} $E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$ |

 $\Gamma(\mu^- f_0(980) \rightarrow \mu^- \pi^+ \pi^-)/\Gamma_{\text{total}}$ Γ_{203}/Γ

| <u>VALUE</u> | <u>CL%</u> | <u>DOCUMENT ID</u> | <u>TECN</u> | <u>COMMENT</u> |
|------------------------|------------|--------------------|-------------|---|
| $< 3.4 \times 10^{-8}$ | 90 | MIYAZAKI | 09 | BELL 671 fb^{-1} $E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$ |

 $\Gamma(e^- \phi)/\Gamma_{\text{total}}$ Γ_{204}/Γ

Test of lepton family number conservation.

| <u>VALUE</u> | <u>CL%</u> | <u>DOCUMENT ID</u> | <u>TECN</u> | <u>COMMENT</u> |
|---|------------|--------------------|-------------|---|
| $< 3.1 \times 10^{-8}$ | 90 | MIYAZAKI | 11 | BELL 854 fb^{-1} $E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$ |
| $< 3.1 \times 10^{-8}$ | 90 | AUBERT | 09W | BABR 451 fb^{-1} $E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$ |
| $\bullet \bullet \bullet$ We do not use the following data for averages, fits, limits, etc. $\bullet \bullet \bullet$ | | | | |
| $< 7.3 \times 10^{-8}$ | 90 | NISHIO | 08 | BELL 543 fb^{-1} $E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$ |
| $< 7.3 \times 10^{-7}$ | 90 | YUSA | 06 | BELL 158 fb^{-1} $E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$ |
| $< 6.9 \times 10^{-6}$ | 90 | BLISS | 98 | CLEO $E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$ |

$\Gamma(\mu^- \phi)/\Gamma_{\text{total}}$ Γ_{205}/Γ

Test of lepton family number conservation.

| <u>VALUE</u> | <u>CL%</u> | <u>DOCUMENT ID</u> | <u>TECN</u> | <u>COMMENT</u> |
|---|------------|--------------------|-------------|---|
| $< 8.4 \times 10^{-8}$ | 90 | MIYAZAKI | 11 | BELL $854 \text{ fb}^{-1} E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$ |
| $\bullet \bullet \bullet$ We do not use the following data for averages, fits, limits, etc. $\bullet \bullet \bullet$ | | | | |
| $< 1.9 \times 10^{-7}$ | 90 | AUBERT | 09W | BABR $451 \text{ fb}^{-1} E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$ |
| $< 1.3 \times 10^{-7}$ | 90 | NISHIO | 08 | BELL $543 \text{ fb}^{-1} E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$ |
| $< 7.7 \times 10^{-7}$ | 90 | YUSA | 06 | BELL $158 \text{ fb}^{-1} E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$ |
| $< 7.0 \times 10^{-6}$ | 90 | BLISS | 98 | CLEO $E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$ |

 $\Gamma(e^- e^+ e^-)/\Gamma_{\text{total}}$ Γ_{206}/Γ

Test of lepton family number conservation.

| <u>VALUE</u> | <u>CL%</u> | <u>DOCUMENT ID</u> | <u>TECN</u> | <u>COMMENT</u> |
|---|------------|-----------------------|-------------|--|
| $< 2.7 \times 10^{-8}$ | 90 | HAYASAKA | 10 | BELL $782 \text{ fb}^{-1} E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$ |
| $\bullet \bullet \bullet$ We do not use the following data for averages, fits, limits, etc. $\bullet \bullet \bullet$ | | | | |
| $< 2.9 \times 10^{-8}$ | 90 | LEES | 10A | BABR $468 \text{ fb}^{-1} E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$ |
| $< 3.6 \times 10^{-8}$ | 90 | MIYAZAKI | 08 | BELL $535 \text{ fb}^{-1} E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$ |
| $< 4.3 \times 10^{-8}$ | 90 | AUBERT | 07BK | BABR $376 \text{ fb}^{-1} E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$ |
| $< 2.0 \times 10^{-7}$ | 90 | AUBERT | 04J | BABR $91.5 \text{ fb}^{-1} E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$ |
| $< 3.5 \times 10^{-7}$ | 90 | YUSA | 04 | BELL $87.1 \text{ fb}^{-1} E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$ |
| $< 2.9 \times 10^{-6}$ | 90 | BLISS | 98 | CLEO $E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$ |
| $< 0.33 \times 10^{-5}$ | 90 | ¹ BARTELTT | 94 | CLEO Repl. by BLISS 98 |
| $< 1.3 \times 10^{-5}$ | 90 | ALBRECHT | 92K | ARG $E_{\text{cm}}^{\text{ee}} = 10 \text{ GeV}$ |
| $< 2.7 \times 10^{-5}$ | 90 | BOWCOCK | 90 | CLEO $E_{\text{cm}}^{\text{ee}} = 10.4\text{--}10.9$ |
| $< 40 \times 10^{-5}$ | 90 | HAYES | 82 | MRK2 $E_{\text{cm}}^{\text{ee}} = 3.8\text{--}6.8 \text{ GeV}$ |

¹ BARTELTT 94 assume phase space decays. $\Gamma(e^- \mu^+ \mu^-)/\Gamma_{\text{total}}$ Γ_{207}/Γ

Test of lepton family number conservation.

| <u>VALUE</u> | <u>CL%</u> | <u>DOCUMENT ID</u> | <u>TECN</u> | <u>COMMENT</u> |
|---|------------|-----------------------|-------------|--|
| $< 2.7 \times 10^{-8}$ | 90 | HAYASAKA | 10 | BELL $782 \text{ fb}^{-1} E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$ |
| $\bullet \bullet \bullet$ We do not use the following data for averages, fits, limits, etc. $\bullet \bullet \bullet$ | | | | |
| $< 3.2 \times 10^{-8}$ | 90 | LEES | 10A | BABR $468 \text{ fb}^{-1} E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$ |
| $< 4.1 \times 10^{-8}$ | 90 | MIYAZAKI | 08 | BELL $535 \text{ fb}^{-1} E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$ |
| $< 3.7 \times 10^{-8}$ | 90 | AUBERT | 07BK | BABR $376 \text{ fb}^{-1} E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$ |
| $< 3.3 \times 10^{-7}$ | 90 | AUBERT | 04J | BABR $91.5 \text{ fb}^{-1} E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$ |
| $< 2.0 \times 10^{-7}$ | 90 | YUSA | 04 | BELL $87.1 \text{ fb}^{-1} E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$ |
| $< 1.8 \times 10^{-6}$ | 90 | BLISS | 98 | CLEO $E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$ |
| $< 0.36 \times 10^{-5}$ | 90 | ¹ BARTELTT | 94 | CLEO Repl. by BLISS 98 |
| $< 1.9 \times 10^{-5}$ | 90 | ALBRECHT | 92K | ARG $E_{\text{cm}}^{\text{ee}} = 10 \text{ GeV}$ |
| $< 2.7 \times 10^{-5}$ | 90 | BOWCOCK | 90 | CLEO $E_{\text{cm}}^{\text{ee}} = 10.4\text{--}10.9$ |
| $< 33 \times 10^{-5}$ | 90 | HAYES | 82 | MRK2 $E_{\text{cm}}^{\text{ee}} = 3.8\text{--}6.8 \text{ GeV}$ |

¹ BARTELTT 94 assume phase space decays.

$\Gamma(e^+ \mu^- \mu^-)/\Gamma_{\text{total}}$ Γ_{208}/Γ

Test of lepton family number conservation.

| VALUE | CL% | DOCUMENT ID | TECN | COMMENT |
|---|-----|-----------------------|------|--|
| $< 1.7 \times 10^{-8}$ | 90 | HAYASAKA | 10 | BELL 782 fb^{-1} $E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$ |
| $\bullet \bullet \bullet$ We do not use the following data for averages, fits, limits, etc. $\bullet \bullet \bullet$ | | | | |
| $< 2.6 \times 10^{-8}$ | 90 | LEES | 10A | BABR 468 fb^{-1} $E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$ |
| $< 2.3 \times 10^{-8}$ | 90 | MIYAZAKI | 08 | BELL 535 fb^{-1} $E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$ |
| $< 5.6 \times 10^{-8}$ | 90 | AUBERT | 07BK | BABR 376 fb^{-1} $E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$ |
| $< 1.3 \times 10^{-7}$ | 90 | AUBERT | 04J | BABR 91.5 fb^{-1} $E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$ |
| $< 2.0 \times 10^{-7}$ | 90 | YUSA | 04 | BELL 87.1 fb^{-1} $E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$ |
| $< 1.5 \times 10^{-6}$ | 90 | BLISS | 98 | CLEO $E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$ |
| $< 0.35 \times 10^{-5}$ | 90 | ¹ BARTELTT | 94 | CLEO Repl. by BLISS 98 |
| $< 1.8 \times 10^{-5}$ | 90 | ALBRECHT | 92K | ARG $E_{\text{cm}}^{\text{ee}} = 10 \text{ GeV}$ |
| $< 1.6 \times 10^{-5}$ | 90 | BOWCOCK | 90 | CLEO $E_{\text{cm}}^{\text{ee}} = 10.4\text{--}10.9$ |

¹ BARTELTT 94 assume phase space decays. $\Gamma(\mu^- e^+ e^-)/\Gamma_{\text{total}}$ Γ_{209}/Γ

Test of lepton family number conservation.

| VALUE | CL% | DOCUMENT ID | TECN | COMMENT |
|---|-----|-----------------------|------|--|
| $< 1.8 \times 10^{-8}$ | 90 | HAYASAKA | 10 | BELL 782 fb^{-1} $E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$ |
| $\bullet \bullet \bullet$ We do not use the following data for averages, fits, limits, etc. $\bullet \bullet \bullet$ | | | | |
| $< 2.2 \times 10^{-8}$ | 90 | LEES | 10A | BABR 468 fb^{-1} $E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$ |
| $< 2.7 \times 10^{-8}$ | 90 | MIYAZAKI | 08 | BELL 535 fb^{-1} $E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$ |
| $< 8.0 \times 10^{-8}$ | 90 | AUBERT | 07BK | BABR 376 fb^{-1} $E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$ |
| $< 2.7 \times 10^{-7}$ | 90 | AUBERT | 04J | BABR 91.5 fb^{-1} $E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$ |
| $< 1.9 \times 10^{-7}$ | 90 | YUSA | 04 | BELL 87.1 fb^{-1} $E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$ |
| $< 1.7 \times 10^{-6}$ | 90 | BLISS | 98 | CLEO $E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$ |
| $< 0.34 \times 10^{-5}$ | 90 | ¹ BARTELTT | 94 | CLEO Repl. by BLISS 98 |
| $< 1.4 \times 10^{-5}$ | 90 | ALBRECHT | 92K | ARG $E_{\text{cm}}^{\text{ee}} = 10 \text{ GeV}$ |
| $< 2.7 \times 10^{-5}$ | 90 | BOWCOCK | 90 | CLEO $E_{\text{cm}}^{\text{ee}} = 10.4\text{--}10.9$ |
| $< 44 \times 10^{-5}$ | 90 | HAYES | 82 | MRK2 $E_{\text{cm}}^{\text{ee}} = 3.8\text{--}6.8 \text{ GeV}$ |

¹ BARTELTT 94 assume phase space decays. $\Gamma(\mu^+ e^- e^-)/\Gamma_{\text{total}}$ Γ_{210}/Γ

Test of lepton family number conservation.

| VALUE | CL% | DOCUMENT ID | TECN | COMMENT |
|---|-----|-----------------------|------|--|
| $< 1.5 \times 10^{-8}$ | 90 | HAYASAKA | 10 | BELL 782 fb^{-1} $E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$ |
| $\bullet \bullet \bullet$ We do not use the following data for averages, fits, limits, etc. $\bullet \bullet \bullet$ | | | | |
| $< 1.8 \times 10^{-8}$ | 90 | LEES | 10A | BABR 468 fb^{-1} $E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$ |
| $< 2.0 \times 10^{-8}$ | 90 | MIYAZAKI | 08 | BELL 535 fb^{-1} $E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$ |
| $< 5.8 \times 10^{-8}$ | 90 | AUBERT | 07BK | BABR 376 fb^{-1} $E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$ |
| $< 1.1 \times 10^{-7}$ | 90 | AUBERT | 04J | BABR 91.5 fb^{-1} $E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$ |
| $< 2.0 \times 10^{-7}$ | 90 | YUSA | 04 | BELL 87.1 fb^{-1} $E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$ |
| $< 1.5 \times 10^{-6}$ | 90 | BLISS | 98 | CLEO $E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$ |
| $< 0.34 \times 10^{-5}$ | 90 | ¹ BARTELTT | 94 | CLEO Repl. by BLISS 98 |
| $< 1.4 \times 10^{-5}$ | 90 | ALBRECHT | 92K | ARG $E_{\text{cm}}^{\text{ee}} = 10 \text{ GeV}$ |
| $< 1.6 \times 10^{-5}$ | 90 | BOWCOCK | 90 | CLEO $E_{\text{cm}}^{\text{ee}} = 10.4\text{--}10.9$ |

¹ BARTELT 94 assume phase space decays.

$\Gamma(\mu^- \mu^+ \mu^-)/\Gamma_{\text{total}}$

Test of lepton family number conservation.

Γ_{211}/Γ

| VALUE | CL% | DOCUMENT ID | TECN | COMMENT |
|---|-----|----------------------|------|--|
| $< 2.1 \times 10^{-8}$ | 90 | HAYASAKA | 10 | BELL $782 \text{ fb}^{-1} E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$ |
| $\bullet \bullet \bullet$ We do not use the following data for averages, fits, limits, etc. $\bullet \bullet \bullet$ | | | | |
| $< 4.6 \times 10^{-8}$ | 90 | AAIJ | 15AI | LHCb $3.0 \text{ fb}^{-1} \sqrt{s} = 7, 8 \text{ TeV}$ |
| $< 8.0 \times 10^{-8}$ | 90 | ¹ AAIJ | 13AH | LHCb $1.0 \text{ fb}^{-1}, \sqrt{s} = 7 \text{ TeV}$ |
| $< 3.3 \times 10^{-8}$ | 90 | LEES | 10A | BABR $468 \text{ fb}^{-1} E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$ |
| $< 3.2 \times 10^{-8}$ | 90 | MIYAZAKI | 08 | BELL $535 \text{ fb}^{-1} E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$ |
| $< 5.3 \times 10^{-8}$ | 90 | AUBERT | 07BK | BABR $376 \text{ fb}^{-1} E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$ |
| $< 1.9 \times 10^{-7}$ | 90 | AUBERT | 04J | BABR $91.5 \text{ fb}^{-1} E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$ |
| $< 2.0 \times 10^{-7}$ | 90 | YUSA | 04 | BELL $87.1 \text{ fb}^{-1} E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$ |
| $< 1.9 \times 10^{-6}$ | 90 | BLISS | 98 | CLEO $E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$ |
| $< 0.43 \times 10^{-5}$ | 90 | ² BARTELT | 94 | CLEO Repl. by BLISS 98 |
| $< 1.9 \times 10^{-5}$ | 90 | ALBRECHT | 92K | ARG $E_{\text{cm}}^{\text{ee}} = 10 \text{ GeV}$ |
| $< 1.7 \times 10^{-5}$ | 90 | BOWCOCK | 90 | CLEO $E_{\text{cm}}^{\text{ee}} = 10.4\text{--}10.9$ |
| $< 49 \times 10^{-5}$ | 90 | HAYES | 82 | MRK2 $E_{\text{cm}}^{\text{ee}} = 3.8\text{--}6.8 \text{ GeV}$ |

¹ Repl. by AAIJ 15AI.

² BARTELT 94 assume phase space decays.

$\Gamma(e^- \pi^+ \pi^-)/\Gamma_{\text{total}}$

Test of lepton family number conservation.

Γ_{212}/Γ

| VALUE | CL% | DOCUMENT ID | TECN | COMMENT |
|---|-----|----------------------|------|--|
| $< 2.3 \times 10^{-8}$ | 90 | MIYAZAKI | 13 | BELL $854 \text{ fb}^{-1} E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$ |
| $\bullet \bullet \bullet$ We do not use the following data for averages, fits, limits, etc. $\bullet \bullet \bullet$ | | | | |
| $< 4.4 \times 10^{-8}$ | 90 | MIYAZAKI | 10 | BELL Repl. by MIYAZAKI 13 |
| $< 7.3 \times 10^{-7}$ | 90 | YUSA | 06 | BELL $158 \text{ fb}^{-1} E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$ |
| $< 1.2 \times 10^{-7}$ | 90 | AUBERT,BE | 05D | BABR $221 \text{ fb}^{-1}, E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$ |
| $< 2.2 \times 10^{-6}$ | 90 | BLISS | 98 | CLEO $E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$ |
| $< 4.4 \times 10^{-6}$ | 90 | ¹ BARTELT | 94 | CLEO Repl. by BLISS 98 |
| $< 2.7 \times 10^{-5}$ | 90 | ALBRECHT | 92K | ARG $E_{\text{cm}}^{\text{ee}} = 10 \text{ GeV}$ |
| $< 6.0 \times 10^{-5}$ | 90 | BOWCOCK | 90 | CLEO $E_{\text{cm}}^{\text{ee}} = 10.4\text{--}10.9$ |

¹ BARTELT 94 assume phase space decays.

$\Gamma(e^+ \pi^- \pi^-)/\Gamma_{\text{total}}$

Test of lepton number conservation.

Γ_{213}/Γ

| VALUE | CL% | DOCUMENT ID | TECN | COMMENT |
|---|-----|----------------------|------|--|
| $< 2.0 \times 10^{-8}$ | 90 | MIYAZAKI | 13 | BELL $854 \text{ fb}^{-1} E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$ |
| $\bullet \bullet \bullet$ We do not use the following data for averages, fits, limits, etc. $\bullet \bullet \bullet$ | | | | |
| $< 8.8 \times 10^{-8}$ | 90 | MIYAZAKI | 10 | BELL Repl. by MIYAZAKI 13 |
| $< 2.0 \times 10^{-7}$ | 90 | YUSA | 06 | BELL $158 \text{ fb}^{-1} E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$ |
| $< 2.7 \times 10^{-7}$ | 90 | AUBERT,BE | 05D | BABR $221 \text{ fb}^{-1}, E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$ |
| $< 1.9 \times 10^{-6}$ | 90 | BLISS | 98 | CLEO $E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$ |
| $< 4.4 \times 10^{-6}$ | 90 | ¹ BARTELT | 94 | CLEO Repl. by BLISS 98 |
| $< 1.8 \times 10^{-5}$ | 90 | ALBRECHT | 92K | ARG $E_{\text{cm}}^{\text{ee}} = 10 \text{ GeV}$ |
| $< 1.7 \times 10^{-5}$ | 90 | BOWCOCK | 90 | CLEO $E_{\text{cm}}^{\text{ee}} = 10.4\text{--}10.9$ |

¹ BARTELT 94 assume phase space decays.

$\Gamma(\mu^-\pi^+\pi^-)/\Gamma_{\text{total}}$ Γ_{214}/Γ

Test of lepton family number conservation.

| VALUE | CL% | DOCUMENT ID | TECN | COMMENT |
|---|-----|----------------------|------|--|
| $<2.1 \times 10^{-8}$ | 90 | MIYAZAKI | 13 | BELL $854 \text{ fb}^{-1} E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$ |
| $\bullet \bullet \bullet$ We do not use the following data for averages, fits, limits, etc. $\bullet \bullet \bullet$ | | | | |
| $<3.3 \times 10^{-8}$ | 90 | MIYAZAKI | 10 | BELL Repl. by MIYAZAKI 13 |
| $<4.8 \times 10^{-7}$ | 90 | YUSA | 06 | BELL $158 \text{ fb}^{-1} E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$ |
| $<2.9 \times 10^{-7}$ | 90 | AUBERT,BE | 05D | BABR $221 \text{ fb}^{-1}, E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$ |
| $<8.2 \times 10^{-6}$ | 90 | BLISS | 98 | CLEO $E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$ |
| $<7.4 \times 10^{-6}$ | 90 | ¹ BARTELT | 94 | CLEO Repl. by BLISS 98 |
| $<3.6 \times 10^{-5}$ | 90 | ALBRECHT | 92K | ARG $E_{\text{cm}}^{\text{ee}} = 10 \text{ GeV}$ |
| $<3.9 \times 10^{-5}$ | 90 | BOWCOCK | 90 | CLEO $E_{\text{cm}}^{\text{ee}} = 10.4\text{--}10.9$ |

¹ BARTELT 94 assume phase space decays.

$\Gamma(\mu^+\pi^-\pi^-)/\Gamma_{\text{total}}$ Γ_{215}/Γ

Test of lepton number conservation.

| VALUE | CL% | DOCUMENT ID | TECN | COMMENT |
|---|-----|----------------------|------|--|
| $<3.9 \times 10^{-8}$ | 90 | MIYAZAKI | 13 | BELL $854 \text{ fb}^{-1} E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$ |
| $\bullet \bullet \bullet$ We do not use the following data for averages, fits, limits, etc. $\bullet \bullet \bullet$ | | | | |
| $<3.7 \times 10^{-8}$ | 90 | MIYAZAKI | 10 | BELL Repl. by MIYAZAKI 13 |
| $<3.4 \times 10^{-7}$ | 90 | YUSA | 06 | BELL $158 \text{ fb}^{-1} E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$ |
| $<7 \times 10^{-8}$ | 90 | AUBERT,BE | 05D | BABR $221 \text{ fb}^{-1}, E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$ |
| $<3.4 \times 10^{-6}$ | 90 | BLISS | 98 | CLEO $E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$ |
| $<6.9 \times 10^{-6}$ | 90 | ¹ BARTELT | 94 | CLEO Repl. by BLISS 98 |
| $<6.3 \times 10^{-5}$ | 90 | ALBRECHT | 92K | ARG $E_{\text{cm}}^{\text{ee}} = 10 \text{ GeV}$ |
| $<3.9 \times 10^{-5}$ | 90 | BOWCOCK | 90 | CLEO $E_{\text{cm}}^{\text{ee}} = 10.4\text{--}10.9$ |

¹ BARTELT 94 assume phase space decays.

$\Gamma(e^-\pi^+K^-)/\Gamma_{\text{total}}$ Γ_{216}/Γ

Test of lepton family number conservation.

| VALUE | CL% | DOCUMENT ID | TECN | COMMENT |
|---|-----|----------------------|------|--|
| $<3.7 \times 10^{-8}$ | 90 | MIYAZAKI | 13 | BELL $854 \text{ fb}^{-1} E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$ |
| $\bullet \bullet \bullet$ We do not use the following data for averages, fits, limits, etc. $\bullet \bullet \bullet$ | | | | |
| $<5.8 \times 10^{-8}$ | 90 | MIYAZAKI | 10 | BELL Repl. by MIYAZAKI 13 |
| $<7.2 \times 10^{-7}$ | 90 | YUSA | 06 | BELL $158 \text{ fb}^{-1} E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$ |
| $<3.2 \times 10^{-7}$ | 90 | AUBERT,BE | 05D | BABR $221 \text{ fb}^{-1}, E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$ |
| $<6.4 \times 10^{-6}$ | 90 | BLISS | 98 | CLEO $E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$ |
| $<7.7 \times 10^{-6}$ | 90 | ¹ BARTELT | 94 | CLEO Repl. by BLISS 98 |
| $<2.9 \times 10^{-5}$ | 90 | ALBRECHT | 92K | ARG $E_{\text{cm}}^{\text{ee}} = 10 \text{ GeV}$ |
| $<5.8 \times 10^{-5}$ | 90 | BOWCOCK | 90 | CLEO $E_{\text{cm}}^{\text{ee}} = 10.4\text{--}10.9$ |

¹ BARTELT 94 assume phase space decays.

$\Gamma(e^- \pi^- K^+)/\Gamma_{\text{total}}$ Γ_{217}/Γ

Test of lepton family number conservation.

| VALUE | CL% | DOCUMENT ID | TECN | COMMENT |
|---|-----|---------------------|------|--|
| $<3.1 \times 10^{-8}$ | 90 | MIYAZAKI | 13 | BELL $854 \text{ fb}^{-1} E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$ |
| $\bullet \bullet \bullet$ We do not use the following data for averages, fits, limits, etc. $\bullet \bullet \bullet$ | | | | |
| $<5.2 \times 10^{-8}$ | 90 | MIYAZAKI | 10 | BELL Repl. by MIYAZAKI 13 |
| $<1.6 \times 10^{-7}$ | 90 | YUSA | 06 | BELL $158 \text{ fb}^{-1} E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$ |
| $<1.7 \times 10^{-7}$ | 90 | AUBERT,BE | 05D | BABR $221 \text{ fb}^{-1}, E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$ |
| $<3.8 \times 10^{-6}$ | 90 | BLISS | 98 | CLEO $E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$ |
| $<4.6 \times 10^{-6}$ | 90 | ¹ BARTEL | 94 | CLEO Repl. by BLISS 98 |
| $<5.8 \times 10^{-5}$ | 90 | BOWCOCK | 90 | CLEO $E_{\text{cm}}^{\text{ee}} = 10.4\text{--}10.9$ |

¹ BARTEL 94 assume phase space decays. $\Gamma(e^+ \pi^- K^-)/\Gamma_{\text{total}}$ Γ_{218}/Γ

Test of lepton number conservation.

| VALUE | CL% | DOCUMENT ID | TECN | COMMENT |
|---|-----|---------------------|------|--|
| $<3.2 \times 10^{-8}$ | 90 | MIYAZAKI | 13 | BELL $854 \text{ fb}^{-1} E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$ |
| $\bullet \bullet \bullet$ We do not use the following data for averages, fits, limits, etc. $\bullet \bullet \bullet$ | | | | |
| $<6.7 \times 10^{-8}$ | 90 | MIYAZAKI | 10 | BELL Repl. by MIYAZAKI 13 |
| $<1.9 \times 10^{-7}$ | 90 | YUSA | 06 | BELL $158 \text{ fb}^{-1} E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$ |
| $<1.8 \times 10^{-7}$ | 90 | AUBERT,BE | 05D | BABR $221 \text{ fb}^{-1}, E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$ |
| $<2.1 \times 10^{-6}$ | 90 | BLISS | 98 | CLEO $E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$ |
| $<4.5 \times 10^{-6}$ | 90 | ¹ BARTEL | 94 | CLEO Repl. by BLISS 98 |
| $<2.0 \times 10^{-5}$ | 90 | ALBRECHT | 92K | ARG $E_{\text{cm}}^{\text{ee}} = 10 \text{ GeV}$ |
| $<4.9 \times 10^{-5}$ | 90 | BOWCOCK | 90 | CLEO $E_{\text{cm}}^{\text{ee}} = 10.4\text{--}10.9$ |

¹ BARTEL 94 assume phase space decays. $\Gamma(e^- K_S^0 K_S^0)/\Gamma_{\text{total}}$ Γ_{219}/Γ

Test of lepton family number conservation.

| VALUE | CL% | DOCUMENT ID | TECN | COMMENT |
|---|-----|-------------|------|---|
| $<7.1 \times 10^{-8}$ | 90 | MIYAZAKI | 10A | BELL $671 \text{ fb}^{-1} E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$ |
| $\bullet \bullet \bullet$ We do not use the following data for averages, fits, limits, etc. $\bullet \bullet \bullet$ | | | | |
| $<2.2 \times 10^{-6}$ | 90 | CHEN | 02C | CLEO $E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$ |

 $\Gamma(e^- K^+ K^-)/\Gamma_{\text{total}}$ Γ_{220}/Γ

Test of lepton family number conservation.

| VALUE | CL% | DOCUMENT ID | TECN | COMMENT |
|---|-----|-------------|------|--|
| $<3.4 \times 10^{-8}$ | 90 | MIYAZAKI | 13 | BELL $854 \text{ fb}^{-1} E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$ |
| $\bullet \bullet \bullet$ We do not use the following data for averages, fits, limits, etc. $\bullet \bullet \bullet$ | | | | |
| $<5.4 \times 10^{-8}$ | 90 | MIYAZAKI | 10 | BELL Repl. by MIYAZAKI 13 |
| $<3.0 \times 10^{-7}$ | 90 | YUSA | 06 | BELL $158 \text{ fb}^{-1} E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$ |
| $<1.4 \times 10^{-7}$ | 90 | AUBERT,BE | 05D | BABR $221 \text{ fb}^{-1}, E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$ |
| $<6.0 \times 10^{-6}$ | 90 | BLISS | 98 | CLEO $E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$ |

$\Gamma(e^+ K^- K^-)/\Gamma_{\text{total}}$ Γ_{221}/Γ

Test of lepton number conservation.

| VALUE | CL% | DOCUMENT ID | TECN | COMMENT |
|---|-----|-------------|------|--|
| $<3.3 \times 10^{-8}$ | 90 | MIYAZAKI | 13 | BELL $854 \text{ fb}^{-1} E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$ |
| $\bullet \bullet \bullet$ We do not use the following data for averages, fits, limits, etc. $\bullet \bullet \bullet$ | | | | |
| $<6.0 \times 10^{-8}$ | 90 | MIYAZAKI | 10 | BELL Repl. by MIYAZAKI 13 |
| $<3.1 \times 10^{-7}$ | 90 | YUSA | 06 | BELL $158 \text{ fb}^{-1} E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$ |
| $<1.5 \times 10^{-7}$ | 90 | AUBERT,BE | 05D | BABR $221 \text{ fb}^{-1}, E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$ |
| $<3.8 \times 10^{-6}$ | 90 | BLISS | 98 | CLEO $E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$ |

 $\Gamma(\mu^- \pi^+ K^-)/\Gamma_{\text{total}}$ Γ_{222}/Γ

Test of lepton family number conservation.

| VALUE | CL% | DOCUMENT ID | TECN | COMMENT |
|---|-----|---------------------|------|--|
| $< 8.6 \times 10^{-8}$ | 90 | MIYAZAKI | 13 | BELL $854 \text{ fb}^{-1} E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$ |
| $\bullet \bullet \bullet$ We do not use the following data for averages, fits, limits, etc. $\bullet \bullet \bullet$ | | | | |
| $< 1.6 \times 10^{-7}$ | 90 | MIYAZAKI | 10 | BELL Repl. by MIYAZAKI 13 |
| $< 2.7 \times 10^{-7}$ | 90 | YUSA | 06 | BELL $158 \text{ fb}^{-1} E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$ |
| $< 2.6 \times 10^{-7}$ | 90 | AUBERT,BE | 05D | BABR $221 \text{ fb}^{-1}, E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$ |
| $< 7.5 \times 10^{-6}$ | 90 | BLISS | 98 | CLEO $E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$ |
| $< 8.7 \times 10^{-6}$ | 90 | ¹ BARTEL | 94 | CLEO Repl. by BLISS 98 |
| $< 11 \times 10^{-5}$ | 90 | ALBRECHT | 92K | ARG $E_{\text{cm}}^{\text{ee}} = 10 \text{ GeV}$ |
| $< 7.7 \times 10^{-5}$ | 90 | BOWCOCK | 90 | CLEO $E_{\text{cm}}^{\text{ee}} = 10.4\text{--}10.9$ |

¹ BARTEL 94 assume phase space decays.
 $\Gamma(\mu^- \pi^- K^+)/\Gamma_{\text{total}}$ Γ_{223}/Γ

Test of lepton family number conservation.

| VALUE | CL% | DOCUMENT ID | TECN | COMMENT |
|---|-----|---------------------|------|--|
| $< 4.5 \times 10^{-8}$ | 90 | MIYAZAKI | 13 | BELL $854 \text{ fb}^{-1} E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$ |
| $\bullet \bullet \bullet$ We do not use the following data for averages, fits, limits, etc. $\bullet \bullet \bullet$ | | | | |
| $< 1.0 \times 10^{-7}$ | 90 | MIYAZAKI | 10 | BELL Repl. by MIYAZAKI 13 |
| $< 7.3 \times 10^{-7}$ | 90 | YUSA | 06 | BELL $158 \text{ fb}^{-1} E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$ |
| $< 3.2 \times 10^{-7}$ | 90 | AUBERT,BE | 05D | BABR $221 \text{ fb}^{-1}, E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$ |
| $< 7.4 \times 10^{-6}$ | 90 | BLISS | 98 | CLEO $E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$ |
| $< 1.5 \times 10^{-5}$ | 90 | ¹ BARTEL | 94 | CLEO Repl. by BLISS 98 |
| $< 7.7 \times 10^{-5}$ | 90 | BOWCOCK | 90 | CLEO $E_{\text{cm}}^{\text{ee}} = 10.4\text{--}10.9$ |

¹ BARTEL 94 assume phase space decays.
 $\Gamma(\mu^+ \pi^- K^-)/\Gamma_{\text{total}}$ Γ_{224}/Γ

Test of lepton number conservation.

| VALUE | CL% | DOCUMENT ID | TECN | COMMENT |
|---|-----|---------------------|------|--|
| $< 4.8 \times 10^{-8}$ | 90 | MIYAZAKI | 13 | BELL $854 \text{ fb}^{-1} E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$ |
| $\bullet \bullet \bullet$ We do not use the following data for averages, fits, limits, etc. $\bullet \bullet \bullet$ | | | | |
| $< 9.4 \times 10^{-8}$ | 90 | MIYAZAKI | 10 | BELL Repl. by MIYAZAKI 13 |
| $< 2.9 \times 10^{-7}$ | 90 | YUSA | 06 | BELL $158 \text{ fb}^{-1} E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$ |
| $< 2.2 \times 10^{-7}$ | 90 | AUBERT,BE | 05D | BABR $221 \text{ fb}^{-1}, E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$ |
| $< 7.0 \times 10^{-6}$ | 90 | BLISS | 98 | CLEO $E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$ |
| $< 2.0 \times 10^{-5}$ | 90 | ¹ BARTEL | 94 | CLEO Repl. by BLISS 98 |
| $< 5.8 \times 10^{-5}$ | 90 | ALBRECHT | 92K | ARG $E_{\text{cm}}^{\text{ee}} = 10 \text{ GeV}$ |
| $< 4.0 \times 10^{-5}$ | 90 | BOWCOCK | 90 | CLEO $E_{\text{cm}}^{\text{ee}} = 10.4\text{--}10.9$ |

¹ BARTELT 94 assume phase space decays.

$\Gamma(\mu^- K_S^0 K_S^0)/\Gamma_{\text{total}}$ Γ_{225}/Γ

| VALUE | CL% | DOCUMENT ID | TECN | COMMENT |
|---|-----|-------------|------|---|
| $< 8.0 \times 10^{-8}$ | 90 | MIYAZAKI | 10A | BELL 671 fb^{-1} $E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$ |
| $\bullet \bullet \bullet$ We do not use the following data for averages, fits, limits, etc. $\bullet \bullet \bullet$ | | | | |
| $< 3.4 \times 10^{-6}$ | 90 | CHEN | 02C | CLEO $E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$ |

$\Gamma(\mu^- K^+ K^-)/\Gamma_{\text{total}}$ Γ_{226}/Γ

Test of lepton family number conservation.

| VALUE | CL% | DOCUMENT ID | TECN | COMMENT |
|---|-----|-------------|------|---|
| $< 4.4 \times 10^{-8}$ | 90 | MIYAZAKI | 13 | BELL 854 fb^{-1} $E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$ |
| $\bullet \bullet \bullet$ We do not use the following data for averages, fits, limits, etc. $\bullet \bullet \bullet$ | | | | |
| $< 6.8 \times 10^{-8}$ | 90 | MIYAZAKI | 10 | BELL Repl. by MIYAZAKI 13 |
| $< 8.0 \times 10^{-7}$ | 90 | YUSA | 06 | BELL 158 fb^{-1} $E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$ |
| $< 2.5 \times 10^{-7}$ | 90 | AUBERT,BE | 05D | BABR 221 fb^{-1} , $E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$ |
| $< 15 \times 10^{-6}$ | 90 | BLISS | 98 | CLEO $E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$ |

$\Gamma(\mu^+ K^- K^-)/\Gamma_{\text{total}}$ Γ_{227}/Γ

Test of lepton number conservation.

| VALUE | CL% | DOCUMENT ID | TECN | COMMENT |
|---|-----|-------------|------|---|
| $< 4.7 \times 10^{-8}$ | 90 | MIYAZAKI | 13 | BELL 854 fb^{-1} $E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$ |
| $\bullet \bullet \bullet$ We do not use the following data for averages, fits, limits, etc. $\bullet \bullet \bullet$ | | | | |
| $< 9.6 \times 10^{-8}$ | 90 | MIYAZAKI | 10 | BELL Repl. by MIYAZAKI 13 |
| $< 4.4 \times 10^{-7}$ | 90 | YUSA | 06 | BELL 158 fb^{-1} $E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$ |
| $< 4.8 \times 10^{-7}$ | 90 | AUBERT,BE | 05D | BABR 221 fb^{-1} , $E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$ |
| $< 6.0 \times 10^{-6}$ | 90 | BLISS | 98 | CLEO $E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$ |

$\Gamma(e^- \pi^0 \pi^0)/\Gamma_{\text{total}}$ Γ_{228}/Γ

Test of lepton family number conservation.

| VALUE | CL% | DOCUMENT ID | TECN | COMMENT |
|------------------------|-----|-------------|------|---|
| $< 6.5 \times 10^{-6}$ | 90 | BONVICINI | 97 | CLEO $E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$ |

$\Gamma(\mu^- \pi^0 \pi^0)/\Gamma_{\text{total}}$ Γ_{229}/Γ

Test of lepton family number conservation.

| VALUE | CL% | DOCUMENT ID | TECN | COMMENT |
|-----------------------|-----|-------------|------|---|
| $< 14 \times 10^{-6}$ | 90 | BONVICINI | 97 | CLEO $E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$ |

$\Gamma(e^- \eta \eta)/\Gamma_{\text{total}}$ Γ_{230}/Γ

Test of lepton family number conservation.

| VALUE | CL% | DOCUMENT ID | TECN | COMMENT |
|-----------------------|-----|-------------|------|---|
| $< 35 \times 10^{-6}$ | 90 | BONVICINI | 97 | CLEO $E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$ |

$\Gamma(\mu^- \eta \eta)/\Gamma_{\text{total}}$ Γ_{231}/Γ

Test of lepton family number conservation.

| VALUE | CL% | DOCUMENT ID | TECN | COMMENT |
|-----------------------|-----|-------------|------|---|
| $< 60 \times 10^{-6}$ | 90 | BONVICINI | 97 | CLEO $E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$ |

$\Gamma(e^- \pi^0 \eta)/\Gamma_{\text{total}}$ Γ_{232}/Γ

Test of lepton family number conservation.

| <u>VALUE</u> | <u>CL%</u> | <u>DOCUMENT ID</u> | <u>TECN</u> | <u>COMMENT</u> |
|-----------------------|------------|--------------------|-------------|--|
| $< 24 \times 10^{-6}$ | 90 | BONVICINI | 97 | CLEO $E_{\text{cm}}^{ee} = 10.6 \text{ GeV}$ |

 $\Gamma(\mu^- \pi^0 \eta)/\Gamma_{\text{total}}$ Γ_{233}/Γ

Test of lepton family number conservation.

| <u>VALUE</u> | <u>CL%</u> | <u>DOCUMENT ID</u> | <u>TECN</u> | <u>COMMENT</u> |
|-----------------------|------------|--------------------|-------------|--|
| $< 22 \times 10^{-6}$ | 90 | BONVICINI | 97 | CLEO $E_{\text{cm}}^{ee} = 10.6 \text{ GeV}$ |

 $\Gamma(p \mu^- \mu^-)/\Gamma_{\text{total}}$ Γ_{234}/Γ

| <u>VALUE</u> | <u>CL%</u> | <u>DOCUMENT ID</u> | <u>TECN</u> | <u>COMMENT</u> |
|------------------------|------------|--------------------|-------------|---|
| $< 4.4 \times 10^{-7}$ | 90 | AAIJ | 13AH LHCb | $1.0 \text{ fb}^{-1}, \sqrt{s} = 7 \text{ TeV}$ |

 $\Gamma(\bar{p} \mu^+ \mu^-)/\Gamma_{\text{total}}$ Γ_{235}/Γ

| <u>VALUE</u> | <u>CL%</u> | <u>DOCUMENT ID</u> | <u>TECN</u> | <u>COMMENT</u> |
|------------------------|------------|--------------------|-------------|---|
| $< 3.3 \times 10^{-7}$ | 90 | AAIJ | 13AH LHCb | $1.0 \text{ fb}^{-1}, \sqrt{s} = 7 \text{ TeV}$ |

 $\Gamma(\bar{p} \gamma)/\Gamma_{\text{total}}$ Γ_{236}/Γ

Test of lepton number and baryon number conservation.

| <u>VALUE</u> | <u>CL%</u> | <u>DOCUMENT ID</u> | <u>TECN</u> | <u>COMMENT</u> |
|---|------------|--------------------|-------------|--|
| $< 3.5 \times 10^{-6}$ | 90 | GODANG | 99 | CLEO $E_{\text{cm}}^{ee} = 10.6 \text{ GeV}$ |
| $\bullet \bullet \bullet$ We do not use the following data for averages, fits, limits, etc. $\bullet \bullet \bullet$ | | | | |
| $< 29 \times 10^{-5}$ | 90 | ALBRECHT | 92K ARG | $E_{\text{cm}}^{ee} = 10 \text{ GeV}$ |

 $\Gamma(\bar{p} \pi^0)/\Gamma_{\text{total}}$ Γ_{237}/Γ

Test of lepton number and baryon number conservation.

| <u>VALUE</u> | <u>CL%</u> | <u>DOCUMENT ID</u> | <u>TECN</u> | <u>COMMENT</u> |
|---|------------|--------------------|-------------|--|
| $< 15 \times 10^{-6}$ | 90 | GODANG | 99 | CLEO $E_{\text{cm}}^{ee} = 10.6 \text{ GeV}$ |
| $\bullet \bullet \bullet$ We do not use the following data for averages, fits, limits, etc. $\bullet \bullet \bullet$ | | | | |
| $< 66 \times 10^{-5}$ | 90 | ALBRECHT | 92K ARG | $E_{\text{cm}}^{ee} = 10 \text{ GeV}$ |

 $\Gamma(\bar{p} 2\pi^0)/\Gamma_{\text{total}}$ Γ_{238}/Γ

Test of lepton number and baryon number conservation.

| <u>VALUE</u> | <u>CL%</u> | <u>DOCUMENT ID</u> | <u>TECN</u> | <u>COMMENT</u> |
|-----------------------|------------|--------------------|-------------|--|
| $< 33 \times 10^{-6}$ | 90 | GODANG | 99 | CLEO $E_{\text{cm}}^{ee} = 10.6 \text{ GeV}$ |

 $\Gamma(\bar{p} \eta)/\Gamma_{\text{total}}$ Γ_{239}/Γ

Test of lepton number and baryon number conservation.

| <u>VALUE</u> | <u>CL%</u> | <u>DOCUMENT ID</u> | <u>TECN</u> | <u>COMMENT</u> |
|---|------------|--------------------|-------------|--|
| $< 8.9 \times 10^{-6}$ | 90 | GODANG | 99 | CLEO $E_{\text{cm}}^{ee} = 10.6 \text{ GeV}$ |
| $\bullet \bullet \bullet$ We do not use the following data for averages, fits, limits, etc. $\bullet \bullet \bullet$ | | | | |
| $< 130 \times 10^{-5}$ | 90 | ALBRECHT | 92K ARG | $E_{\text{cm}}^{ee} = 10 \text{ GeV}$ |

 $\Gamma(\bar{p} \pi^0 \eta)/\Gamma_{\text{total}}$ Γ_{240}/Γ

Test of lepton number and baryon number conservation.

| <u>VALUE</u> | <u>CL%</u> | <u>DOCUMENT ID</u> | <u>TECN</u> | <u>COMMENT</u> |
|-----------------------|------------|--------------------|-------------|--|
| $< 27 \times 10^{-6}$ | 90 | GODANG | 99 | CLEO $E_{\text{cm}}^{ee} = 10.6 \text{ GeV}$ |

$\Gamma(\Lambda\pi^-)/\Gamma_{\text{total}}$ Γ_{241}/Γ

Test of lepton number and baryon number conservation.

| <u>VALUE</u> | <u>CL%</u> | <u>DOCUMENT ID</u> | <u>TECN</u> | <u>COMMENT</u> |
|------------------------|------------|--------------------|-------------|---|
| $<0.72 \times 10^{-7}$ | 90 | MIYAZAKI | 06 | BELL 154 fb^{-1} , $E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$ |

 $\Gamma(\bar{\Lambda}\pi^-)/\Gamma_{\text{total}}$ Γ_{242}/Γ

Test of lepton number and baryon number conservation.

| <u>VALUE</u> | <u>CL%</u> | <u>DOCUMENT ID</u> | <u>TECN</u> | <u>COMMENT</u> |
|-----------------------|------------|--------------------|-------------|---|
| $<1.4 \times 10^{-7}$ | 90 | MIYAZAKI | 06 | BELL 154 fb^{-1} , $E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$ |

 $\Gamma(e^- \text{ light boson})/\Gamma(e^- \bar{\nu}_e \nu_\tau)$ Γ_{243}/Γ_5

Test of lepton family number conservation.

| <u>VALUE</u> | <u>CL%</u> | <u>DOCUMENT ID</u> | <u>TECN</u> | <u>COMMENT</u> |
|---|------------|-----------------------------|-------------|--|
| <0.015 | 95 | ¹ ALBRECHT | 95G | ARG $E_{\text{cm}}^{\text{ee}} = 9.4\text{--}10.6 \text{ GeV}$ |
| $\bullet \bullet \bullet$ We do not use the following data for averages, fits, limits, etc. $\bullet \bullet \bullet$ | | | | |
| <0.018 | 95 | ² ALBRECHT | 90E | ARG $E_{\text{cm}}^{\text{ee}} = 9.4\text{--}10.6 \text{ GeV}$ |
| <0.040 | 95 | ³ BALTRUSAIT..85 | MRK3 | $E_{\text{cm}}^{\text{ee}} = 3.77 \text{ GeV}$ |

¹ ALBRECHT 95G limit holds for bosons with mass $< 0.4 \text{ GeV}$. The limit rises to 0.036 for a mass of 1.0 GeV, then falls to 0.006 at the upper mass limit of 1.6 GeV.

² ALBRECHT 90E limit applies for spinless boson with mass $< 100 \text{ MeV}$, and rises to 0.050 for mass = 500 MeV.

³ BALTRUSAITIS 85 limit applies for spinless boson with mass $< 100 \text{ MeV}$.

 $\Gamma(\mu^- \text{ light boson})/\Gamma(e^- \bar{\nu}_e \nu_\tau)$ Γ_{244}/Γ_5

Test of lepton family number conservation.

| <u>VALUE</u> | <u>CL%</u> | <u>DOCUMENT ID</u> | <u>TECN</u> | <u>COMMENT</u> |
|---|------------|-----------------------------|-------------|--|
| <0.026 | 95 | ¹ ALBRECHT | 95G | ARG $E_{\text{cm}}^{\text{ee}} = 9.4\text{--}10.6 \text{ GeV}$ |
| $\bullet \bullet \bullet$ We do not use the following data for averages, fits, limits, etc. $\bullet \bullet \bullet$ | | | | |
| <0.033 | 95 | ² ALBRECHT | 90E | ARG $E_{\text{cm}}^{\text{ee}} = 9.4\text{--}10.6 \text{ GeV}$ |
| <0.125 | 95 | ³ BALTRUSAIT..85 | MRK3 | $E_{\text{cm}}^{\text{ee}} = 3.77 \text{ GeV}$ |

¹ ALBRECHT 95G limit holds for bosons with mass $< 1.3 \text{ GeV}$. The limit rises to 0.034 for a mass of 1.4 GeV, then falls to 0.003 at the upper mass limit of 1.6 GeV.

² ALBRECHT 90E limit applies for spinless boson with mass $< 100 \text{ MeV}$, and rises to 0.071 for mass = 500 MeV.

³ BALTRUSAITIS 85 limit applies for spinless boson with mass $< 100 \text{ MeV}$.

 τ -DECAY PARAMETERS

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 $\rho(e \text{ or } \mu)$ PARAMETER(V-A) theory predicts $\rho = 0.75$.

| <u>VALUE</u> | <u>EVTS</u> | <u>DOCUMENT ID</u> | <u>TECN</u> | <u>COMMENT</u> |
|---|-------------|-----------------------|-------------|--|
| 0.745 ± 0.008 OUR FIT | | | | |
| 0.749 ± 0.008 OUR AVERAGE | | | | |
| 0.742 $\pm 0.014 \pm 0.006$ | 81k | HEISTER | 01E | ALEP 1991–1995 LEP runs |
| 0.775 $\pm 0.023 \pm 0.020$ | 36k | ABREU | 00L | DLPH 1992–1995 runs |
| 0.781 $\pm 0.028 \pm 0.018$ | 46k | ACKERSTAFF | 99D | OPAL 1990–1995 LEP runs |
| 0.762 ± 0.035 | 54k | ACCIARRI | 98R | L3 1991–1995 LEP runs |
| 0.731 ± 0.031 | | ¹ ALBRECHT | 98 | ARG $E_{\text{cm}}^{\text{ee}} = 9.5\text{--}10.6 \text{ GeV}$ |

| | | | | | |
|--|------|-----------------------|-----|------|--|
| $0.72 \pm 0.09 \pm 0.03$ | | ² ABE | 970 | SLD | 1993–1995 SLC runs |
| $0.747 \pm 0.010 \pm 0.006$ | 55k | ALEXANDER | 97F | CLEO | $E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$ |
| $0.79 \pm 0.10 \pm 0.10$ | 3732 | FORD | 87B | MAC | $E_{\text{cm}}^{\text{ee}} = 29 \text{ GeV}$ |
| $0.71 \pm 0.09 \pm 0.03$ | 1426 | BEHREND | 85 | CLEO | $e^+ e^-$ near $\gamma(4S)$ |
| • • • We do not use the following data for averages, fits, limits, etc. • • • | | | | | |
| $0.735 \pm 0.013 \pm 0.008$ | 31k | AMMAR | 97B | CLEO | Repl. by ALEXANDER 97F |
| $0.794 \pm 0.039 \pm 0.031$ | 18k | ACCIARRI | 96H | L3 | Repl. by ACCIARRI 98R |
| $0.732 \pm 0.034 \pm 0.020$ | 8.2k | ³ ALBRECHT | 95 | ARG | $E_{\text{cm}}^{\text{ee}} = 9.5\text{--}10.6 \text{ GeV}$ |
| 0.738 ± 0.038 | | ⁴ ALBRECHT | 95C | ARG | Repl. by ALBRECHT 98 |
| $0.751 \pm 0.039 \pm 0.022$ | | BUSKULIC | 95D | ALEP | Repl. by HEISTER 01E |
| $0.742 \pm 0.035 \pm 0.020$ | 8000 | ALBRECHT | 90E | ARG | $E_{\text{cm}}^{\text{ee}} = 9.4\text{--}10.6 \text{ GeV}$ |

¹ Combined fit to ARGUS tau decay parameter measurements in ALBRECHT 98, ALBRECHT 95C, ALBRECHT 93G, and ALBRECHT 94E. ALBRECHT 98 use tau pair events of the type $\tau^- \tau^+ \rightarrow (\ell^- \bar{\nu}_\ell \nu_\tau)(\pi^+ \pi^0 \bar{\nu}_\tau)$, and their charged conjugates.

² ABE 970 assume $\eta = 0$ in their fit. Letting η vary in the fit gives a ρ value of $0.69 \pm 0.13 \pm 0.05$.

³ Value is from a simultaneous fit for the ρ and η decay parameters to the lepton energy spectrum. Not independent of ALBRECHT 90E $\rho(e$ or $\mu)$ value which assumes $\eta = 0$. Result is strongly correlated with ALBRECHT 95C.

⁴ Combined fit to ARGUS tau decay parameter measurements in ALBRECHT 95C, ALBRECHT 93G, and ALBRECHT 94E.

$\rho(e)$ PARAMETER

(V - A) theory predicts $\rho = 0.75$.

| VALUE | EVTS | DOCUMENT ID | TECN | COMMENT |
|--|------|-----------------------|------|---|
| 0.747 ± 0.010 OUR FIT | | | | |
| 0.744 ± 0.010 OUR AVERAGE | | | | |
| $0.747 \pm 0.019 \pm 0.014$ | 44k | HEISTER | 01E | ALEP 1991–1995 LEP runs |
| $0.744 \pm 0.036 \pm 0.037$ | 17k | ABREU | 00L | DLPH 1992–1995 runs |
| $0.779 \pm 0.047 \pm 0.029$ | 25k | ACKERSTAFF | 99D | OPAL 1990–1995 LEP runs |
| $0.68 \pm 0.04 \pm 0.07$ | | ¹ ALBRECHT | 98 | ARG $E_{\text{cm}}^{\text{ee}} = 9.5\text{--}10.6 \text{ GeV}$ |
| $0.71 \pm 0.14 \pm 0.05$ | | ABE | 970 | SLD 1993–1995 SLC runs |
| $0.747 \pm 0.012 \pm 0.004$ | 34k | ALEXANDER | 97F | CLEO $E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$ |
| $0.735 \pm 0.036 \pm 0.020$ | 4.7k | ² ALBRECHT | 95 | ARG $E_{\text{cm}}^{\text{ee}} = 9.5\text{--}10.6 \text{ GeV}$ |
| $0.79 \pm 0.08 \pm 0.06$ | 3230 | ³ ALBRECHT | 93G | ARG $E_{\text{cm}}^{\text{ee}} = 9.4\text{--}10.6 \text{ GeV}$ |
| $0.64 \pm 0.06 \pm 0.07$ | 2753 | JANSSEN | 89 | CBAL $E_{\text{cm}}^{\text{ee}} = 9.4\text{--}10.6 \text{ GeV}$ |
| $0.62 \pm 0.17 \pm 0.14$ | 1823 | FORD | 87B | MAC $E_{\text{cm}}^{\text{ee}} = 29 \text{ GeV}$ |
| 0.60 ± 0.13 | 699 | BEHREND | 85 | CLEO $e^+ e^-$ near $\gamma(4S)$ |
| $0.72 \pm 0.10 \pm 0.11$ | 594 | BACINO | 79B | DLCO $E_{\text{cm}}^{\text{ee}} = 3.5\text{--}7.4 \text{ GeV}$ |
| • • • We do not use the following data for averages, fits, limits, etc. • • • | | | | |
| $0.732 \pm 0.014 \pm 0.009$ | 19k | AMMAR | 97B | CLEO Repl. by ALEXANDER 97F |
| $0.793 \pm 0.050 \pm 0.025$ | | BUSKULIC | 95D | ALEP Repl. by HEISTER 01E |
| $0.747 \pm 0.045 \pm 0.028$ | 5106 | ALBRECHT | 90E | ARG Repl. by ALBRECHT 95 |

¹ ALBRECHT 98 use tau pair events of the type $\tau^- \tau^+ \rightarrow (\ell^- \bar{\nu}_\ell \nu_\tau)(\pi^+ \pi^0 \bar{\nu}_\tau)$, and their charged conjugates.

² ALBRECHT 95 use tau pair events of the type $\tau^- \tau^+ \rightarrow (\ell^- \bar{\nu}_\ell \nu_\tau)(h^+ h^- h^+ (\pi^0) \bar{\nu}_\tau)$ and their charged conjugates.

³ ALBRECHT 93G use tau pair events of the type $\tau^- \tau^+ \rightarrow (\mu^- \bar{\nu}_\mu \nu_\tau)(e^+ \nu_e \bar{\nu}_\tau)$ and their charged conjugates.

$\rho(\mu)$ PARAMETER(V-A) theory predicts $\rho = 0.75$.

| VALUE | EVTS | DOCUMENT ID | TECN | COMMENT |
|---|---------------|----------------|--|--|
| 0.763±0.020 OUR FIT | | | | |
| 0.770±0.022 OUR AVERAGE | | | | |
| 0.776±0.045±0.019 | 46k | HEISTER 01E | ALEP | 1991–1995 LEP runs |
| 0.999±0.098±0.045 | 22k | ABREU 00L | DLPH | 1992–1995 runs |
| 0.777±0.044±0.016 | 27k | ACKERSTAFF 99D | OPAL | 1990–1995 LEP runs |
| 0.69 ±0.06 ±0.06 | 1 ALBRECHT 98 | ARG | $E_{cm}^{ee} = 9.5\text{--}10.6 \text{ GeV}$ | |
| 0.54 ±0.28 ±0.14 | ABE 970 | SLD | 1993–1995 SLC runs | |
| 0.750±0.017±0.045 | 22k | ALEXANDER 97F | CLEO | $E_{cm}^{ee} = 10.6 \text{ GeV}$ |
| 0.76 ±0.07 ±0.08 | 3230 | ALBRECHT 93G | ARG | $E_{cm}^{ee} = 9.4\text{--}10.6 \text{ GeV}$ |
| 0.734±0.055±0.027 | 3041 | ALBRECHT 90E | ARG | $E_{cm}^{ee} = 9.4\text{--}10.6 \text{ GeV}$ |
| 0.89 ±0.14 ±0.08 | 1909 | FORD 87B | MAC | $E_{cm}^{ee} = 29 \text{ GeV}$ |
| 0.81 ±0.13 | 727 | BEHRENDS 85 | CLEO | e^+e^- near $\gamma(4S)$ |
| • • • We do not use the following data for averages, fits, limits, etc. • • • | | | | |
| 0.747±0.048±0.044 | 13k | AMMAR 97B | CLEO | Repl. by ALEXANDER 97F |
| 0.693±0.057±0.028 | | BUSKULIC 95D | ALEP | Repl. by HEISTER 01E |

¹ ALBRECHT 98 use tau pair events of the type $\tau^-\tau^+ \rightarrow (\ell^-\bar{\nu}_\ell\nu_\tau)(\pi^+\pi^0\bar{\nu}_\tau)$, and their charged conjugates.

 $\xi(e$ or $\mu)$ PARAMETER(V-A) theory predicts $\xi = 1$.

| VALUE | EVTS | DOCUMENT ID | TECN | COMMENT |
|---|---------------------------|---------------------------|--|--|
| 0.985±0.030 OUR FIT | | | | |
| 0.981±0.031 OUR AVERAGE | | | | |
| 0.986±0.068±0.031 | 81k | HEISTER 01E | ALEP | 1991–1995 LEP runs |
| 0.929±0.070±0.030 | 36k | ABREU 00L | DLPH | 1992–1995 runs |
| 0.98 ±0.22 ±0.10 | 46k | ACKERSTAFF 99D | OPAL | 1990–1995 LEP runs |
| 0.70 ±0.16 | 54k | ACCIARRI 98R | L3 | 1991–1995 LEP runs |
| 1.03 ±0.11 | 1 ALBRECHT 98 | ARG | $E_{cm}^{ee} = 9.5\text{--}10.6 \text{ GeV}$ | |
| 1.05 ±0.35 ±0.04 | ² ABE 970 | SLD | 1993–1995 SLC runs | |
| 1.007±0.040±0.015 | 55k | ALEXANDER 97F | CLEO | $E_{cm}^{ee} = 10.6 \text{ GeV}$ |
| • • • We do not use the following data for averages, fits, limits, etc. • • • | | | | |
| 0.94 ±0.21 ±0.07 | 18k | ACCIARRI 96H | L3 | Repl. by ACCIARRI 98R |
| 0.97 ±0.14 | ³ ALBRECHT 95C | ARG | Repl. by ALBRECHT 98 | |
| 1.18 ±0.15 ±0.16 | BUSKULIC 95D | ALEP | Repl. by HEISTER 01E | |
| 0.90 ±0.15 ±0.10 | 3230 | ⁴ ALBRECHT 93G | ARG | $E_{cm}^{ee} = 9.4\text{--}10.6 \text{ GeV}$ |

¹ Combined fit to ARGUS tau decay parameter measurements in ALBRECHT 98, ALBRECHT 95C, ALBRECHT 93G, and ALBRECHT 94E. ALBRECHT 98 use tau pair events of the type $\tau^-\tau^+ \rightarrow (\ell^-\bar{\nu}_\ell\nu_\tau)(\pi^+\pi^0\bar{\nu}_\tau)$, and their charged conjugates.

² ABE 970 assume $\eta = 0$ in their fit. Letting η vary in the fit gives a ξ value of $1.02 \pm 0.36 \pm 0.05$.

³ Combined fit to ARGUS tau decay parameter measurements in ALBRECHT 95C, ALBRECHT 93G, and ALBRECHT 94E. ALBRECHT 95C uses events of the type $\tau^-\tau^+ \rightarrow (\ell^-\bar{\nu}_\ell\nu_\tau)(h^+h^-h^+\bar{\nu}_\tau)$ and their charged conjugates.

⁴ ALBRECHT 93G measurement determines $|\xi|$ for the case $\xi(e) = \xi(\mu)$, but the authors point out that other LEP experiments determine the sign to be positive.

$\xi(e)$ PARAMETER $(V-A)$ theory predicts $\xi = 1$.

| VALUE | EVTS | DOCUMENT ID | TECN | COMMENT |
|---|------|--------------------------|------|--|
| 0.994±0.040 OUR FIT | | | | |
| 1.00 ±0.04 OUR AVERAGE | | | | |
| 1.011±0.094±0.038 | 44k | HEISTER 01E | ALEP | 1991–1995 LEP runs |
| 1.01 ±0.12 ±0.05 | 17k | ABREU 00L | DLPH | 1992–1995 runs |
| 1.13 ±0.39 ±0.14 | 25k | ACKERSTAFF 99D | OPAL | 1990–1995 LEP runs |
| 1.11 ±0.20 ±0.08 | | ¹ ALBRECHT 98 | ARG | $E_{cm}^{ee} = 9.5\text{--}10.6 \text{ GeV}$ |
| 1.16 ±0.52 ±0.06 | | ABE 970 | SLD | 1993–1995 SLC runs |
| 0.979±0.048±0.016 | 34k | ALEXANDER 97F | CLEO | $E_{cm}^{ee} = 10.6 \text{ GeV}$ |
| • • • We do not use the following data for averages, fits, limits, etc. • • • | | | | |
| 1.03 ±0.23 ±0.09 | | BUSKULIC 95D | ALEP | Repl. by HEISTER 01E |

¹ ALBRECHT 98 use tau pair events of the type $\tau^- \tau^+ \rightarrow (\ell^- \bar{\nu}_\ell \nu_\tau)(\pi^+ \pi^0 \bar{\nu}_\tau)$, and their charged conjugates.

 $\xi(\mu)$ PARAMETER $(V-A)$ theory predicts $\xi = 1$.

| VALUE | EVTS | DOCUMENT ID | TECN | COMMENT |
|---|------|--------------------------|------|--|
| 1.030±0.059 OUR FIT | | | | |
| 1.06 ±0.06 OUR AVERAGE | | | | |
| 1.030±0.120±0.050 | 46k | HEISTER 01E | ALEP | 1991–1995 LEP runs |
| 1.16 ±0.19 ±0.06 | 22k | ABREU 00L | DLPH | 1992–1995 runs |
| 0.79 ±0.41 ±0.09 | 27k | ACKERSTAFF 99D | OPAL | 1990–1995 LEP runs |
| 1.26 ±0.27 ±0.14 | | ¹ ALBRECHT 98 | ARG | $E_{cm}^{ee} = 9.5\text{--}10.6 \text{ GeV}$ |
| 0.75 ±0.50 ±0.14 | | ABE 970 | SLD | 1993–1995 SLC runs |
| 1.054±0.069±0.047 | 22k | ALEXANDER 97F | CLEO | $E_{cm}^{ee} = 10.6 \text{ GeV}$ |
| • • • We do not use the following data for averages, fits, limits, etc. • • • | | | | |
| 1.23 ±0.22 ±0.10 | | BUSKULIC 95D | ALEP | Repl. by HEISTER 01E |

¹ ALBRECHT 98 use tau pair events of the type $\tau^- \tau^+ \rightarrow (\ell^- \bar{\nu}_\ell \nu_\tau)(\pi^+ \pi^0 \bar{\nu}_\tau)$, and their charged conjugates.

 $\eta(e \text{ or } \mu)$ PARAMETER $(V-A)$ theory predicts $\eta = 0$.

| VALUE | EVTS | DOCUMENT ID | TECN | COMMENT |
|---|------|----------------|------|--|
| 0.013±0.020 OUR FIT | | | | |
| 0.015±0.021 OUR AVERAGE | | | | |
| 0.012±0.026±0.004 | 81k | HEISTER 01E | ALEP | 1991–1995 LEP runs |
| -0.005±0.036±0.037 | | ABREU 00L | DLPH | 1992–1995 runs |
| 0.027±0.055±0.005 | 46k | ACKERSTAFF 99D | OPAL | 1990–1995 LEP runs |
| 0.27 ±0.14 | 54k | ACCIARRI 98R | L3 | 1991–1995 LEP runs |
| -0.13 ±0.47 ±0.15 | | ABE 970 | SLD | 1993–1995 SLC runs |
| -0.015±0.061±0.062 | 31k | AMMAR 97B | CLEO | $E_{cm}^{ee} = 10.6 \text{ GeV}$ |
| 0.03 ±0.18 ±0.12 | 8.2k | ALBRECHT 95 | ARG | $E_{cm}^{ee} = 9.5\text{--}10.6 \text{ GeV}$ |
| • • • We do not use the following data for averages, fits, limits, etc. • • • | | | | |
| 0.25 ±0.17 ±0.11 | 18k | ACCIARRI 96H | L3 | Repl. by ACCIARRI 98R |
| -0.04 ±0.15 ±0.11 | | BUSKULIC 95D | ALEP | Repl. by HEISTER 01E |

$\eta(\mu)$ PARAMETER $(V-A)$ theory predicts $\eta = 0$.

| VALUE | EVTS | DOCUMENT ID | TECN | COMMENT |
|---|------|-------------------------|------|-------------------------------|
| 0.094±0.073 OUR FIT | | | | |
| 0.17 ± 0.15 OUR AVERAGE | | | | |
| 0.160±0.150±0.060 | 46k | HEISTER | 01E | ALEP 1991–1995 LEP runs |
| 0.72 ± 0.32 ± 0.15 | | ABREU | 00L | DLPH 1992–1995 runs |
| −0.59 ± 0.82 ± 0.45 | | ¹ ABE | 970 | SLD 1993–1995 SLC runs |
| 0.010±0.149±0.171 | 13k | ² AMMAR | 97B | CLEO $E_{cm}^{ee} = 10.6$ GeV |
| • • • We do not use the following data for averages, fits, limits, etc. • • • | | | | |
| 0.010±0.065±0.001 | 27k | ³ ACKERSTAFF | 99D | OPAL 1990–1995 LEP runs |
| −0.24 ± 0.23 ± 0.18 | | BUSKULIC | 95D | ALEP Repl. by HEISTER 01E |

¹ Highly correlated (corr. = 0.92) with ABE 970 $\rho(\mu)$ measurement.² Highly correlated (corr. = 0.949) with AMMAR 97B $\rho(\mu)$ value.³ ACKERSTAFF 99D result is dominated by a constraint on η from the OPAL measurements of the τ lifetime and $B(\tau^- \rightarrow \mu^- \bar{\nu}_\mu \nu_\tau)$ assuming lepton universality for the total coupling strength. **$(\delta\xi)(e \text{ or } \mu)$ PARAMETER** $(V-A)$ theory predicts $(\delta\xi) = 0.75$.

| VALUE | EVTS | DOCUMENT ID | TECN | COMMENT |
|---|------|-----------------------|------|--|
| 0.746±0.021 OUR FIT | | | | |
| 0.744±0.022 OUR AVERAGE | | | | |
| 0.776±0.045±0.024 | 81k | HEISTER | 01E | ALEP 1991–1995 LEP runs |
| 0.779±0.070±0.028 | 36k | ABREU | 00L | DLPH 1992–1995 runs |
| 0.65 ± 0.14 ± 0.07 | 46k | ACKERSTAFF | 99D | OPAL 1990–1995 LEP runs |
| 0.70 ± 0.11 | 54k | ACCIARRI | 98R | L3 1991–1995 LEP runs |
| 0.63 ± 0.09 | | ¹ ALBRECHT | 98 | ARG $E_{cm}^{ee} = 9.5\text{--}10.6$ GeV |
| 0.88 ± 0.27 ± 0.04 | | ² ABE | 970 | SLD 1993–1995 SLC runs |
| 0.745±0.026±0.009 | 55k | ALEXANDER | 97F | CLEO $E_{cm}^{ee} = 10.6$ GeV |
| • • • We do not use the following data for averages, fits, limits, etc. • • • | | | | |
| 0.81 ± 0.14 ± 0.06 | 18k | ACCIARRI | 96H | L3 Repl. by ACCIARRI 98R |
| 0.65 ± 0.12 | | ³ ALBRECHT | 95C | ARG Repl. by ALBRECHT 98 |
| 0.88 ± 0.11 ± 0.07 | | BUSKULIC | 95D | ALEP Repl. by HEISTER 01E |

¹ Combined fit to ARGUS tau decay parameter measurements in ALBRECHT 98, ALBRECHT 95C, ALBRECHT 93G, and ALBRECHT 94E. ALBRECHT 98 use tau pair events of the type $\tau^- \tau^+ \rightarrow (\ell^- \bar{\nu}_\ell \nu_\tau)(\pi^+ \pi^0 \bar{\nu}_\tau)$, and their charged conjugates.² ABE 970 assume $\eta = 0$ in their fit. Letting η vary in the fit gives a $(\delta\xi)$ value of $0.87 \pm 0.27 \pm 0.04$.³ Combined fit to ARGUS tau decay parameter measurements in ALBRECHT 95C, ALBRECHT 93G, and ALBRECHT 94E. ALBRECHT 95C uses events of the type $\tau^- \tau^+ \rightarrow (\ell^- \bar{\nu}_\ell \nu_\tau)(h^+ h^- h^+ \bar{\nu}_\tau)$ and their charged conjugates. **$(\delta\xi)(e)$ PARAMETER** $(V-A)$ theory predicts $(\delta\xi) = 0.75$.

| VALUE | EVTS | DOCUMENT ID | TECN | COMMENT |
|--------------------------------|------|-------------|------|-------------------------|
| 0.734±0.028 OUR FIT | | | | |
| 0.731±0.029 OUR AVERAGE | | | | |
| 0.778±0.066±0.024 | 44k | HEISTER | 01E | ALEP 1991–1995 LEP runs |
| 0.85 ± 0.12 ± 0.04 | 17k | ABREU | 00L | DLPH 1992–1995 runs |
| 0.72 ± 0.31 ± 0.14 | 25k | ACKERSTAFF | 99D | OPAL 1990–1995 LEP runs |

| | |
|---|--|
| 0.56 $\pm 0.14 \pm 0.06$ | 1 ALBRECHT 98 ARG $E_{\text{cm}}^{\text{ee}} = 9.5\text{--}10.6 \text{ GeV}$ |
| 0.85 $\pm 0.43 \pm 0.08$ | ABE 970 SLD 1993–1995 SLC runs |
| 0.720 $\pm 0.032 \pm 0.010$ | 34k ALEXANDER 97F CLEO $E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$ |
| $\bullet \bullet \bullet$ We do not use the following data for averages, fits, limits, etc. $\bullet \bullet \bullet$ | |
| 1.11 $\pm 0.17 \pm 0.07$ | BUSKULIC 95D ALEP Repl. by HEISTER 01E |

¹ ALBRECHT 98 use tau pair events of the type $\tau^- \tau^+ \rightarrow (\ell^- \bar{\nu}_\ell \nu_\tau)(\pi^+ \pi^0 \bar{\nu}_\tau)$, and their charged conjugates.

$(\delta\xi)(\mu)$ PARAMETER

$(V-A)$ theory predicts $(\delta\xi) = 0.75$.

| VALUE | EVTS | DOCUMENT ID | TECN | COMMENT |
|---|---------------|----------------|--|--|
| 0.778 ± 0.037 OUR FIT | | | | |
| 0.79 ± 0.04 OUR AVERAGE | | | | |
| 0.786 $\pm 0.066 \pm 0.028$ | 46k | HEISTER 01E | ALEP | 1991–1995 LEP runs |
| 0.86 $\pm 0.13 \pm 0.04$ | 22k | ABREU 00L | DLPH | 1992–1995 runs |
| 0.63 $\pm 0.23 \pm 0.05$ | 27k | ACKERSTAFF 99D | OPAL | 1990–1995 LEP runs |
| 0.73 $\pm 0.18 \pm 0.10$ | 1 ALBRECHT 98 | ARG | $E_{\text{cm}}^{\text{ee}} = 9.5\text{--}10.6 \text{ GeV}$ | |
| 0.82 $\pm 0.32 \pm 0.07$ | ABE 970 | SLD | 1993–1995 SLC runs | |
| 0.786 $\pm 0.041 \pm 0.032$ | 22k | ALEXANDER 97F | CLEO | $E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$ |
| $\bullet \bullet \bullet$ We do not use the following data for averages, fits, limits, etc. $\bullet \bullet \bullet$ | | | | |
| 0.71 $\pm 0.14 \pm 0.06$ | BUSKULIC 95D | ALEP | Repl. by HEISTER 01E | |

¹ ALBRECHT 98 use tau pair events of the type $\tau^- \tau^+ \rightarrow (\ell^- \bar{\nu}_\ell \nu_\tau)(\pi^+ \pi^0 \bar{\nu}_\tau)$, and their charged conjugates.

$\xi(\pi)$ PARAMETER

$(V-A)$ theory predicts $\xi(\pi) = 1$.

| VALUE | EVTS | DOCUMENT ID | TECN | COMMENT |
|---|---------------------------|-------------|----------------------|--|
| 0.993 ± 0.022 OUR FIT | | | | |
| 0.994 ± 0.023 OUR AVERAGE | | | | |
| 0.994 $\pm 0.020 \pm 0.014$ | 27k | HEISTER 01E | ALEP | 1991–1995 LEP runs |
| 0.81 $\pm 0.17 \pm 0.02$ | ABE 970 | SLD | 1993–1995 SLC runs | |
| 1.03 $\pm 0.06 \pm 0.04$ | 2.0k | COAN 97 | CLEO | $E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$ |
| $\bullet \bullet \bullet$ We do not use the following data for averages, fits, limits, etc. $\bullet \bullet \bullet$ | | | | |
| 0.987 $\pm 0.057 \pm 0.027$ | BUSKULIC 95D | ALEP | Repl. by HEISTER 01E | |
| 0.95 $\pm 0.11 \pm 0.05$ | ¹ BUSKULIC 94D | ALEP | 1990+1991 LEP run | |

¹ Superseded by BUSKULIC 95D.

$\xi(\rho)$ PARAMETER

$(V-A)$ theory predicts $\xi(\rho) = 1$.

| VALUE | EVTS | DOCUMENT ID | TECN | COMMENT |
|---|---------------------------|---------------------------|----------------------|--|
| 0.994 ± 0.008 OUR FIT | | | | |
| 0.994 ± 0.009 OUR AVERAGE | | | | |
| 0.987 $\pm 0.012 \pm 0.011$ | 59k | HEISTER 01E | ALEP | 1991–1995 LEP runs |
| 0.99 $\pm 0.12 \pm 0.04$ | ABE 970 | SLD | 1993–1995 SLC runs | |
| 0.995 $\pm 0.010 \pm 0.003$ | 66k | ALEXANDER 97F | CLEO | $E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$ |
| 1.022 $\pm 0.028 \pm 0.030$ | 1.7k | ¹ ALBRECHT 94E | ARG | $E_{\text{cm}}^{\text{ee}} = 9.4\text{--}10.6 \text{ GeV}$ |
| $\bullet \bullet \bullet$ We do not use the following data for averages, fits, limits, etc. $\bullet \bullet \bullet$ | | | | |
| 1.045 $\pm 0.058 \pm 0.032$ | BUSKULIC 95D | ALEP | Repl. by HEISTER 01E | |
| 1.03 $\pm 0.11 \pm 0.05$ | ² BUSKULIC 94D | ALEP | 1990+1991 LEP run | |

¹ ALBRECHT 94E measure the square of this quantity and use the sign determined by ALBRECHT 90I to obtain the quoted result.

² Superseded by BUSKULIC 95D.

$\xi(a_1)$ PARAMETER $(V-A)$ theory predicts $\xi(a_1) = 1$.

| VALUE | EVTS | DOCUMENT ID | TECN | COMMENT |
|---|-------|-----------------------------|------|--------------------------------------|
| 1.001±0.027 OUR FIT | | | | |
| 1.002±0.028 OUR AVERAGE | | | | |
| 1.000 ± 0.016 ± 0.024 | 35k | ¹ HEISTER 01E | ALEP | 1991–1995 LEP runs |
| 1.02 ± 0.13 ± 0.03 | 17.2k | ASNER 00 | CLEO | $E_{cm}^{ee} = 10.6$ GeV |
| 1.29 ± 0.26 ± 0.11 | 7.4k | ² ACKERSTAFF 97R | OPAL | 1992–1994 LEP runs |
| 0.85 $^{+0.15}_{-0.17}$ ± 0.05 | | ALBRECHT 95C | ARG | $E_{cm}^{ee} = 9.5\text{--}10.6$ GeV |
| 1.25 ± 0.23 $^{+0.15}_{-0.08}$ | 7.5k | ALBRECHT 93C | ARG | $E_{cm}^{ee} = 9.4\text{--}10.6$ GeV |
| • • • We do not use the following data for averages, fits, limits, etc. • • • | | | | |
| 1.08 $^{+0.46}_{-0.41}$ $^{+0.14}_{-0.25}$ | 2.6k | ³ AKERS 95P | OPAL | Repl. by ACKER-STAFF 97R |
| 0.937 ± 0.116 ± 0.064 | | BUSKULIC 95D | ALEP | Repl. by HEISTER 01E |

¹ HEISTER 01E quote $1.000 \pm 0.016 \pm 0.013 \pm 0.020$ where the errors are statistical, systematic, and an uncertainty due to the final state model. We combine the systematic error and model uncertainty.

² ACKERSTAFF 97R obtain this result with a model independent fit to the hadronic structure functions. Fitting with the model of Kuhn and Santamaria (ZPHY **C48**, 445 (1990)) gives $0.87 \pm 0.16 \pm 0.04$, and with the model of Isgur *et al.* (PR **D39**, 1357 (1989)) they obtain $1.20 \pm 0.21 \pm 0.14$.

³ AKERS 95P obtain this result with a model independent fit to the hadronic structure functions. Fitting with the model of Kuhn and Santamaria (ZPHY **C48**, 445 (1990)) gives $0.87 \pm 0.27 $^{+0.05}_{-0.06}$$, and with the model of Isgur *et al.* (PR **D39**, 1357 (1989)) they obtain $1.10 \pm 0.31 $^{+0.13}_{-0.14}$$.

 $\xi(\text{all hadronic modes})$ PARAMETER $(V-A)$ theory predicts $\xi = 1$.

| VALUE | EVTS | DOCUMENT ID | TECN | COMMENT |
|---|-------|-----------------------------|------|--------------------------------------|
| 0.995±0.007 OUR FIT | | | | |
| 0.997±0.007 OUR AVERAGE | | | | |
| 0.992 ± 0.007 ± 0.008 | 102k | ¹ HEISTER 01E | ALEP | 1991–1995 LEP runs |
| 0.997 ± 0.027 ± 0.011 | 39k | ² ABREU 00L | DLPH | 1992–1995 runs |
| 1.02 ± 0.13 ± 0.03 | 17.2k | ³ ASNER 00 | CLEO | $E_{cm}^{ee} = 10.6$ GeV |
| 1.032 ± 0.031 | 37k | ⁴ ACCIARRI 98R | L3 | 1991–1995 LEP runs |
| 0.93 ± 0.10 ± 0.04 | | ABE 970 | SLD | 1993–1995 SLC runs |
| 1.29 ± 0.26 ± 0.11 | 7.4k | ⁵ ACKERSTAFF 97R | OPAL | 1992–1994 LEP runs |
| 0.995 ± 0.010 ± 0.003 | 66k | ⁶ ALEXANDER 97F | CLEO | $E_{cm}^{ee} = 10.6$ GeV |
| 1.03 ± 0.06 ± 0.04 | 2.0k | ⁷ COAN 97 | CLEO | $E_{cm}^{ee} = 10.6$ GeV |
| 1.017 ± 0.039 | | ⁸ ALBRECHT 95C | ARG | $E_{cm}^{ee} = 9.5\text{--}10.6$ GeV |
| 1.25 ± 0.23 $^{+0.15}_{-0.08}$ | 7.5k | ⁹ ALBRECHT 93C | ARG | $E_{cm}^{ee} = 9.4\text{--}10.6$ GeV |
| • • • We do not use the following data for averages, fits, limits, etc. • • • | | | | |
| 0.970 ± 0.053 ± 0.011 | 14k | ¹⁰ ACCIARRI 96H | L3 | Repl. by ACCIARRI 98R |
| 1.08 $^{+0.46}_{-0.41}$ $^{+0.14}_{-0.25}$ | 2.6k | ¹¹ AKERS 95P | OPAL | Repl. by ACKER-STAFF 97R |
| 1.006 ± 0.032 ± 0.019 | | ¹² BUSKULIC 95D | ALEP | Repl. by HEISTER 01E |
| 1.022 ± 0.028 ± 0.030 | 1.7k | ¹³ ALBRECHT 94E | ARG | $E_{cm}^{ee} = 9.4\text{--}10.6$ GeV |
| 0.99 ± 0.07 ± 0.04 | | ¹⁴ BUSKULIC 94D | ALEP | 1990+1991 LEP run |

- ¹ HEISTER 01E quote $0.992 \pm 0.007 \pm 0.006 \pm 0.005$ where the errors are statistical, systematic, and an uncertainty due to the final state model. We combine the systematic error and model uncertainty. They use $\tau \rightarrow \pi \nu_\tau$, $\tau \rightarrow K \nu_\tau$, $\tau \rightarrow \rho \nu_\tau$, and $\tau \rightarrow a_1 \nu_\tau$ decays.
- ² ABREU 00L use $\tau^- \rightarrow h^- \geq 0 \pi^0 \nu_\tau$ decays.
- ³ ASNER 00 use $\tau^- \rightarrow \pi^- 2\pi^0 \nu_\tau$ decays.
- ⁴ ACCIARRI 98R use $\tau \rightarrow \pi \nu_\tau$, $\tau \rightarrow K \nu_\tau$, and $\tau \rightarrow \rho \nu_\tau$ decays.
- ⁵ ACKERSTAFF 97R use $\tau \rightarrow a_1 \nu_\tau$ decays.
- ⁶ ALEXANDER 97F use $\tau \rightarrow \rho \nu_\tau$ decays.
- ⁷ COAN 97 use $h^+ h^-$ energy correlations.
- ⁸ Combined fit to ARGUS tau decay parameter measurements in ALBRECHT 95C, ALBRECHT 93G, and ALBRECHT 94E.
- ⁹ Uses $\tau \rightarrow a_1 \nu_\tau$ decays. Replaced by ALBRECHT 95C.
- ¹⁰ ACCIARRI 96H use $\tau \rightarrow \pi \nu_\tau$, $\tau \rightarrow K \nu_\tau$, and $\tau \rightarrow \rho \nu_\tau$ decays.
- ¹¹ AKERS 95P use $\tau \rightarrow a_1 \nu_\tau$ decays.
- ¹² BUSKULIC 95D use $\tau \rightarrow \pi \nu_\tau$, $\tau \rightarrow \rho \nu_\tau$, and $\tau \rightarrow a_1 \nu_\tau$ decays.
- ¹³ ALBRECHT 94E measure the square of this quantity and use the sign determined by ALBRECHT 90I to obtain the quoted result. Uses $\tau \rightarrow a_1 \nu_\tau$ decays. Replaced by ALBRECHT 95C.
- ¹⁴ BUSKULIC 94D use $\tau \rightarrow \pi \nu_\tau$ and $\tau \rightarrow \rho \nu_\tau$ decays. Superseded by BUSKULIC 95D.

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| AUBERT | 10B | PRL 104 021802 | B. Aubert <i>et al.</i> | (BABAR Collab.) |
| AUBERT | 10F | PRL 105 051602 | B. Aubert <i>et al.</i> | (BABAR Collab.) |
| HAYASAKA | 10 | PL B687 139 | K. Hayasaka <i>et al.</i> | (BELLE Collab.) |
| LEE | 10 | PR D81 113007 | M.J. Lee <i>et al.</i> | (BELLE Collab.) |
| LEES | 10A | PR D81 111101 | J.P. Lees <i>et al.</i> | (BABAR Collab.) |
| MIYAZAKI | 10 | PL B682 355 | Y. Miyazaki <i>et al.</i> | (BELLE Collab.) |
| MIYAZAKI | 10A | PL B692 4 | Y. Miyazaki <i>et al.</i> | (BELLE Collab.) |
| AUBERT | 09AK | PR D80 092005 | B. Aubert <i>et al.</i> | (BABAR Collab.) |
| AUBERT | 09D | PR D79 012004 | B. Aubert <i>et al.</i> | (BABAR Collab.) |
| AUBERT | 09W | PRL 103 021801 | B. Aubert <i>et al.</i> | (BABAR Collab.) |
| GROZIN | 09A | PAN 72 1203 | A.G. Grozin, I.B. Khriplovich, A.S. Rudenko | (NOVO) |
| INAMI | 09 | PL B672 209 | K. Inami <i>et al.</i> | (BELLE Collab.) |
| MIYAZAKI | 09 | PL B672 317 | Y. Miyazaki <i>et al.</i> | (BELLE Collab.) |
| AUBERT | 08 | PRL 100 011801 | B. Aubert <i>et al.</i> | (BABAR Collab.) |
| AUBERT | 08AE | PR D77 112002 | B. Aubert <i>et al.</i> | (BABAR Collab.) |
| AUBERT | 08K | PRL 100 071802 | B. Aubert <i>et al.</i> | (BABAR Collab.) |
| FUJIKAWA | 08 | PR D78 072006 | M. Fujikawa <i>et al.</i> | (BELLE Collab.) |
| HAYASAKA | 08 | PL B666 16 | K. Hayasaka <i>et al.</i> | (BELLE Collab.) |
| MIYAZAKI | 08 | PL B660 154 | Y. Miyazaki <i>et al.</i> | (BELLE Collab.) |
| NISHIO | 08 | PL B664 35 | Y. Nishio <i>et al.</i> | (BELLE Collab.) |
| ANASHIN | 07 | JETPL 85 347 | V.V. Anashin <i>et al.</i> | (KEDR Collab.) |

Translated from ZETFP 85 429.

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| AUBERT | 07AP | PR D76 051104 | B. Aubert <i>et al.</i> | (BABAR Collab.) |
| AUBERT | 07BK | PRL 99 251803 | B. Aubert <i>et al.</i> | (BABAR Collab.) |
| AUBERT | 07I | PRL 98 061803 | B. Aubert <i>et al.</i> | (BABAR Collab.) |
| BELLOUS | 07 | PRL 99 011801 | K. Belous <i>et al.</i> | (BELLE Collab.) |
| EIDELMAN | 07 | MPL A22 159 | S. Eidelman, M. Passera | (NOVO, PADO) |
| EPIFANOV | 07 | PL B654 65 | D. Epifanov <i>et al.</i> | (BELLE Collab.) |
| MIYAZAKI | 07 | PL B648 341 | Y. Miyazaki <i>et al.</i> | (BELLE Collab.) |
| ABDALLAH | 06A | EPJ C46 1 | J. Abdallah <i>et al.</i> | (DELPHI Collab.) |
| AUBERT | 06C | PRL 96 041801 | B. Aubert <i>et al.</i> | (BABAR Collab.) |
| AUBERT,B | 06 | PR D73 112003 | B. Aubert <i>et al.</i> | (BABAR Collab.) |
| INAMI | 06 | PL B643 5 | K. Inami <i>et al.</i> | (BELLE Collab.) |
| MIYAZAKI | 06 | PL B632 51 | Y. Miyazaki <i>et al.</i> | (BELLE Collab.) |
| MIYAZAKI | 06A | PL B639 159 | Y. Miyazaki <i>et al.</i> | (BELLE Collab.) |
| PDG | 06 | JP G33 1 | W.-M. Yao <i>et al.</i> | (PDG Collab.) |
| YUSA | 06 | PL B640 138 | Y. Yusa <i>et al.</i> | (BELLE Collab.) |
| ARMS | 05 | PRL 94 241802 | K. Arms <i>et al.</i> | (CLEO Collab.) |
| AUBERT,B | 05A | PRL 95 041802 | B. Aubert <i>et al.</i> | (BABAR Collab.) |
| AUBERT,B | 05F | PR D72 012003 | B. Aubert <i>et al.</i> | (BABAR Collab.) |
| AUBERT,B | 05W | PR D72 072001 | B. Aubert <i>et al.</i> | (BABAR Collab.) |
| AUBERT,BE | 05D | PRL 95 191801 | B. Aubert <i>et al.</i> | (BABAR Collab.) |
| ENARI | 05 | PL B622 218 | Y. Enari <i>et al.</i> | (BELLE Collab.) |
| HAYASAKA | 05 | PL B613 20 | K. Hayasaka <i>et al.</i> | (BELLE Collab.) |
| SCHAEL | 05C | PRPL 421 191 | S. Schael <i>et al.</i> | (ALEPH Collab.) |
| ABBIENDI | 04J | EPJ C35 437 | G. Abbiendi <i>et al.</i> | (OPAL Collab.) |
| ABDALLAH | 04K | EPJ C35 159 | J. Abdallah <i>et al.</i> | (DELPHI Collab.) |
| ABDALLAH | 04T | EPJ C36 283 | J. Abdallah <i>et al.</i> | (DELPHI Collab.) |
| ABE | 04B | PRL 92 171802 | K. Abe <i>et al.</i> | (BELLE Collab.) |
| ACHARD | 04G | PL B585 53 | P. Achard <i>et al.</i> | (L3 Collab.) |
| AUBERT | 04J | PRL 92 121801 | B. Aubert <i>et al.</i> | (BABAR Collab.) |
| ENARI | 04 | PRL 93 081803 | Y. Enari <i>et al.</i> | (BELLE Collab.) |
| PDG | 04 | PL B592 1 | S. Eidelman <i>et al.</i> | (PDG Collab.) |
| YUSA | 04 | PL B589 103 | Y. Yusa <i>et al.</i> | (BELLE Collab.) |
| ABBIENDI | 03 | PL B551 35 | G. Abbiendi <i>et al.</i> | (OPAL Collab.) |
| BRIERE | 03 | PRL 90 181802 | R. A. Briere <i>et al.</i> | (CLEO Collab.) |
| HEISTER | 03F | EPJ C30 291 | A. Heister <i>et al.</i> | (ALEPH Collab.) |
| INAMI | 03 | PL B551 16 | K. Inami <i>et al.</i> | (BELLE Collab.) |
| CHEH | 02C | PR D66 071101 | S. Chen <i>et al.</i> | (CLEO Collab.) |
| REGAN | 02 | PRL 88 071805 | B.C. Regan <i>et al.</i> | (OPAL Collab.) |
| ABBIENDI | 01J | EPJ C19 653 | G. Abbiendi <i>et al.</i> | (DELPHI Collab.) |
| ABREU | 01M | EPJ C20 617 | P. Abreu <i>et al.</i> | (L3 Collab.) |
| ACCIARRI | 01F | PL B507 47 | M. Acciarri <i>et al.</i> | (L3 Collab.) |
| ACHARD | 01D | PL B519 189 | P. Achard <i>et al.</i> | (CLEO Collab.) |
| ANASTASSOV | 01 | PRL 86 4467 | A. Anastassov <i>et al.</i> | (OPAL Collab.) |
| HEISTER | 01E | EPJ C22 217 | A. Heister <i>et al.</i> | (ALEPH Collab.) |
| ABBIENDI | 00A | PL B492 23 | G. Abbiendi <i>et al.</i> | (OPAL Collab.) |
| ABBIENDI | 00C | EPJ C13 213 | G. Abbiendi <i>et al.</i> | (OPAL Collab.) |
| ABBIENDI | 00D | EPJ C13 197 | G. Abbiendi <i>et al.</i> | (OPAL Collab.) |
| ABREU | 00L | EPJ C16 229 | P. Abreu <i>et al.</i> | (DELPHI Collab.) |
| ACCIARRI | 00B | PL B479 67 | M. Acciarri <i>et al.</i> | (L3 Collab.) |
| AHMED | 00 | PR D61 071101 | S. Ahmed <i>et al.</i> | (CLEO Collab.) |
| ALBRECHT | 00 | PL B485 37 | H. Albrecht <i>et al.</i> | (ARGUS Collab.) |
| ASNER | 00 | PR D61 012002 | D.M. Asner <i>et al.</i> | (CLEO Collab.) |
| ASNER | 00B | PR D62 072006 | D.M. Asner <i>et al.</i> | (CLEO Collab.) |
| BERGFELD | 00 | PRL 84 830 | T. Bergfeld <i>et al.</i> | (CLEO Collab.) |
| BROWDER | 00 | PR D61 052004 | T.E. Browder <i>et al.</i> | (CLEO Collab.) |
| EDWARDS | 00A | PR D61 072003 | K.W. Edwards <i>et al.</i> | (CLEO Collab.) |
| GONZALEZ-S... | 00 | NP B582 3 | G.A. Gonzalez-Sprinberg <i>et al.</i> | (CLEO Collab.) |
| ABBIENDI | 99H | PL B447 134 | G. Abbiendi <i>et al.</i> | (OPAL Collab.) |
| ABREU | 99X | EPJ C10 201 | P. Abreu <i>et al.</i> | (DELPHI Collab.) |
| ACKERSTAFF | 99D | EPJ C8 3 | K. Ackerstaff <i>et al.</i> | (OPAL Collab.) |
| ACKERSTAFF | 99E | EPJ C8 183 | K. Ackerstaff <i>et al.</i> | (OPAL Collab.) |
| BARATE | 99K | EPJ C10 1 | R. Barate <i>et al.</i> | (ALEPH Collab.) |
| BARATE | 99R | EPJ C11 599 | R. Barate <i>et al.</i> | (ALEPH Collab.) |
| BISHAI | 99 | PRL 82 281 | M. Bishai <i>et al.</i> | (CLEO Collab.) |
| GODANG | 99 | PR D59 091303 | R. Godang <i>et al.</i> | (CLEO Collab.) |
| RICHICHI | 99 | PR D60 112002 | S.J. Richichi <i>et al.</i> | (CLEO Collab.) |
| ACCIARRI | 98C | PL B426 207 | M. Acciarri <i>et al.</i> | (L3 Collab.) |
| ACCIARRI | 98E | PL B434 169 | M. Acciarri <i>et al.</i> | (L3 Collab.) |
| ACCIARRI | 98R | PL B438 405 | M. Acciarri <i>et al.</i> | (L3 Collab.) |
| ACKERSTAFF | 98M | EPJ C4 193 | K. Ackerstaff <i>et al.</i> | (OPAL Collab.) |
| ACKERSTAFF | 98N | PL B431 188 | K. Ackerstaff <i>et al.</i> | (OPAL Collab.) |

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| ALBRECHT | 98 | PL B431 179 | H. Albrecht <i>et al.</i> | (ARGUS Collab.) |
| BARATE | 98 | EPJ C1 65 | R. Barate <i>et al.</i> | (ALEPH Collab.) |
| BARATE | 98E | EPJ C4 29 | R. Barate <i>et al.</i> | (ALEPH Collab.) |
| BLISS | 98 | PR D57 5903 | D.W. Bliss <i>et al.</i> | (CLEO Collab.) |
| ABE | 97O | PRL 78 4691 | K. Abe <i>et al.</i> | (SLD Collab.) |
| ACKERSTAFF | 97J | PL B404 213 | K. Ackerstaff <i>et al.</i> | (OPAL Collab.) |
| ACKERSTAFF | 97L | ZPHY C74 403 | K. Ackerstaff <i>et al.</i> | (OPAL Collab.) |
| ACKERSTAFF | 97R | ZPHY C75 593 | K. Ackerstaff <i>et al.</i> | (OPAL Collab.) |
| ALEXANDER | 97F | PR D56 5320 | J.P. Alexander <i>et al.</i> | (CLEO Collab.) |
| AMMAR | 97B | PRL 78 4686 | R. Ammar <i>et al.</i> | (CLEO Collab.) |
| ANASTASSOV | 97 | PR D55 2559 | A. Anastassov <i>et al.</i> | (CLEO Collab.) |
| Also | | PR D58 119903 (erratum) | A. Anastassov <i>et al.</i> | (CLEO Collab.) |
| ANDERSON | 97 | PRL 79 3814 | S. Anderson <i>et al.</i> | (CLEO Collab.) |
| AVERY | 97 | PR D55 R1119 | P. Avery <i>et al.</i> | (CLEO Collab.) |
| BARATE | 97I | ZPHY C74 387 | R. Barate <i>et al.</i> | (ALEPH Collab.) |
| BARATE | 97R | PL B414 362 | R. Barate <i>et al.</i> | (ALEPH Collab.) |
| BERGFELD | 97 | PRL 79 2406 | T. Bergfeld <i>et al.</i> | (CLEO Collab.) |
| BONVICINI | 97 | PRL 79 1221 | G. Bonvicini <i>et al.</i> | (CLEO Collab.) |
| BUSKULIC | 97C | ZPHY C74 263 | D. Buskulic <i>et al.</i> | (ALEPH Collab.) |
| COAN | 97 | PR D55 7291 | T.E. Coan <i>et al.</i> | (CLEO Collab.) |
| EDWARDS | 97 | PR D55 R3919 | K.W. Edwards <i>et al.</i> | (CLEO Collab.) |
| EDWARDS | 97B | PR D56 R5297 | K.W. Edwards <i>et al.</i> | (CLEO Collab.) |
| ESCRIBANO | 97 | PL B395 369 | R. Escribano, E. Masso | (BARC, PARIT) |
| ABREU | 96B | PL B365 448 | P. Abreu <i>et al.</i> | (DELPHI Collab.) |
| ACCIARRI | 96H | PL B377 313 | M. Acciarri <i>et al.</i> | (L3 Collab.) |
| ACCIARRI | 96K | PL B389 187 | M. Acciarri <i>et al.</i> | (L3 Collab.) |
| ALAM | 96 | PRL 76 2637 | M.S. Alam <i>et al.</i> | (CLEO Collab.) |
| ALBRECHT | 96E | PRPL 276 223 | H. Albrecht <i>et al.</i> | (ARGUS Collab.) |
| ALEXANDER | 96D | PL B369 163 | G. Alexander <i>et al.</i> | (OPAL Collab.) |
| ALEXANDER | 96E | PL B374 341 | G. Alexander <i>et al.</i> | (OPAL Collab.) |
| ALEXANDER | 96S | PL B388 437 | G. Alexander <i>et al.</i> | (OPAL Collab.) |
| BAI | 96 | PR D53 20 | J.Z. Bai <i>et al.</i> | (BES Collab.) |
| BALEST | 96 | PL B388 402 | R. Balest <i>et al.</i> | (CLEO Collab.) |
| BARTELT | 96 | PRL 76 4119 | J.E. Bartelt <i>et al.</i> | (CLEO Collab.) |
| BUSKULIC | 96 | ZPHY C70 579 | D. Buskulic <i>et al.</i> | (ALEPH Collab.) |
| BUSKULIC | 96C | ZPHY C70 561 | D. Buskulic <i>et al.</i> | (ALEPH Collab.) |
| COAN | 96 | PR D53 6037 | T.E. Coan <i>et al.</i> | (CLEO Collab.) |
| ABE | 95Y | PR D52 4828 | K. Abe <i>et al.</i> | (SLD Collab.) |
| ABREU | 95T | PL B357 715 | P. Abreu <i>et al.</i> | (DELPHI Collab.) |
| ABREU | 95U | PL B359 411 | P. Abreu <i>et al.</i> | (DELPHI Collab.) |
| ACCIARRI | 95 | PL B345 93 | M. Acciarri <i>et al.</i> | (L3 Collab.) |
| ACCIARRI | 95F | PL B352 487 | M. Acciarri <i>et al.</i> | (L3 Collab.) |
| AKERS | 95F | ZPHY C66 31 | R. Akers <i>et al.</i> | (OPAL Collab.) |
| AKERS | 95I | ZPHY C66 543 | R. Akers <i>et al.</i> | (OPAL Collab.) |
| AKERS | 95P | ZPHY C67 45 | R. Akers <i>et al.</i> | (OPAL Collab.) |
| AKERS | 95Y | ZPHY C68 555 | R. Akers <i>et al.</i> | (OPAL Collab.) |
| ALBRECHT | 95 | PL B341 441 | H. Albrecht <i>et al.</i> | (ARGUS Collab.) |
| ALBRECHT | 95C | PL B349 576 | H. Albrecht <i>et al.</i> | (ARGUS Collab.) |
| ALBRECHT | 95G | ZPHY C68 25 | H. Albrecht <i>et al.</i> | (ARGUS Collab.) |
| ALBRECHT | 95H | ZPHY C68 215 | H. Albrecht <i>et al.</i> | (ARGUS Collab.) |
| BALEST | 95C | PRL 75 3809 | R. Balest <i>et al.</i> | (CLEO Collab.) |
| BERNABEU | 95 | NP B436 474 | J. Bernabeu <i>et al.</i> | |
| BUSKULIC | 95C | PL B346 371 | D. Buskulic <i>et al.</i> | (ALEPH Collab.) |
| BUSKULIC | 95D | PL B346 379 | D. Buskulic <i>et al.</i> | (ALEPH Collab.) |
| Also | | PL B363 265 (erratum) | D. Buskulic <i>et al.</i> | (ALEPH Collab.) |
| ABREU | 94K | PL B334 435 | P. Abreu <i>et al.</i> | (DELPHI Collab.) |
| AKERS | 94E | PL B328 207 | R. Akers <i>et al.</i> | (OPAL Collab.) |
| AKERS | 94G | PL B339 278 | R. Akers <i>et al.</i> | (OPAL Collab.) |
| ALBRECHT | 94E | PL B337 383 | H. Albrecht <i>et al.</i> | (ARGUS Collab.) |
| ARTUSO | 94 | PRL 72 3762 | M. Artuso <i>et al.</i> | (CLEO Collab.) |
| BARTELT | 94 | PRL 73 1890 | J.E. Bartelt <i>et al.</i> | (CLEO Collab.) |
| BATTLE | 94 | PRL 73 1079 | M. Battle <i>et al.</i> | (CLEO Collab.) |
| BAUER | 94 | PR D50 13 | D.A. Bauer <i>et al.</i> | (TPC/2gamma Collab.) |
| BUSKULIC | 94D | PL B321 168 | D. Buskulic <i>et al.</i> | (ALEPH Collab.) |
| BUSKULIC | 94E | PL B332 209 | D. Buskulic <i>et al.</i> | (ALEPH Collab.) |
| BUSKULIC | 94F | PL B332 219 | D. Buskulic <i>et al.</i> | (ALEPH Collab.) |
| GIBAUT | 94B | PRL 73 934 | D. Gibaut <i>et al.</i> | (CLEO Collab.) |
| ADRIANI | 93M | PRPL 236 1 | O. Adriani <i>et al.</i> | (L3 Collab.) |
| ALBRECHT | 93C | ZPHY C58 61 | H. Albrecht <i>et al.</i> | (ARGUS Collab.) |
| ALBRECHT | 93G | PL B316 608 | H. Albrecht <i>et al.</i> | (ARGUS Collab.) |
| BALEST | 93 | PR D47 R3671 | R. Balest <i>et al.</i> | (CLEO Collab.) |

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| BEAN | 93 | PRL 70 138 | A. Bean <i>et al.</i> | (CLEO Collab.) |
| BORTOLETTO | 93 | PRL 71 1791 | D. Bortoletto <i>et al.</i> | (CLEO Collab.) |
| ESCRIBANO | 93 | PL B301 419 | R. Escribano, E. Masso | (BARC) |
| PROCARIO | 93 | PRL 70 1207 | M. Procario <i>et al.</i> | (CLEO Collab.) |
| ABREU | 92N | ZPHY C55 555 | P. Abreu <i>et al.</i> | (DELPHI Collab.) |
| ACTON | 92F | PL B281 405 | D.P. Acton <i>et al.</i> | (OPAL Collab.) |
| ACTON | 92H | PL B288 373 | P.D. Acton <i>et al.</i> | (OPAL Collab.) |
| AKERIB | 92 | PRL 69 3610 | D.S. Akerib <i>et al.</i> | (CLEO Collab.) |
| Also | | PRL 71 3395 (erratum) | D.S. Akerib <i>et al.</i> | (CLEO Collab.) |
| ALBRECHT | 92D | ZPHY C53 367 | H. Albrecht <i>et al.</i> | (ARGUS Collab.) |
| ALBRECHT | 92K | ZPHY C55 179 | H. Albrecht <i>et al.</i> | (ARGUS Collab.) |
| ALBRECHT | 92M | PL B292 221 | H. Albrecht <i>et al.</i> | (ARGUS Collab.) |
| ALBRECHT | 92Q | ZPHY C56 339 | H. Albrecht <i>et al.</i> | (ARGUS Collab.) |
| AMMAR | 92 | PR D45 3976 | R. Ammar <i>et al.</i> | (CLEO Collab.) |
| ARTUSO | 92 | PRL 69 3278 | M. Artuso <i>et al.</i> | (CLEO Collab.) |
| BAI | 92 | PRL 69 3021 | J.Z. Bai <i>et al.</i> | (BES Collab.) |
| BATTLE | 92 | PL B291 488 | M. Battle <i>et al.</i> | (CLEO Collab.) |
| BUSKULIC | 92J | PL B297 459 | D. Buskulic <i>et al.</i> | (ALEPH Collab.) |
| DECAMP | 92C | ZPHY C54 211 | D. Decamp <i>et al.</i> | (ALEPH Collab.) |
| ADEVA | 91F | PL B265 451 | B. Adeva <i>et al.</i> | (L3 Collab.) |
| ALBRECHT | 91D | PL B260 259 | H. Albrecht <i>et al.</i> | (ARGUS Collab.) |
| ALEXANDER | 91D | PL B266 201 | G. Alexander <i>et al.</i> | (OPAL Collab.) |
| ANTREASYAN | 91 | PL B259 216 | D. Antreasyan <i>et al.</i> | (Crystal Ball Collab.) |
| GRIFOLS | 91 | PL B255 611 | J.A. Grifols, A. Mendez | (BARC) |
| ABACHI | 90 | PR D41 1414 | S. Abachi <i>et al.</i> | (HRS Collab.) |
| ALBRECHT | 90E | PL B246 278 | H. Albrecht <i>et al.</i> | (ARGUS Collab.) |
| ALBRECHT | 90I | PL B250 164 | H. Albrecht <i>et al.</i> | (ARGUS Collab.) |
| BEHREND | 90 | ZPHY C46 537 | H.J. Behrend <i>et al.</i> | (CELLO Collab.) |
| BOWCOCK | 90 | PR D41 805 | T.J.V. Bowcock <i>et al.</i> | (CLEO Collab.) |
| DELAGUILA | 90 | PL B252 116 | F. del Aguila, M. Sher | (BARC, WILL) |
| GOLDBERG | 90 | PL B251 223 | M. Goldberg <i>et al.</i> | (CLEO Collab.) |
| WU | 90 | PR D41 2339 | D.Y. Wu <i>et al.</i> | (Mark II Collab.) |
| ABACHI | 89B | PR D40 902 | S. Abachi <i>et al.</i> | (HRS Collab.) |
| BEHREND | 89B | PL B222 163 | H.J. Behrend <i>et al.</i> | (CELLO Collab.) |
| JANSSEN | 89 | PL B228 273 | H. Janssen <i>et al.</i> | (Crystal Ball Collab.) |
| KLEINWORT | 89 | ZPHY C42 7 | C. Kleinwort <i>et al.</i> | (JADE Collab.) |
| ADEVA | 88 | PR D38 2665 | B. Adeva <i>et al.</i> | (Mark-J Collab.) |
| ALBRECHT | 88B | PL B202 149 | H. Albrecht <i>et al.</i> | (ARGUS Collab.) |
| ALBRECHT | 88L | ZPHY C41 1 | H. Albrecht <i>et al.</i> | (ARGUS Collab.) |
| ALBRECHT | 88M | ZPHY C41 405 | H. Albrecht <i>et al.</i> | (ARGUS Collab.) |
| AMIDEI | 88 | PR D37 1750 | D. Amidei <i>et al.</i> | (Mark II Collab.) |
| BEHREND | 88 | PL B200 226 | H.J. Behrend <i>et al.</i> | (CELLO Collab.) |
| BRAUNSCH... | 88C | ZPHY C39 331 | W. Braunschweig <i>et al.</i> | (TASSO Collab.) |
| KEH | 88 | PL B212 123 | S. Keh <i>et al.</i> | (Crystal Ball Collab.) |
| TSCHIRHART | 88 | PL B205 407 | R. Tschirhart <i>et al.</i> | (HRS Collab.) |
| ABACHI | 87B | PL B197 291 | S. Abachi <i>et al.</i> | (HRS Collab.) |
| ABACHI | 87C | PRL 59 2519 | S. Abachi <i>et al.</i> | (HRS Collab.) |
| ADLER | 87B | PRL 59 1527 | J. Adler <i>et al.</i> | (Mark III Collab.) |
| AIHARA | 87B | PR D35 1553 | H. Aihara <i>et al.</i> | (TPC Collab.) |
| AIHARA | 87C | PRL 59 751 | H. Aihara <i>et al.</i> | (TPC Collab.) |
| ALBRECHT | 87L | PL B185 223 | H. Albrecht <i>et al.</i> | (ARGUS Collab.) |
| ALBRECHT | 87P | PL B199 580 | H. Albrecht <i>et al.</i> | (ARGUS Collab.) |
| BAND | 87 | PL B198 297 | H.R. Band <i>et al.</i> | (MAC Collab.) |
| BAND | 87B | PRL 59 415 | H.R. Band <i>et al.</i> | (MAC Collab.) |
| BARINGER | 87 | PRL 59 1993 | P. Baringer <i>et al.</i> | (CLEO Collab.) |
| BEBEK | 87C | PR D36 690 | C. Bebek <i>et al.</i> | (CLEO Collab.) |
| BURCHAT | 87 | PR D35 27 | P.R. Burchat <i>et al.</i> | (Mark II Collab.) |
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